

CR2: Baseline optics to provide
for a single FFS L^* (QD0 exit – IP
distance) optics configuration

Glen White, SLAC

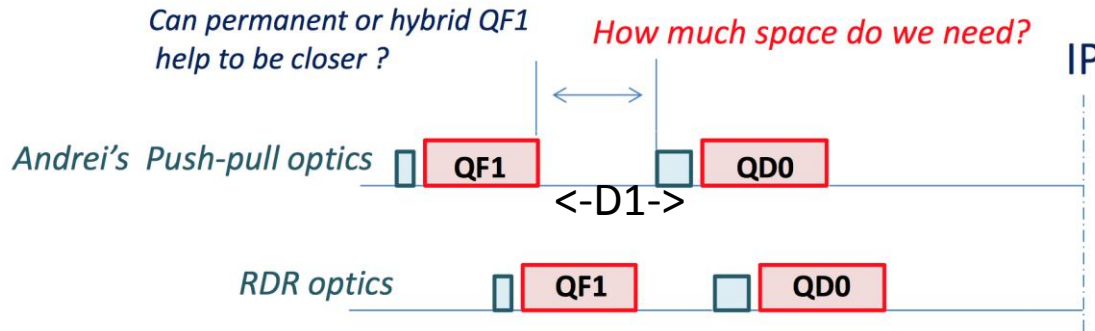
CMB Meeting

Sept. 25, 2014

General Considerations / Comments

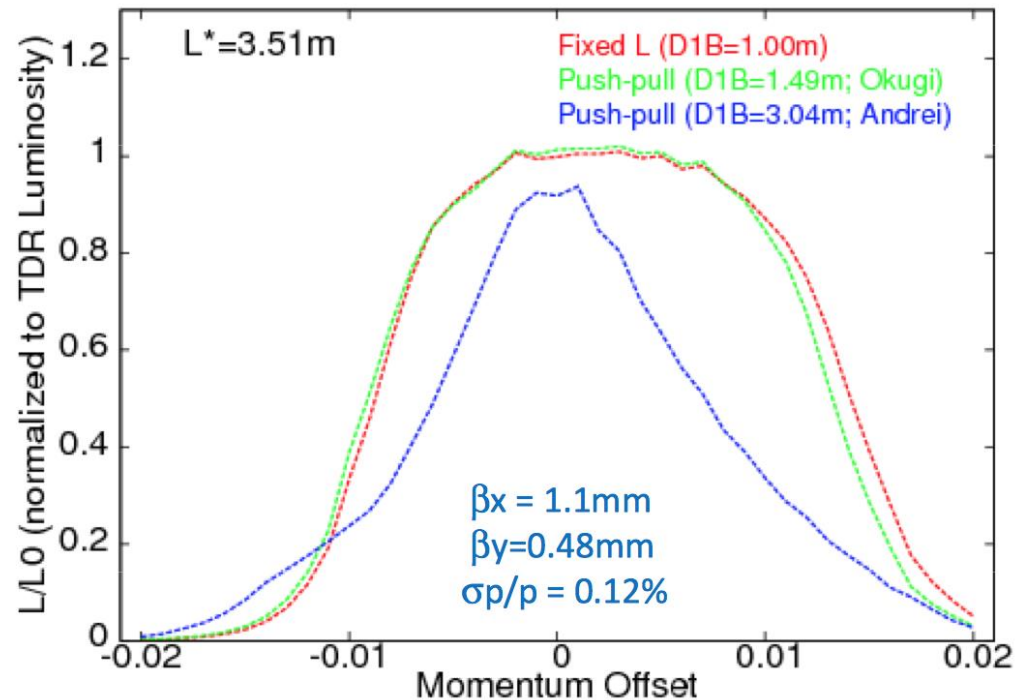
- Unequal L^* is not a *fundamental design or cost issue*
 - **We have feasible optics solutions**
- Primary issue is operational lumi performance and risk mitigation
 - harder to quantify, so arguments tend to be more fuzzy
 - **But based on considerable experimental and theoretical experience with this FFS design**
- L^* is a fundamental parameter that drives many critical design features of the BDS.
As L^* gets longer
 - Chromatic (and geometric) corrections become more challenging
 - Overall larger beta functions drive tolerances (field and alignment) become more demanding
 - Shielding IR from SR fan becomes harder
 - collimation depth becomes tighter for fixed IR apertures
 - tighter collimation tighter jitter tolerances from wakefields etc.
- Bottom line: for the accelerator, shorter is better, and
- Having different L^* will cause significant tuning differences between detectors
 - both lumi and background
 - negative impact on push-pull recovery times
 - **difficult to guarantee equal luminosity performance**

QD0-QF1 Distance Constraint



T.Okugi

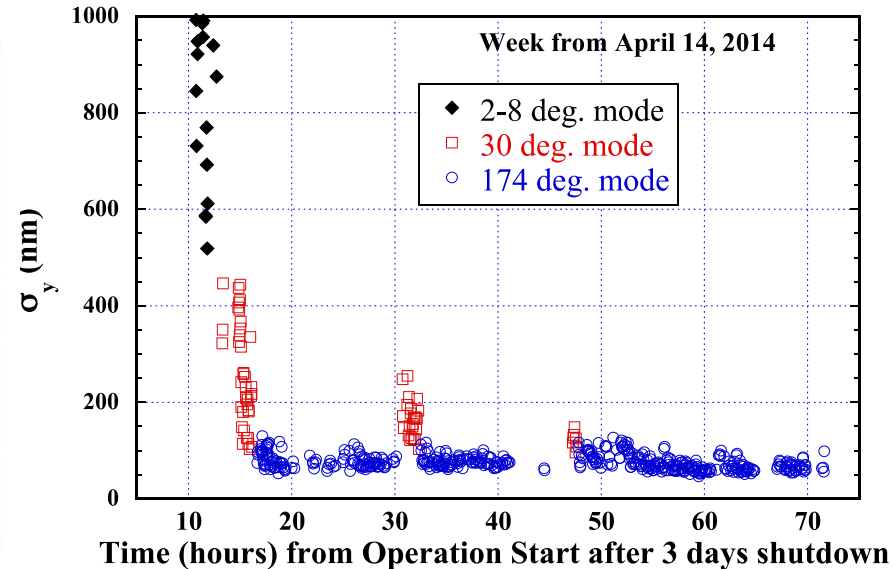
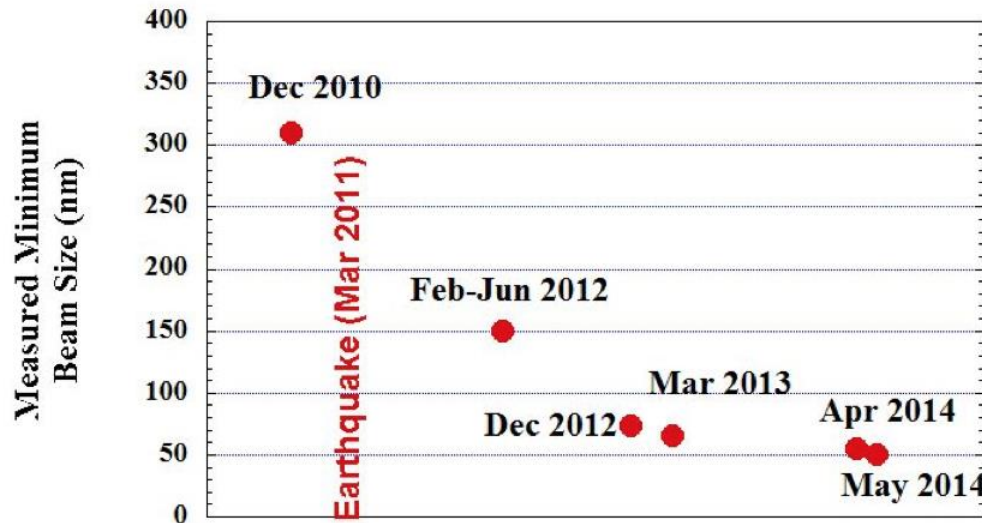
- Constraint of fixed QF1 position complicates “minimize L^* ” argument.
- Increased D1 distance degrades lattice performance
 - More detailed lattice studies required to determine optimum $L^* + D1$
 - May need flexibility to reduce D1
 - **Working assumption that a common $L^*=4\text{m}$ is good compromise**



Impact of Changing L* Optics

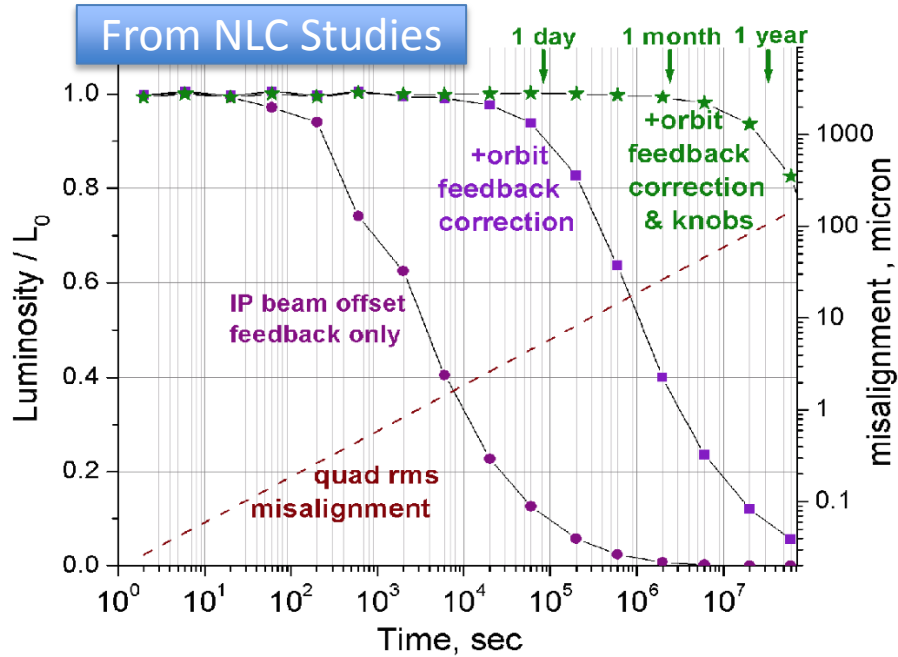
- Errors (misalignment and field errors) within the FFS lead to a specific set of first and higher-order aberrations at the IP specific to the lattice configuration.
 - These are corrected using standard BBA techniques and more complex FFS knobs using sextupoles on movers.
- Tuning can be considered on 2 timescales:
 - (1) the few-hour (and then periodic to counter ground motion) application of multiknobs etc to maximise luminosity after some period of beam-off condition
 - (2) longer timescale period where the result of tuning iteratively gets better (ATF2 & FFTB experience has shown this behaviour).
- Change between L* configurations on periodic basis can be understood to be harmful in terms of losing the iterative work earned from (2).
- Impact of 1 vs. 2 L*'s driven by:
 - Function of (2)
 - Push-pull duration
 - Push-pull frequency
 - Luminosity tuning curve associated with (1)

ATF2 Experience



- Effects of “long-term” tuning efforts shown as beam size improvement over time (left plot)
- Time to re-tune after multi-day shutdown shown on right (~20 hours).

Simulations



Required β COLL Apertures to prevent IR SR hits

| Name | $L^*=3.5\text{m}$ | | $L^*=4.5\text{m}$ | |
|------|------------------------|------------------------|------------------------|------------------------|
| | X / mm ($N\sigma_x$) | Y / mm ($N\sigma_y$) | X / mm ($N\sigma_x$) | Y / mm ($N\sigma_x$) |
| SP1 | - | - | - | - |
| SP2 | - | - | 1.03 (9.3) | 0.21 (25) |
| SP3 | - | - | 0.4 (21) | 0.21 (203) |
| SP4 | - | - | 0.48 (4.3) | 0.2 (24) |
| SP5 | - | - | - | - |

- After O(1 day), orbit control no longer sufficient to regain luminosity, need complex tuning algorithms (using sextupole movers etc).
 - This sextupole-based tuning is very sensitive to the exact lattice configuration and will take longer for a different FFS optics (after a ~ 1 day push-pull operation) than for a single solution.
- Very different collimation requirements for different L^* 's
 - e.g. for IR SR collimation:
 - $L^*=3.5\text{m}$: magnet apertures shield all SR generating particles from causing hits in IR
 - $L^*=4.5\text{m}$: tight collimation required ($\sim 4\sigma$)

Summary

- For least risk & optimal luminosity from accelerator tuning, prefer single L^* optics solution.
- From detector groups:
 - SiD can allow $2.13 < L^* < 4.5$ (m)
 - ILD has more constraints, 4m may be possible as a minimum
- Working assumption and next steps:
 - Common $L^* = 4$ m
 - Understand performance trade-offs for larger vs. smaller D1
 - Develop $L^*=4$ m optics (and associated BDS design items) and present performance (vs. 3.5 & 4.5) at LCWS
 - Detector groups need to design for 4m L^* and (if strongly suggested by design studies) consider flexibility in D1 possible.