# Polarization requirements at the Z-pole

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•Short summary

- Impact of polarization at the Z-pole
- Physics issues at 150 vs. 250 position
- Open issues

# **Short summary**

- Extrapolated from LEP2, we need for calibration
  - 10 pb/detector + couple of pb's over the year
- For calibration: large emittance, low lumi tolerable
- But rather stable energy, not yet completely worked out
  - Only at GigaZ: energy stability, calibration accuracy < 0.1% (scope)</li>
  - Only at WW threshold: beam enrgy calibration ~ few 10<sup>-5</sup> (scope)
- We talk only about 'using calibration data for physics', not the high precision measurements
  - Resolving  $A_{LR}$  and  $A_{FB}$  discrepancy
  - Providing accurate  $sin^2\theta_{eff}$

# Physics gain with $\sin^2\theta_{eff}$ =3 x 10<sup>-5</sup>

Hints for new physics in worst case scenarios:



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Accuracy in sin<sup>2</sup>Ø<sub>eff</sub>

$$A_{\rm LR} = \frac{2(1 - 4\sin^2\theta_W^{\rm eff})}{1 + (1 - 4\sin^2\theta_W^{\rm eff})^2}$$

- → precision in ALR directly transferred to sin<sup>2</sup> 
  ⊕<sub>eff</sub>
- $\stackrel{\text{\tiny T}}{=}$  GigaZ will provide  $\Delta \sin^2 \Theta_{\text{eff}} \sim 1.3 \text{ x } 10^{-5}$  (if Blondel scheme)
- only electron polarization at GigaZ: ~9.5 x 10<sup>-5</sup>
- current value: 16 x 10<sup>-5</sup>
- What could we gain with a 'fraction' of GigaZ ?

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#### **Possible low lumi Z-data**

	ſL	No. of Z's	$\int_{\text{days}} \mathcal{L}_{\text{cal}}$	$P(e^{-})$	$P(e^+)$	$\Delta A_{\rm LR}^0$	$\Delta A_{\rm LR}$	$\sin^2 \theta_{\rm eff}$
	6 pb <sup>-1</sup>	$1.8 \times 10^5$	1	90%	0		$2.7 \times 10^{-3}$	$3.4 \times 10^{-4}$
				90%	40%	$3.3 \times 10^{-3}$	$4.4 \times 10^{-3}$	$5.6 \times 10^{-4}$
				90%	60%	$2.2 \times 10^{-3}$	$3.0 \times 10^{-3}$	$3.8 \times 10^{-4}$
	24 pb <sup>-1</sup>	$7.3  imes 10^5$	4	90%	0	—	$1.5 \times 10^{-3}$	$1.9 \times 10^{-4}$
				90%	40%	$1.6 \times 10^{-3}$	$2.2 \times 10^{-3}$	$2.8 \times 10^{-4}$
				90%	60%	$1.1 \times 10^{-3}$	$1.5 \times 10^{-3}$	$1.9 \times 10^{-4}$
	$60 \text{ pb}^{-1}$	$1.8 \times 10^{6}$	10	90%	0	—	$1.1 \times 10^{-3}$	$1.4 \times 10^{-4}$
				90%	40%	$1.0 \times 10^{-3}$	$1.4 \times 10^{-3}$	$1.8 \times 10^{-4}$
				90%	60%	$7.0 \times 10^{-4}$	$9.4 \times 10^{-4}$	$1.2 \times 10^{-4}$
	0.6 fb <sup>-1</sup>	$18 \times 10^{6}$	100	90%	0		$8.1 \times 10^{-4}$	$1.0 \times 10^{-4}$
				90%	40%	$3.3 \times 10^{-4}$	$4.4 \times 10^{-4}$	$5.6 \times 10^{-5}$
_				90%	60%	$2.2 \times 10^{-4}$	$3.0 \times 10^{-4}$	$3.8 \times 10^{-5}$
	0.9 fb <sup>-1</sup>	$27 \times 10^{6}$	150	90%	0	_	$7.9 \times 10^{-4}$	$1.0 \times 10^{-4}$
				90%	40%	$2.7 \times 10^{-4}$	$3.6 \times 10^{-4}$	$4.6 \times 10^{-5}$
				90%	60%	$1.8 \times 10^{-4}$	$2.4 \times 10^{-4}$	$3.1 \times 10^{-5}$
_	$1.2 \text{ fb}^{-1}$	$36 \times 10^{6}$	200	90%	0	—	$7.9 \times 10^{-4}$	$1.0 \times 10^{-4}$
				90%	40%	$2.3 \times 10^{-4}$	$3.1 \times 10^{-4}$	$4.0 \times 10^{-5}$
				90%	60%	$1.6 \times 10^{-4}$	$2.1 \times 10^{-4}$	$2.7 \times 10^{-5}$
	1.8 fb <sup>-1</sup>	$54 \times 10^{6}$	300	90%	0		$7.8 \times 10^{-4}$	$1.0 \times 10^{-4}$
				90%	40%	$1.9 \times 10^{-4}$	$2.6 \times 10^{-4}$	$3.2 \times 10^{-5}$
				90%	60%	$1.3  imes 10^{-4}$	$1.7 \times 10^{-4}$	$2.2 \times 10^{-5}$

Table 4: Lumi at Z-pole  $\mathcal{L}_{cal} = 7 \times 10^{31} \text{ cm}^{-2} \text{s}^{-1}, \ \sigma(e^+e^- \to Z \to had) \sim 30 \text{ nb}, A_{LR} = 0.154, \ \Delta P/P = 0.5\%, \ \mathcal{L}_{++,--}/\mathcal{L} = 0.1$ 

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# What's the role of polarization?

- Derive the statistical uncertainty of A<sub>LR</sub>
  - If only polarized electrons:

 $\Delta \, A_{LR}$  determined by polarimeter uncertainty

 $A_{LR} = 1 / P(e-) \times [\sigma_{L} - \sigma_{R}] / [\sigma_{L} + \sigma_{R}]$ 

- Pure error propagation: uncertainty depends on  $\Delta \sigma_L$ ,  $\Delta \sigma_R$ , ΔP/P
- For large statistics, σ (ee -> Z -> had) ~ 30 nb:
   main uncertainty from ΔP/P~ 0.5 % up to 0. 25%
- Since 'only' calibration and begin of ILC:  $\Delta P/P = 0.5$  %
- Higher P(e-) better, assumed 90%

#### **Blondel Scheme**

- Two polarized beams available
  - Express  $A_{LR}$  only by cross sections

$$\sigma = \sigma_{\text{unpol}} [1 - P_{e^-} P_{e^+} + A_{\text{LR}} (P_{e^+} - P_{e^-})],$$

$$A_{\text{LR}} = \sqrt{\frac{(\sigma_{++} + \sigma_{+-} - \sigma_{-+} - \sigma_{--})(-\sigma_{++} + \sigma_{+-} - \sigma_{-+} + \sigma_{--})}{(\sigma_{++} + \sigma_{+-} + \sigma_{-+} + \sigma_{--})(-\sigma_{++} + \sigma_{+-} + \sigma_{--} - \sigma_{--})}}}.$$

– Pure error propagation:

uncertainty depends on  $\Delta\sigma_{LL},~\Delta\sigma_{LR},\Delta\sigma_{RL},\Delta\sigma_{RR}$  not on  $\Delta P/P$ 

- Only relative measurements wrt flipping polarization needed
   ΔP / P = 0.5 % sufficient
- Some calibration time in LL and RR required assumed 10%, but that's not the optimum
- Different anal. powers:  $\Delta A_{LR} = \Delta A_{LR}^0 x \sqrt{(1+8/x)}$ , x~10=Ce/Ze

#### **Dependence of A**<sub>LR</sub> **on P(e<sup>+</sup>)**

0.002  $\Delta A_{\rm LR}$  On basis of 10<sup>6</sup> Z's 0.0018 0.0016 0.0014 •P(e<sup>+</sup>) important 0.0012 0.001 •Strictly speaking: 0.0008 0.0006 P(e<sup>+</sup>)=60% desirable 0.0004 0.0002 0

0.1 0.2 0.3 0.4 0.5 0.6 0.7 0.8 0.9 1  $P_{e^+}$ 

# **Dependence of A**<sub>LR</sub> on L<sub>++</sub> and L<sub>--</sub>

- What is the optimum time running in (++) and (--) mode?
- Assume P(e<sup>+</sup>)=40%
- Best value at about
- $(L_{++} L_{--})/L_{int} = 25\%$
- But does not significantly reduce the uncertainty!
- Higher P(e+) more effective



# Other topic: und@150 vs 250 GeV

(See also EUROTEV-Report-2005-015-1)

- Only some physics thoughts (see also weblog, July 08)
- 250 position: higher yield (about a factor ~3) but lumi problems for low  $\sqrt{s}$ 
  - For current parameters: drops below design value 1.5 from  $\sqrt{\rm s}{=}300~{\rm GeV}$  downwards
  - Possible lumi loss could be compensated by using bypass and half rate if lumi drops by factor 2
  - For current parameters this should happen between 200-240 GeV
- What's about expected physics in this energy range?

# **Physics at** $\sqrt{s}$ =200-240 GeV

- Light Higgs:
  - should be in range [115 210] GeV, that means σ(HZ) -> [220-300] GeV
  - Already at the upper edge of this region

  - Higgs mass in continuum up to 50 MeV
  - Threshold scans needed, e.g. for spin verification: 3 steps needed
  - Couplings measurements optimal at 50GeV+threshold: -> almost beyond critical region or at top threshold: anyway ok

# Which other physics is crucial?

- Top threshold: happens at 350 GeV.....ok
- Light SUSY: .....would be lovely, but even if...
  - studies will anyway be done first at 500 GeV
  - If threshold scans required, number of needed energy steps optimized via the continuum measurements (similar as for Higgs)
- Undulator position at 250 should be ok (even without by-pass)

# **Open Issues and Conclusions**

- What is the calibration lumi?
  - For deceleration scheme (only for 150 GeV applicable)
  - For 50 GeV scheme (for 150 GeV and 250 GeV applicable)
- What is the e<sup>+</sup> polarization at which lumi?
- Z-data required to solve A<sub>LR</sub> vs A<sub>FB</sub> and to enable powerful tests even in worst case scenarios

– We should not miss this opportunity!

- Undulator position 150 vs 250: should be ok for physics
  - Expected lower lumi for  $\sqrt{s}$ =200 240 GeV bearable