

A General Framework for SUSY Breaking



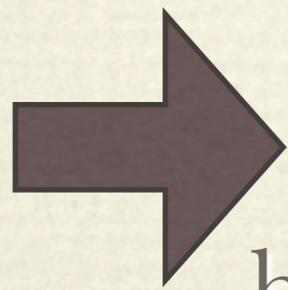
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Benchmark models

- ❖ Particular SUSY breaking models are used in designing triggers, constraining SUSY parameter space
- ❖ SPS points
- ❖ 6 mSUGRA points, 2 GMSB points, 1 AMSB point
- ❖ Theoretical/Experimental prejudices about spectra result

- ❖ These prejudices can have important implications for SUSY searches
- ❖ Jets + mET searches for gluinos
- ❖ in mSUGRA $m_{\tilde{g}} : m_{\tilde{B}} = 6 : 1$
- ❖ This ratio never scanned in mSUGRA motivated searches

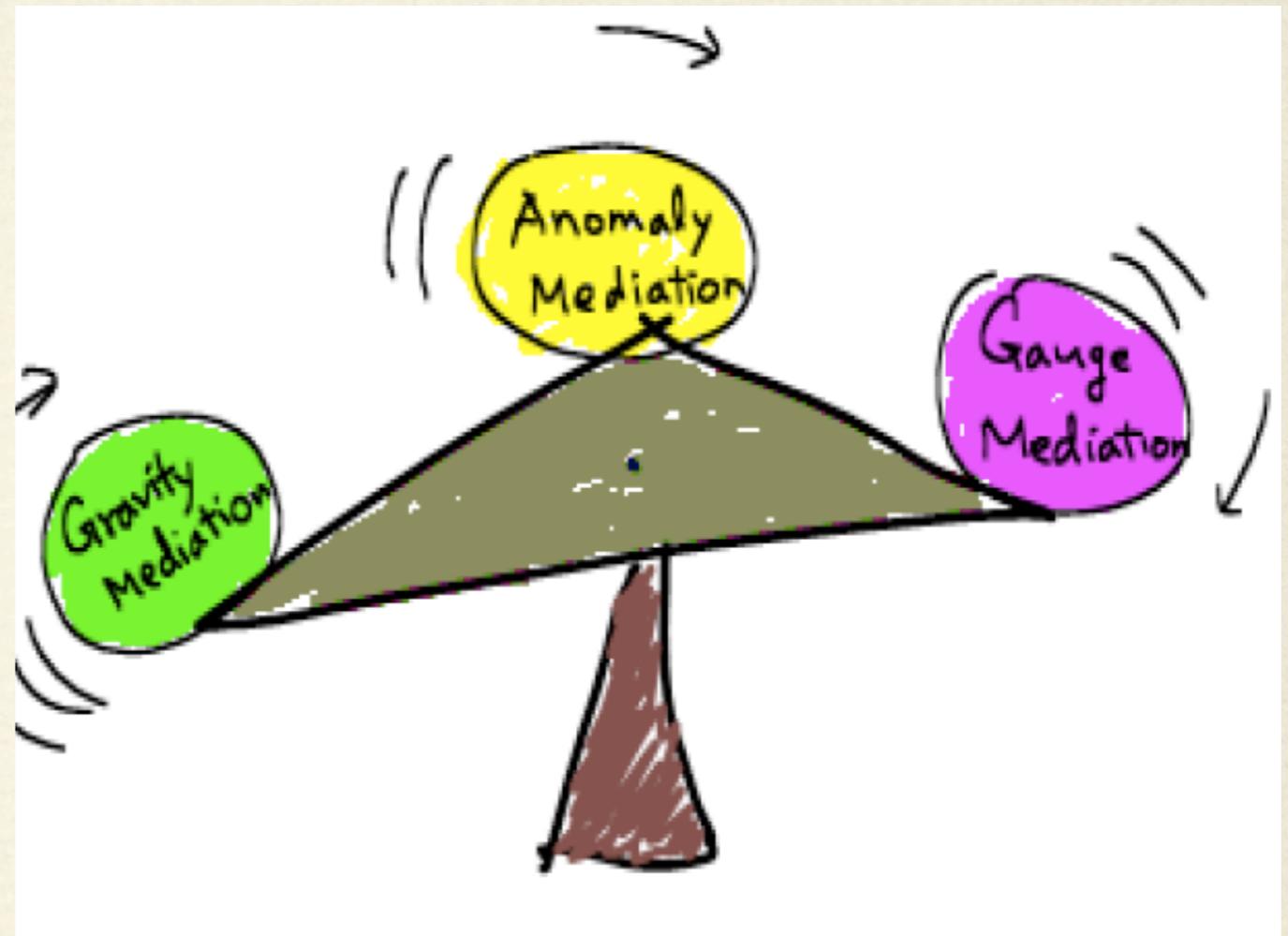


Kinematically accessible regions which have never been scanned over in searches

- ❖ It's been shown simple 1-parameter extensions blow these theoretical prejudices away
 - ❖ Within mSUGRA, small μ FP region occurs for $m_0 \gg m_{1/2}$
 - ❖ scalars are decoupled at LHC
 - ❖ Within mSUGRA wino content of LSP is never large, and we never get Bino-Wino coannihilation
- $M_1(\text{weak}) \simeq M_2(\text{weak}) \implies$ Mixed wino DM (MWDM);
 $M_1(\text{weak}) \simeq -M_2(\text{weak}) \implies$ bino-wino co-annihilation (BWCA);
Low $|M_3|$ or large $M_2 \implies$ Low $|\mu|$, so mixed higgsino DM (MHDM).
- ❖ By adjusting one parameter, all points in $m_0 - m_{1/2}$ plane become relic density allowed!

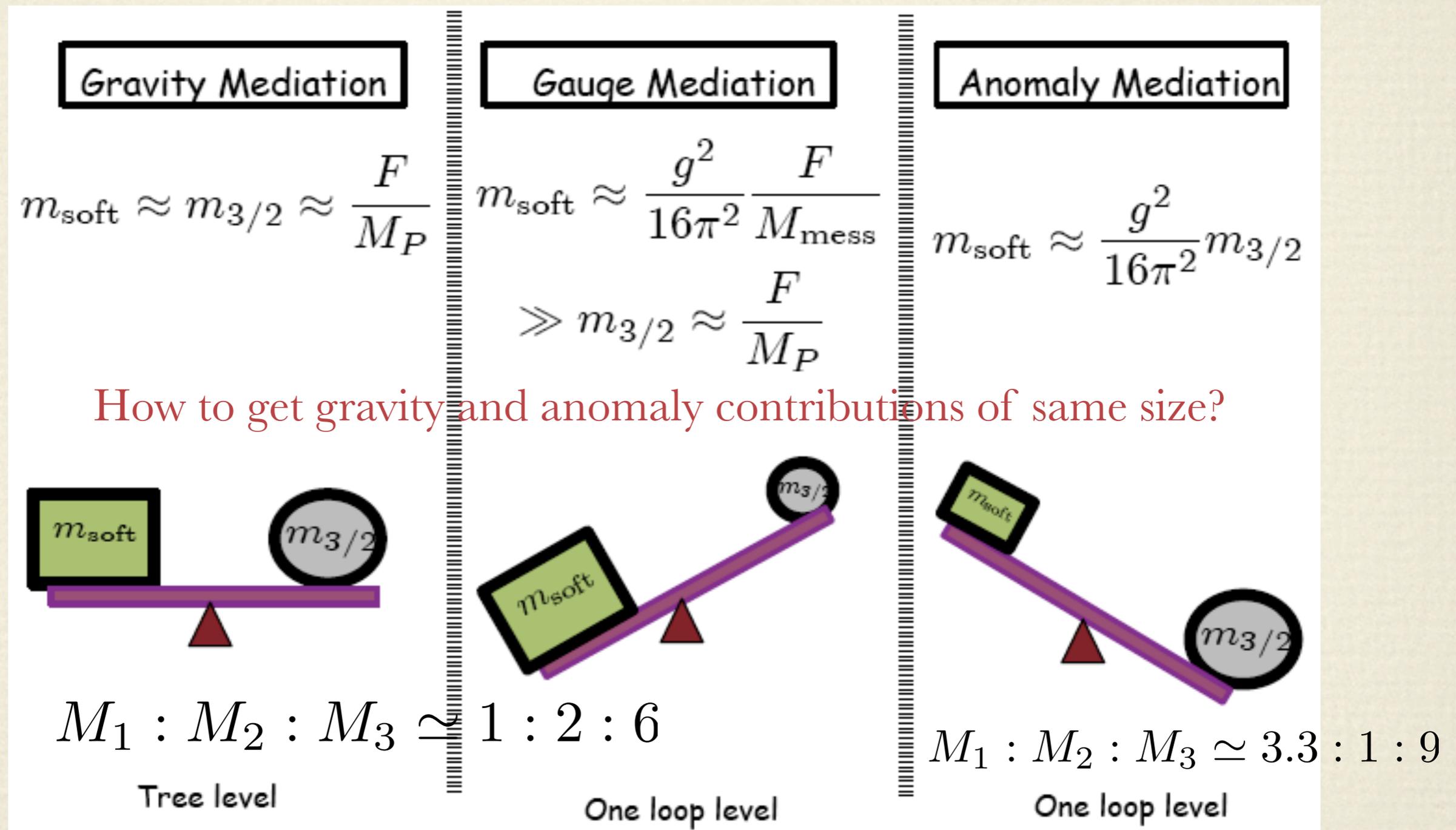
A more general framework?

- ❖ Is there a framework?
- ❖ Is it well-motivated?
- ❖ Does it produce phenomenologically novel MSSM spectra?



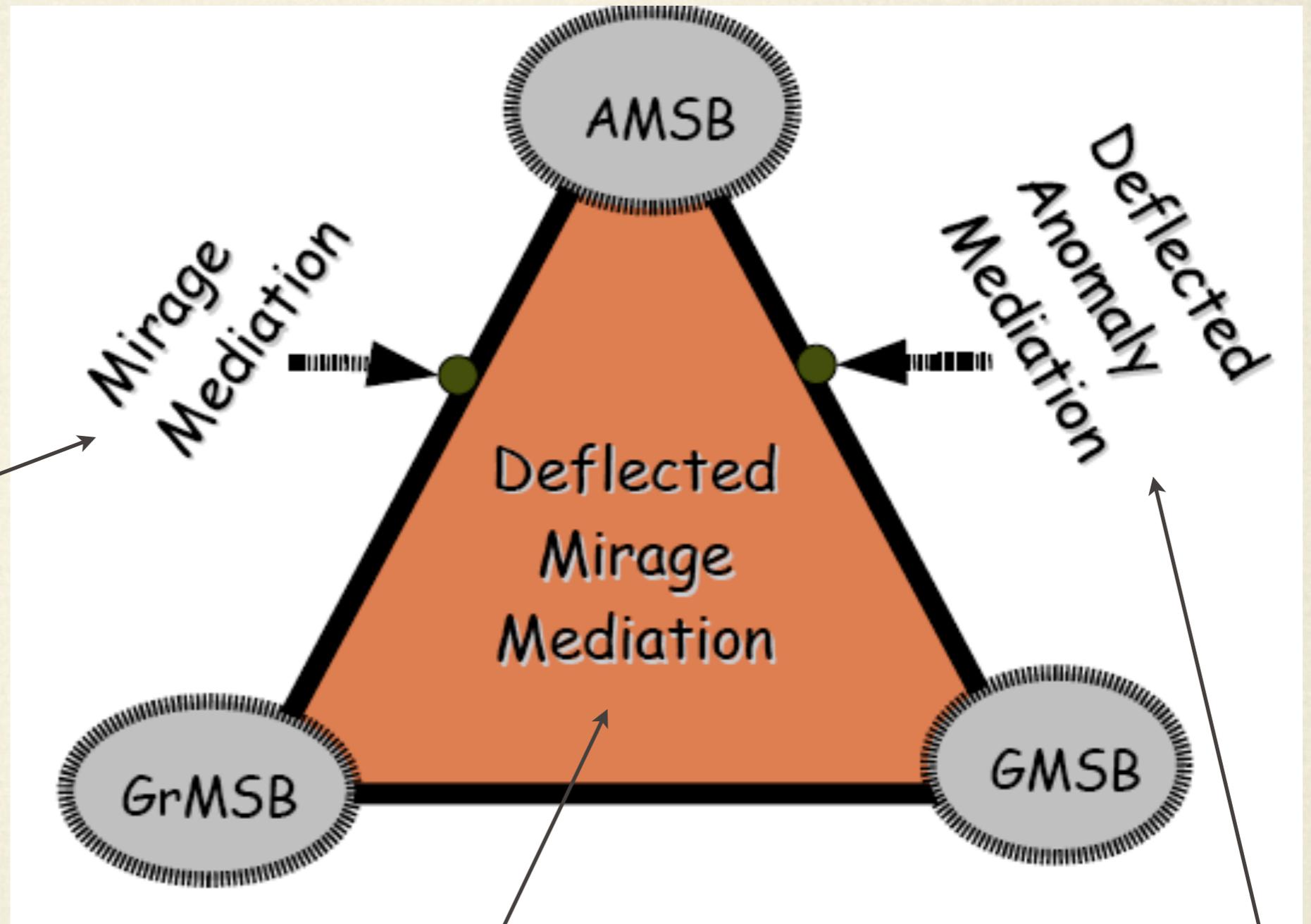
- ❖ A framework for dialing between any of these types of SUSY breaking schemes

Different mediation mechanisms have different spectra



Motivation for contribution from all types

Use moduli stabilization tricks to make on the same size (KKLT) K. Choi et al.



This work

Rattazzi and Wells

A model where this happens

All three contributions active

$$W = w_0 + Ae^{-aT} + \lambda X \psi \tilde{\psi}$$

KKLT Superpotential
GrMSB and AMSB same size

New piece

Take GMSB SUSY breaking field X to be matter modulus

Stabilize X by anomaly mediated SUSY breaking terms

$$\begin{aligned} F^X &= -e^{K/2} K^{X\bar{X}} D_{\bar{X}} \bar{W} \\ &= \underbrace{-e^{K/2} K^{X\bar{X}} \partial_{\bar{X}} \bar{W}}_{(A)} \underbrace{-e^{K/2} K^{X\bar{X}} K_{\bar{X}} \bar{W}}_{(B)} \end{aligned}$$

$$\frac{F^X}{X} = -m_{3/2} + \mathcal{O}\left(\frac{m_{3/2}}{8\pi^2}, \frac{F^T}{T + \bar{T}}\right)$$

Phenomenologically interesting

Just parametrize and dial it....

❖ Three pieces

❖ GrMSB (modular weights)

❖ AMSB

❖ GMSB M_{mess}

$$\begin{aligned}\frac{F^T}{T + \bar{T}} &= m_0 \\ \frac{F^C}{C} &= \alpha_m \ln(m_P / m_{3/2}) m_0 \\ \frac{F^X}{X} &= \alpha_g \frac{F^C}{C}\end{aligned}$$

General prescription to include effects of all types of SUSY breaking

And dial the soft masses...

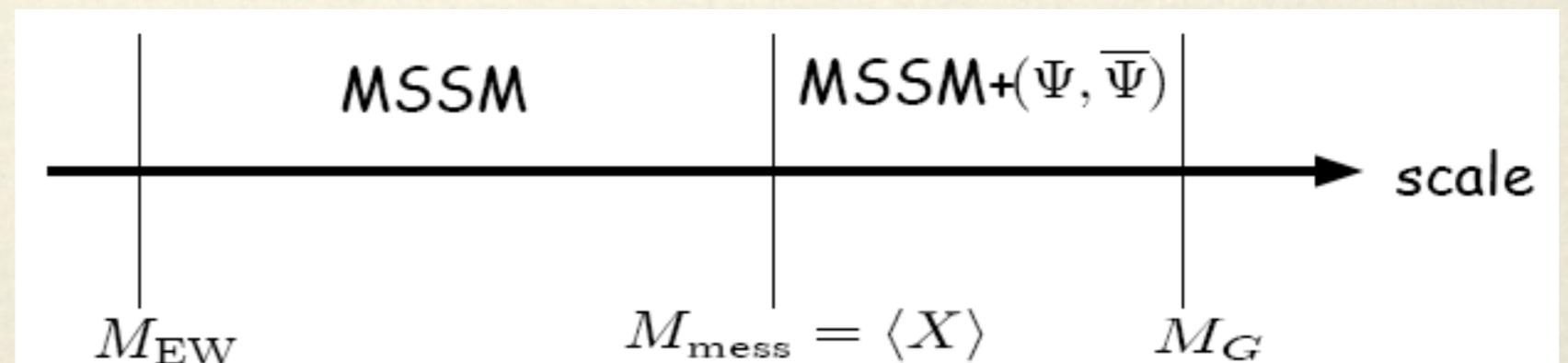
$$M_a = F^n \partial_n \log \text{Re}(\alpha_a^{-1}(\mu))$$

$$A_i = F^n \partial_n \log e^{-K_0/3} Z_i$$

$$m_i^2 = -F^m F^{\bar{n}} \partial_m \partial_{\bar{n}} \log e^{-K_0/3} Z_i$$

$$Z_i(\mu) = Z_i(\Lambda_{UV}) \prod_a \left(\frac{\alpha_a(\Lambda_{UV})}{\alpha_a(X)} \right)^{\frac{2c_a}{b_a - N}} \left(\frac{\alpha_a(X)}{\alpha_a(\mu)} \right)^{\frac{2c_a}{b_a}}$$

- ❖ Compute UV soft masses
- ❖ Run to messenger scale; add in GMSB contribution
- ❖ Run to IR



Example: gaugino masses

$$M_a(\mu = M_{\text{GUT}}) = \frac{F^T}{T + \bar{T}} + \frac{\alpha_{\text{GUT}}}{4\pi} b'_a \frac{F^C}{C}$$

← GUT scale

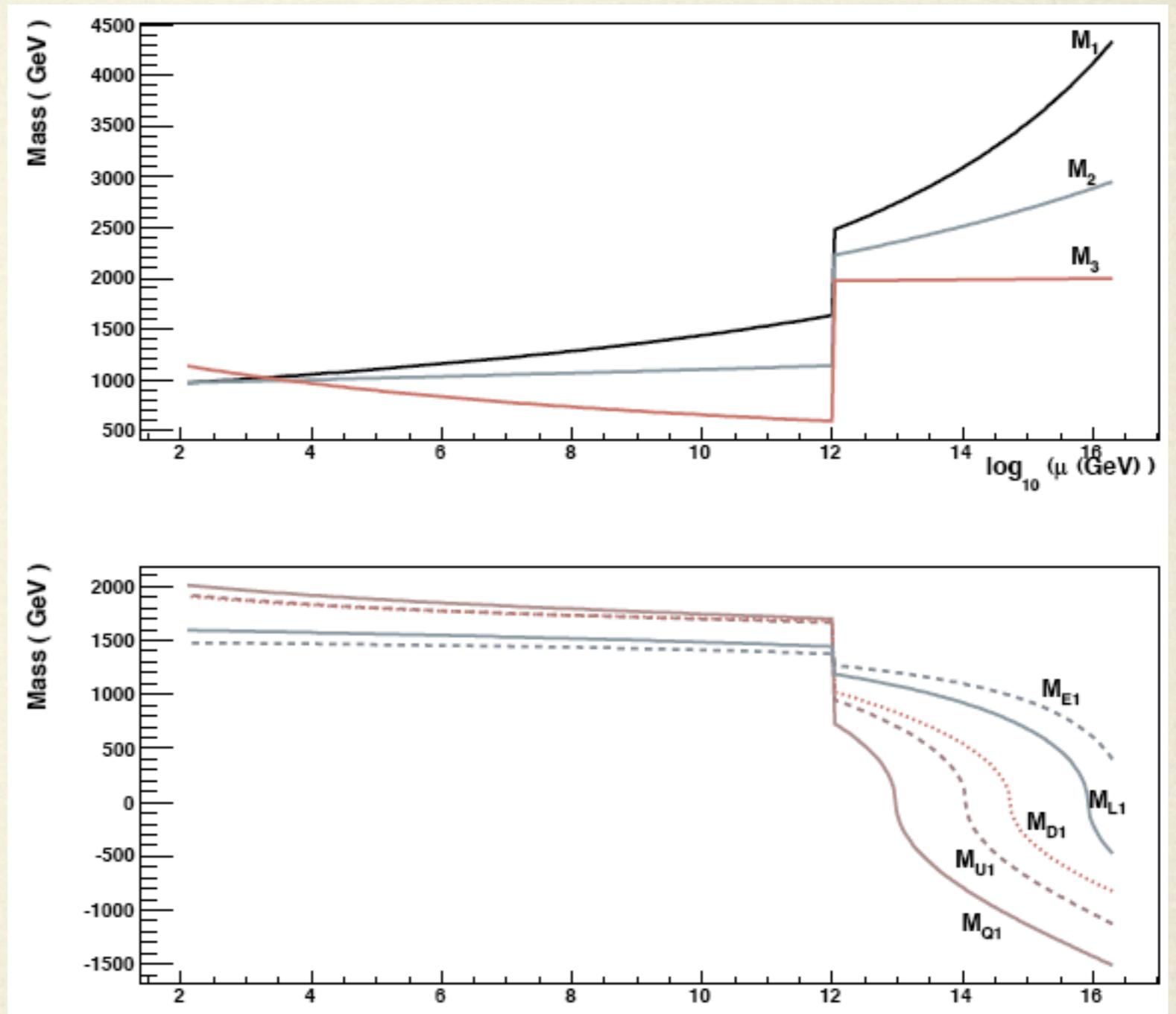
$$M_a(\mu = M_{\text{mess}} - \epsilon) = M_a(\mu = M_{\text{mess}} + \epsilon) + \Delta M_a,$$

← Messenger scale addition

$$\Delta M_a = -\frac{\alpha_a(M_{\text{mess}})}{4\pi} N_{\text{mess}} \left(\frac{F^C}{C} + \frac{F^X}{M_{\text{mess}}} \right)$$

Non-standard points

- ❖ Large negative contribution to gaugino masses
- ❖ Quasi-conformal running of scalar masses
- ❖ Light gluino

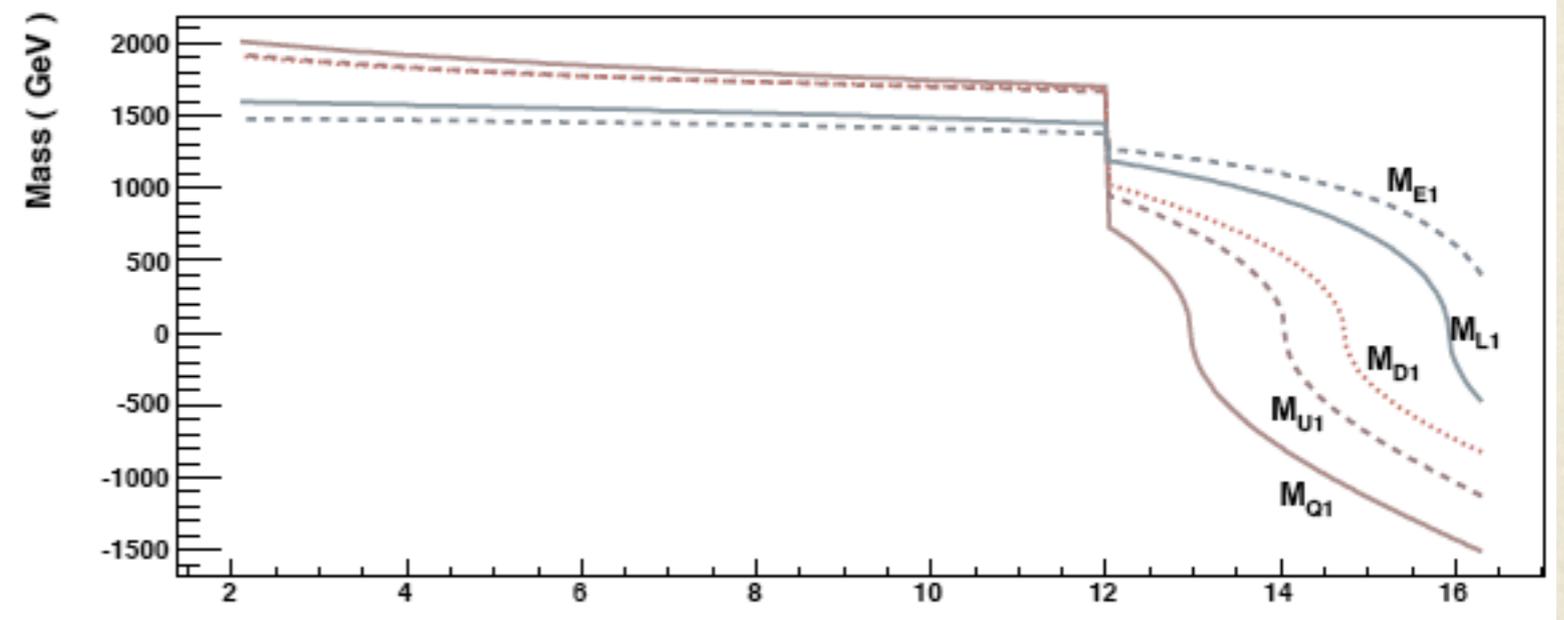
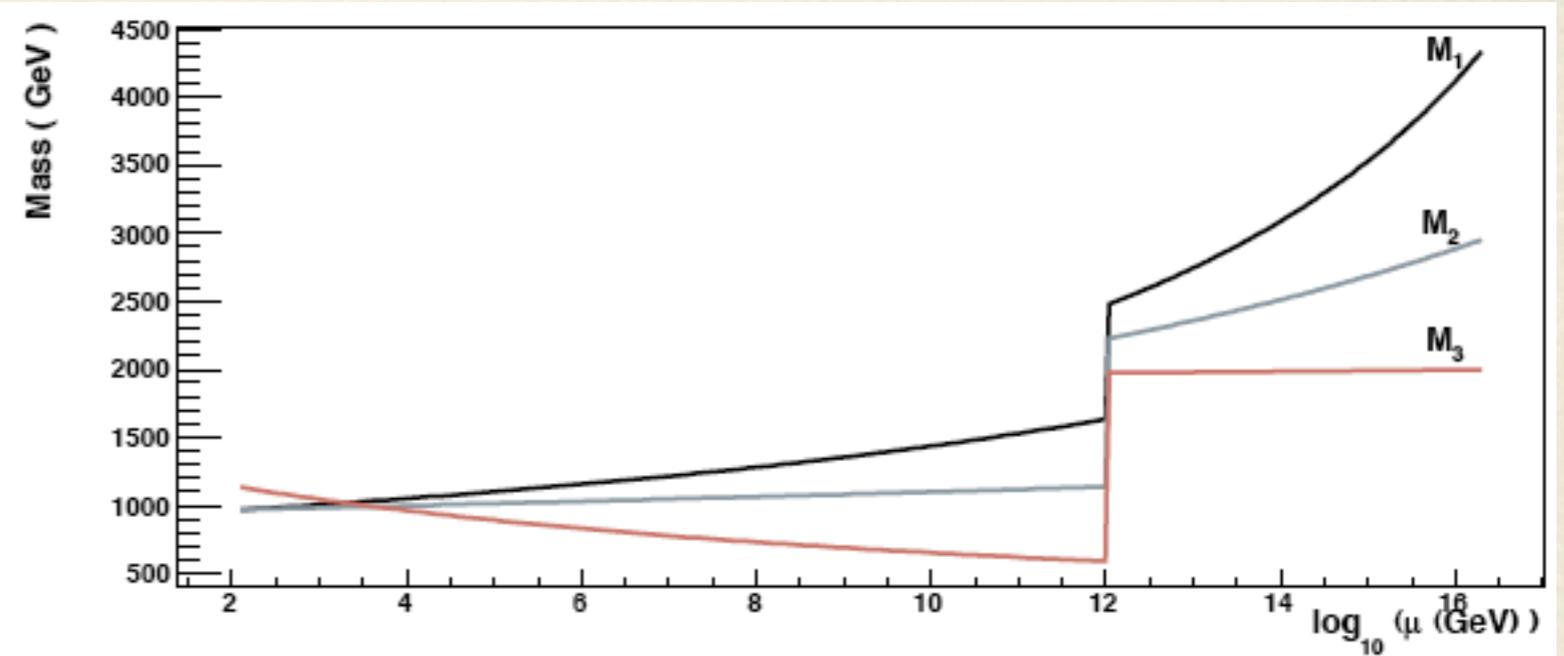


$$\alpha_m = \alpha_g = 1 \quad M_{mess} = 10^{12} \text{ GeV}$$

Non-standard points

❖ Light stop

Point A			
h	117	H, A	1529
\tilde{g}	1170	H^\pm	1531
χ_1^0	1003	χ_2^0	1015
χ_3^0	1374	χ_4^0	1380
χ_1^\pm	1011	χ_2^\pm	1369
\tilde{u}_L	1965	\tilde{u}_R	1890
\tilde{d}_L	1974	\tilde{d}_R	1888
\tilde{e}_L	1587	\tilde{e}_R	1470
$\tilde{\mu}_L$	1587	$\tilde{\mu}_R$	1470
\tilde{t}_1	1420	\tilde{t}_2	1791
\tilde{b}_1	1769	\tilde{b}_2	1872
$\tilde{\tau}_1$	1459	$\tilde{\tau}_2$	1583

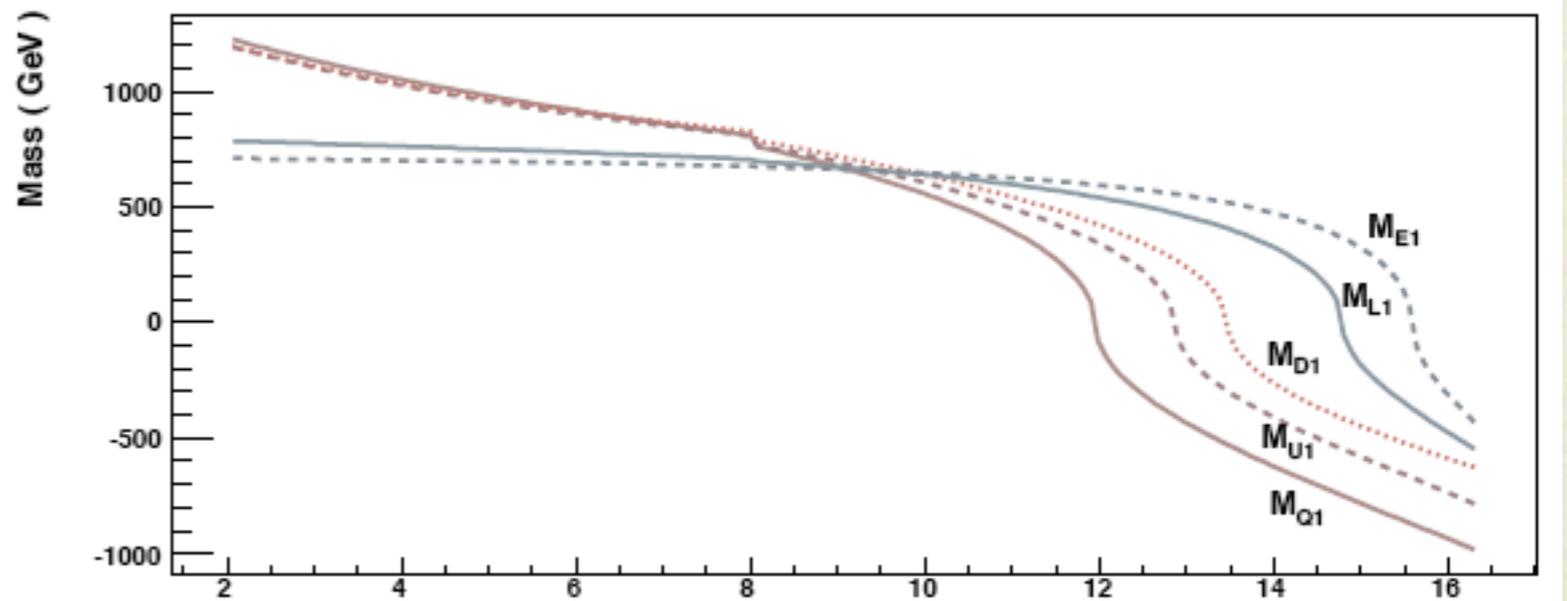
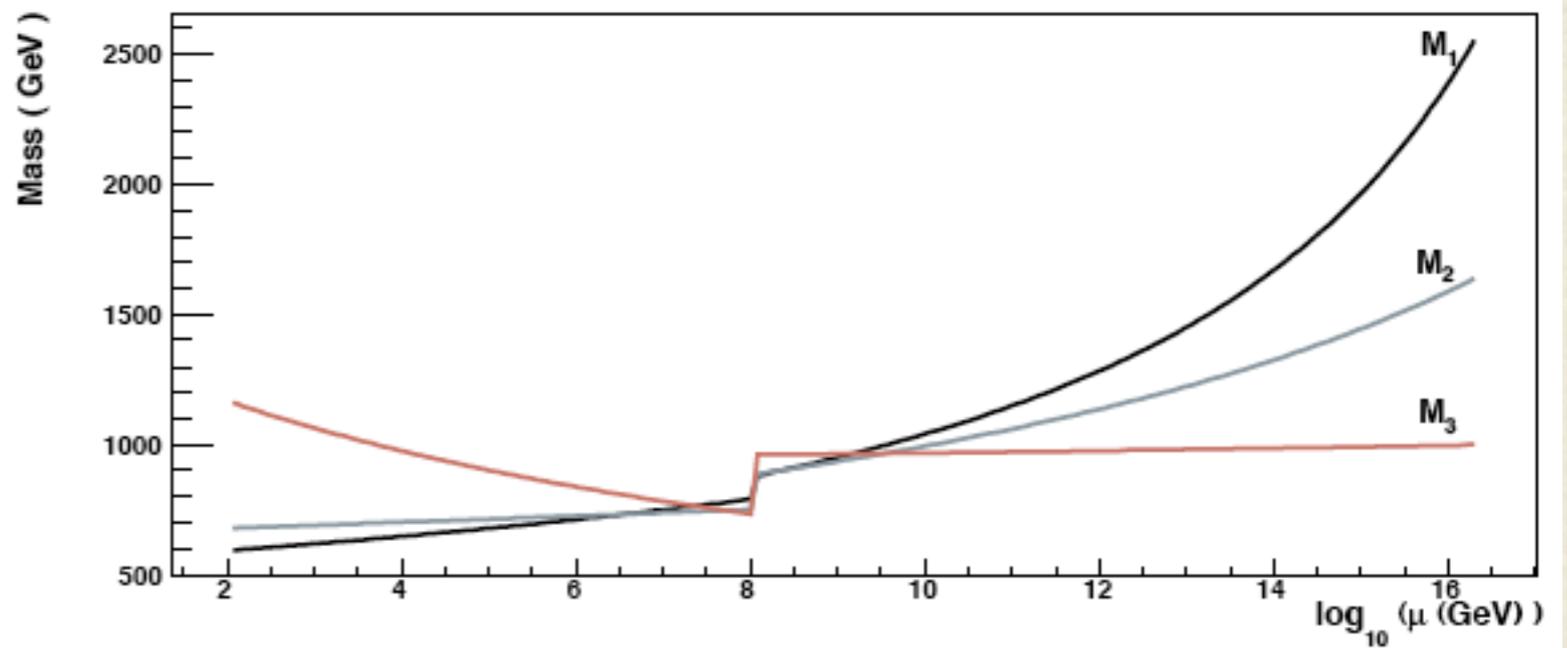


$$\alpha_m = \alpha_g = 1 \quad M_{mess} = 10^{12} \text{ GeV}$$

A point of interest

❖ Compressed spectrum

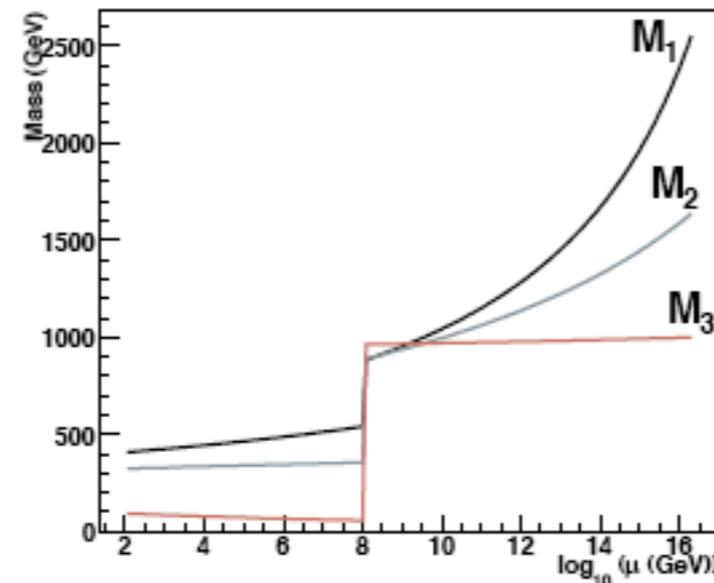
Point B			
h	116	H, A	865
\tilde{g}	1130	H^\pm	869
χ_1^0	608	χ_2^0	683
χ_3^0	818	χ_4^0	844
χ_1^\pm	682	χ_2^\pm	835
\tilde{u}_L	1164	\tilde{u}_R	1140
\tilde{d}_L	1172	\tilde{d}_R	1148
\tilde{e}_L	783	\tilde{e}_R	709
$\tilde{\mu}_L$	783	$\tilde{\mu}_R$	709
\tilde{t}_1	860	\tilde{t}_2	1113
\tilde{b}_1	1059	\tilde{b}_2	1141
$\tilde{\tau}_1$	702	$\tilde{\tau}_2$	782



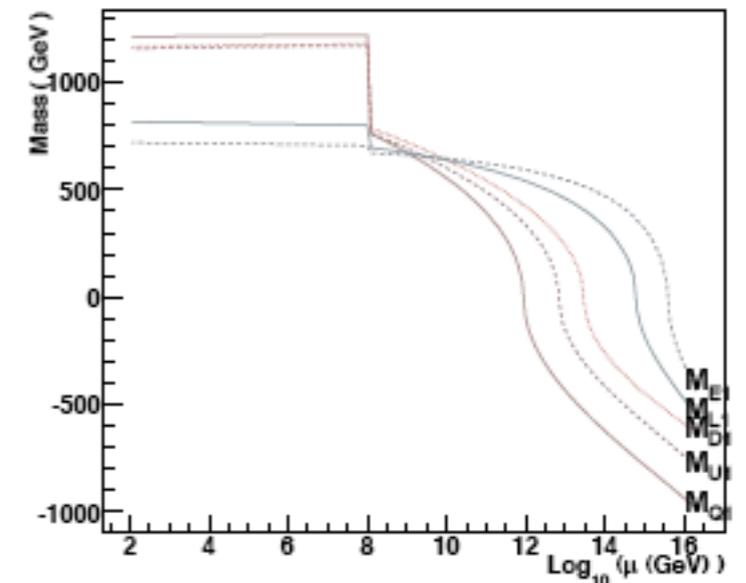
$$\alpha_m = 1 \quad \alpha_g = 1/2 \quad M_{mess} = 10^8 \text{ GeV}$$

Points of interest

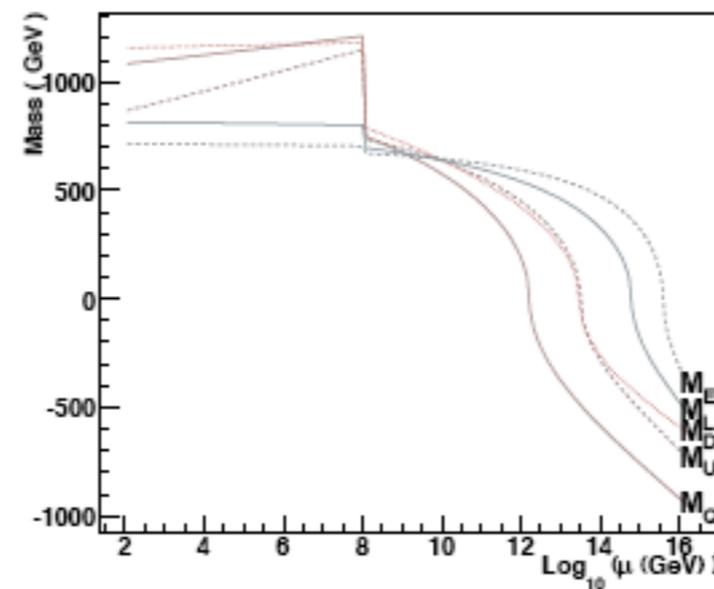
❖ Very light gluino



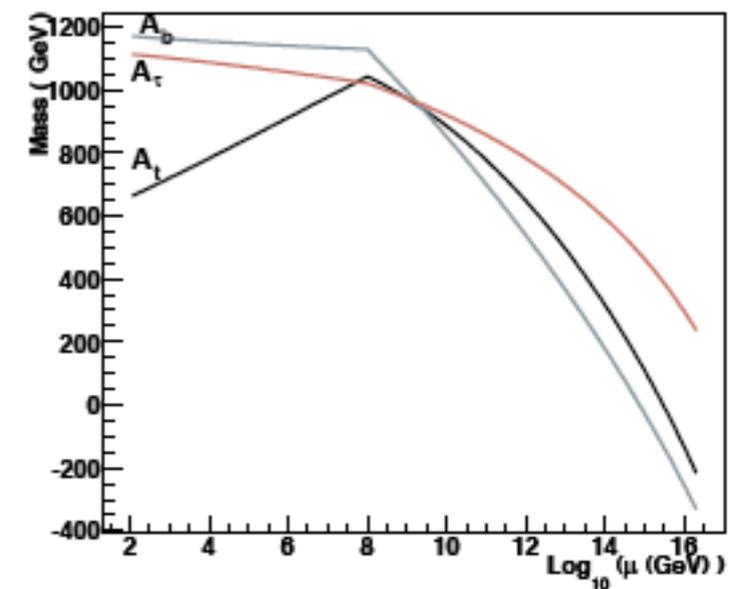
(a) Gaugino masses.



(b) First family soft scalar masses.



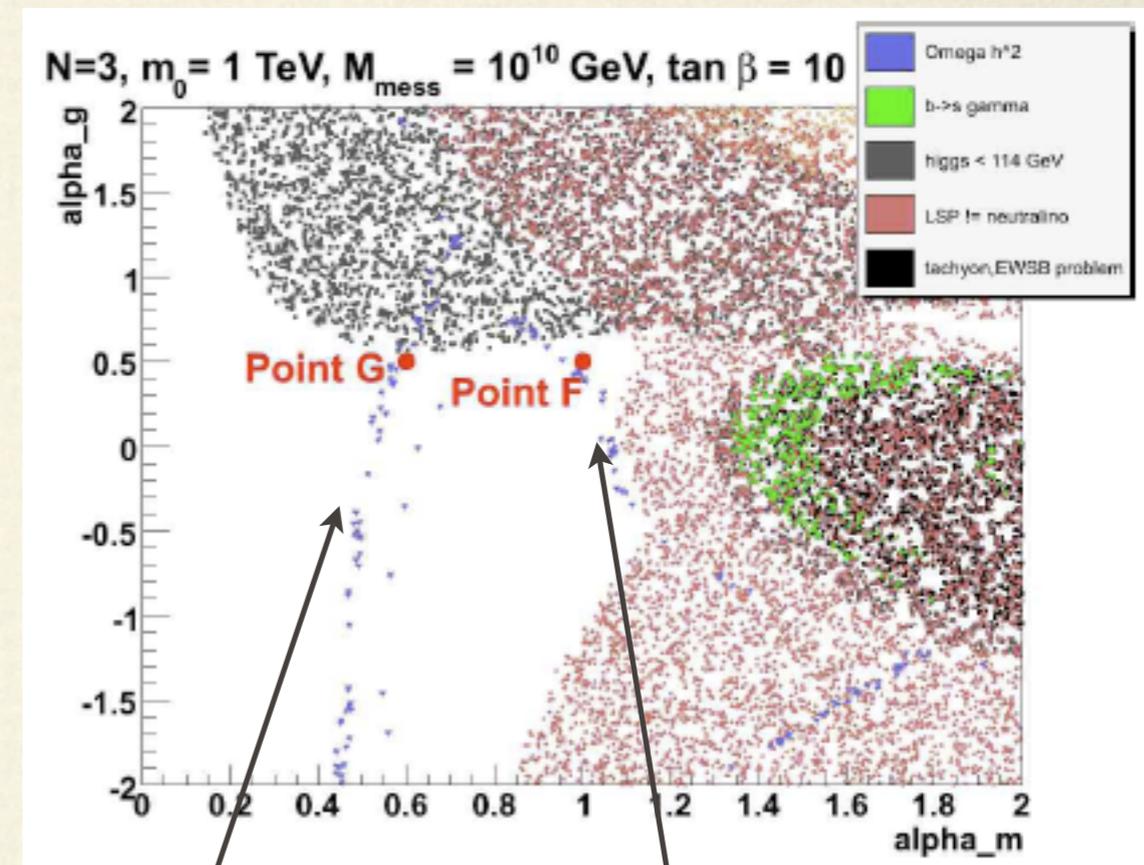
(c) Third family soft scalar masses.



(d) Third family soft trilinear terms.

A word on DM phenomenology

- ❖ Low scale unification lends itself well to “well-tempered neutralino”
- ❖ May have stop co-annihilation
- ❖ May naturally have Bino-Wino coannihilation



Funnel

BWCA

Conclusions

- ❖ A look past mSUGRA
- ❖ Simple framework for generalized SUSY breaking
 $m_0, \alpha_g, \alpha_m, \tan \beta, \text{sgn}(\mu)$
(+ modular weights)
- ❖ Framework can provide for some of the features that have been explored in bottom-up context, such as light gluinos and stops
- ❖ Some mSUGRA folklore shown to be wrong
- ❖ A framework for exploring SUSY more broadly?