

ILC2008

RTML and stray magnetic fields.

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Previous work 1

- Sensitivity to Nano-Tesla Scale Stray Magnetic Fields, by **J. Frisch, T. O. Raubenheimer, P. Tenenbaum**, SLAC, LCC-Note-0140 (June 7, 2004)
 - Analysis for NLC
 - Data from SLC (End station B)
 - Conclusion: *we are mostly OK*
 - **Questions:**
 - **How representative is the data?**

Previous work 2

- Rough estimation of effects of fast changing stray field in long transport of RTML - Emittance dilution in Turnaround, **K. Kubo**, KEK, ILC-Asia-2006-05 October 12, 2006
 - Requirement estimation: rms B < 2 nano-Tesla (ILC RDR), **relaxed to 10 nanotesla (1e-4 Gauss).**

Magnetic fields

- Commercial superconducting solenoid – **10 Tesla (1 e+1)**
- Earth magnetic field -- **50 micro-Tesla (5 e-5)**
- Cell phone – **100 nano-Tesla (1 e-7)**
- **ILC-RDR requirement – 10 nano-Tesla (1 e-8)**
- Beating human heart -- **~ 10 pico-Tesla (1 e-11)**

Classification

Following F.R.T. paper:

- 60 Hz and its harmonics (near-coherent with 5-Hz pulsing)
- Fields from RF systems (coherent with 5-Hz pulsing)
- Others (non-RF technical sources) (uncorrelated with pulses)

Assumptions:

Frequency < 0.1 are compensated by feedback. Higher frequency in uncorrelated noise $\sim 1/f^2$ (to keep energy density approximately constant). Very high frequencies (>100 kHz) attenuate in the structure.

We will look at $\sim 0.1 \text{ Hz} < f < \sim 10 \text{ kHz}$, in particular at 60 Hz problem.

Fermilab data

- Take data at **A0** experimental hall with 8 MW working klystron.
- Use 3-axis fluxgate magnetometer (Bartington Mag-03MC)
 - +/- 1mT full scale, DC to 3 kHz, 20 pT/sqrt(Hz)



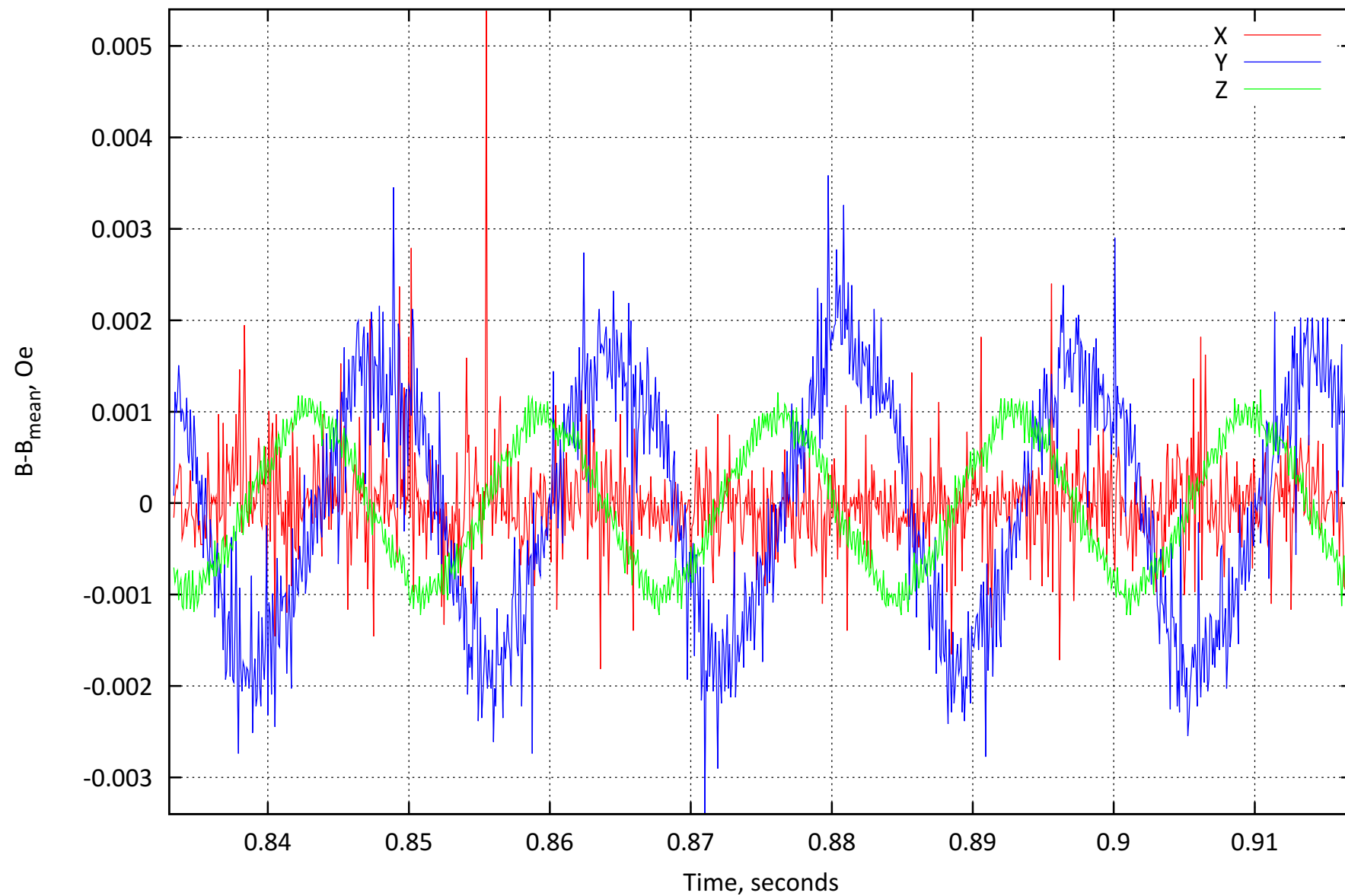


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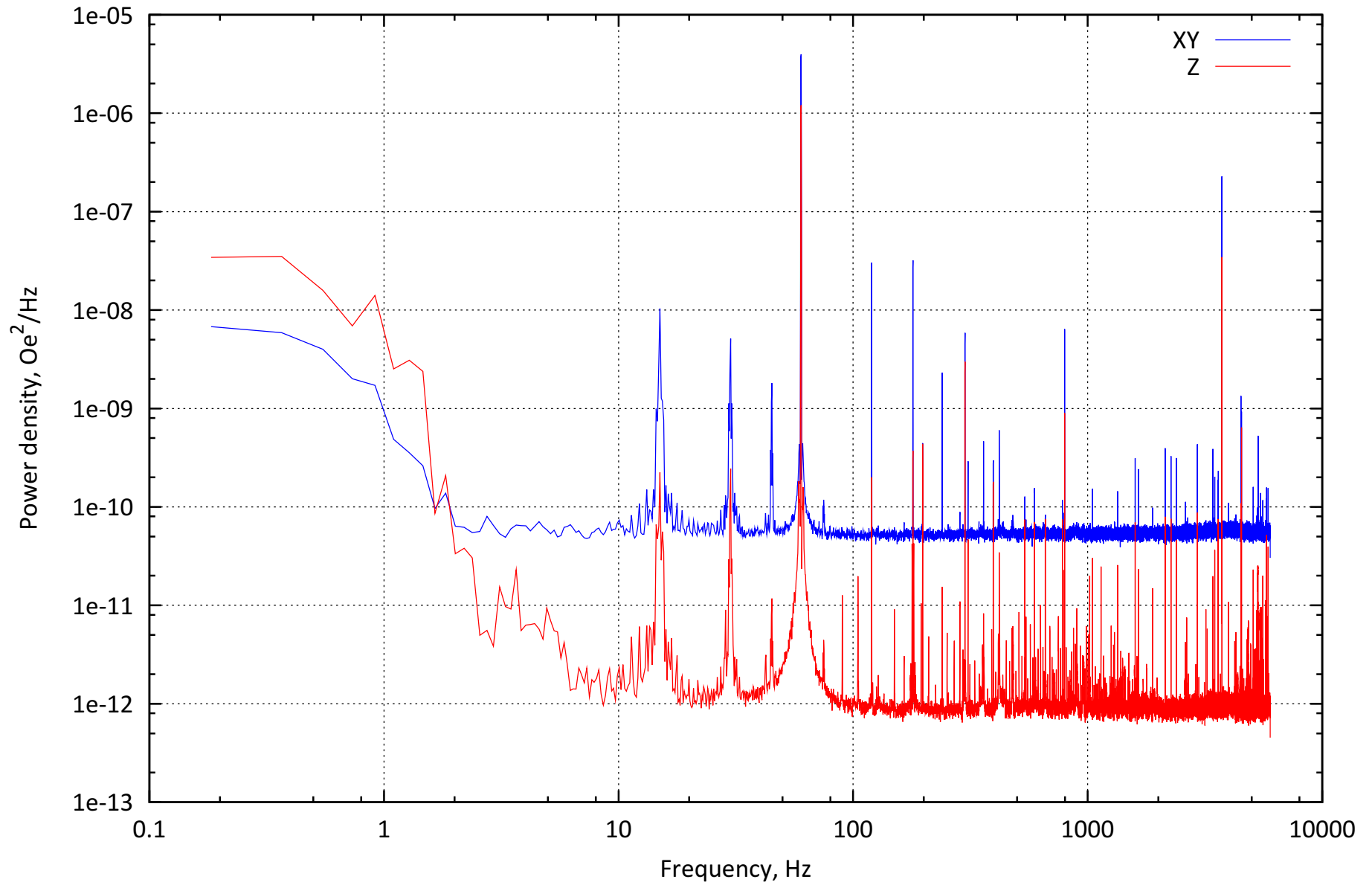
Data taking at Fermilab



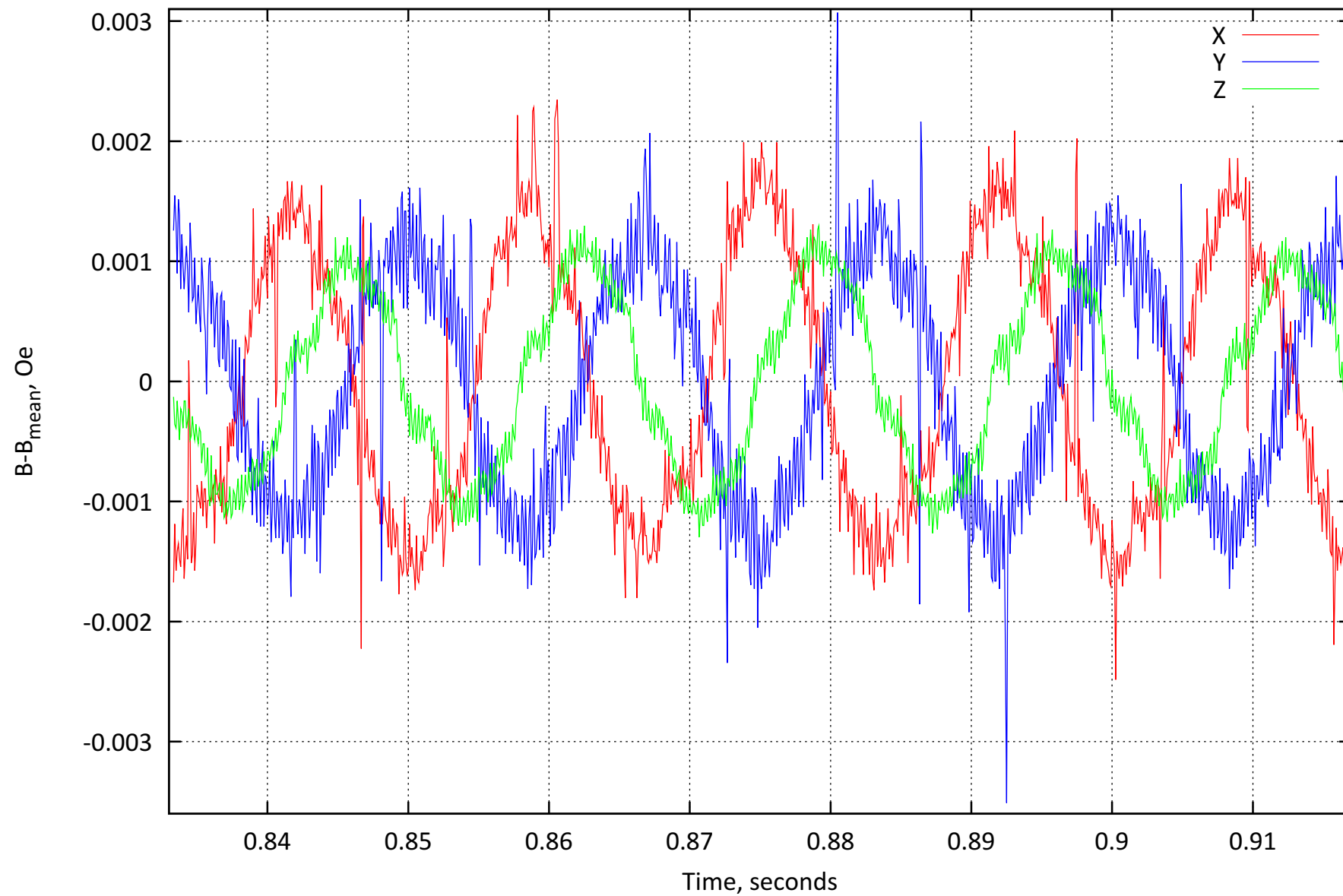
Noise in A0 north cave, klystron OFF.



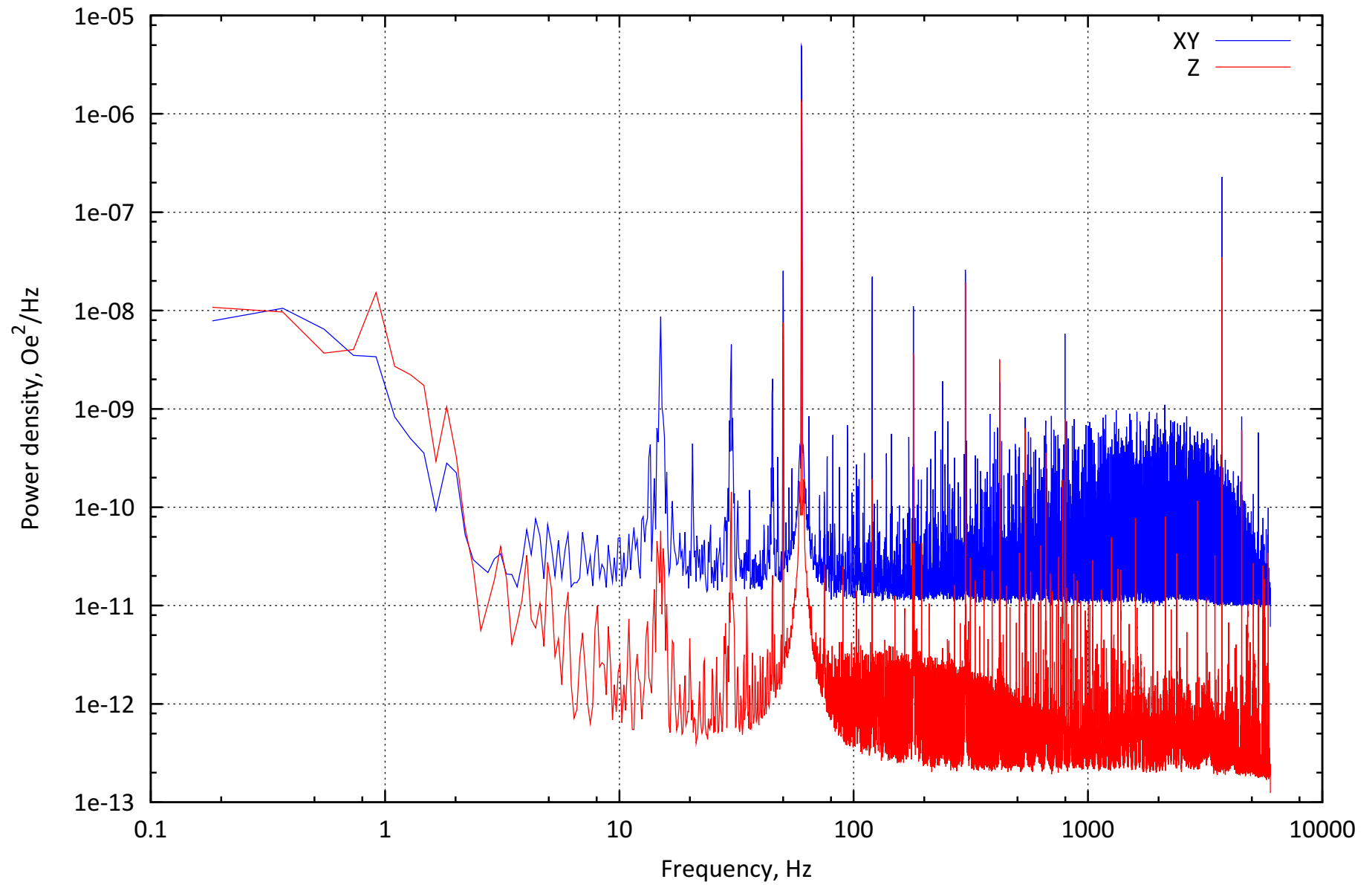
Noise in A0 north cave, klystron OFF.



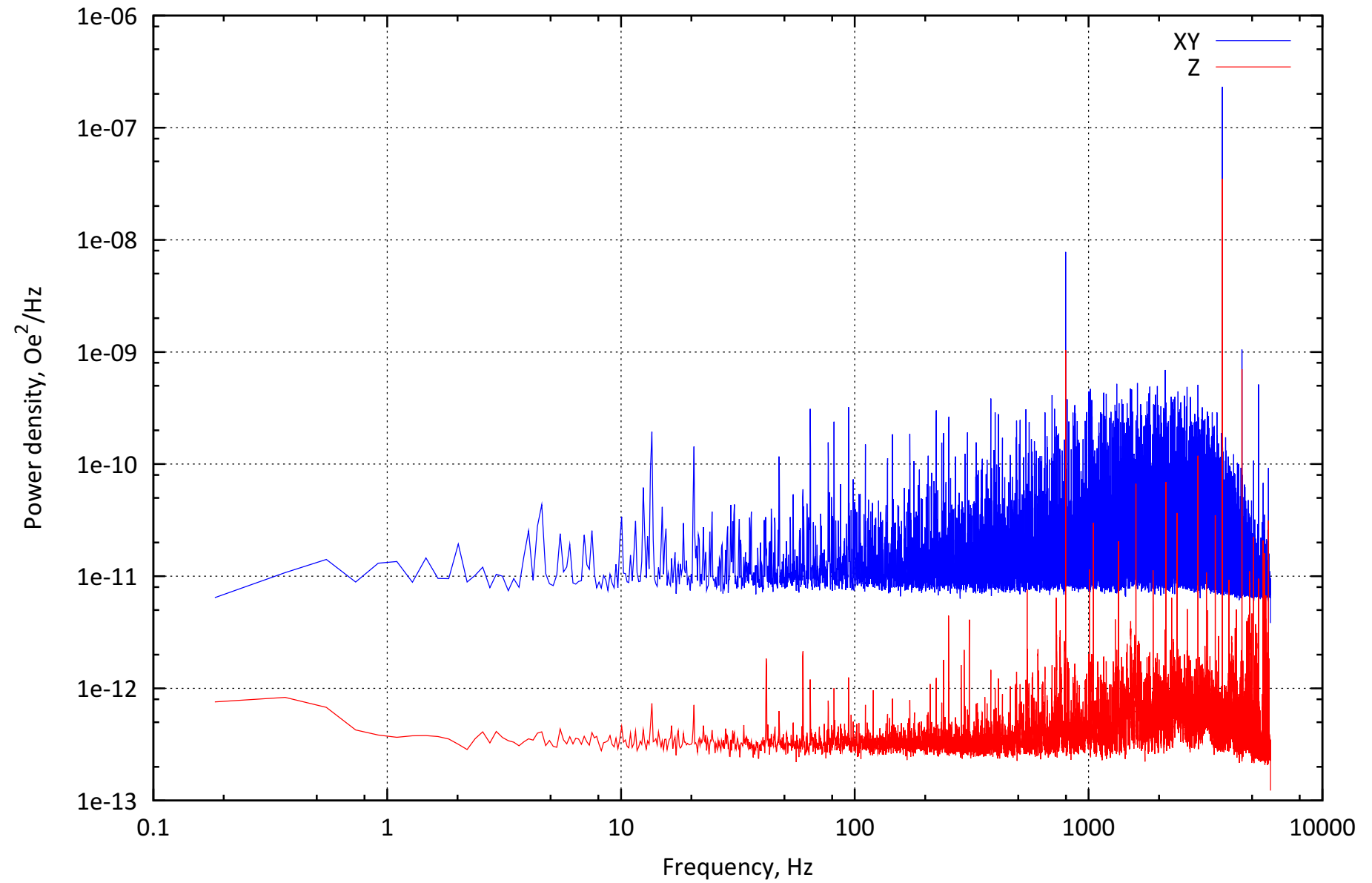
Noise in A0 north cave, klystron ON.



Noise in A0 north cave, klystron ON.



Noise in mu-metal box



Spectral energy balance sheet

	Full	60, 120, 180, 240, 300	Full -line
Mumetal box	4.15e-4	6.7105e-06 4.9973e-06 5.8313e-06 9.2630e-06 4.5480e-06	
Klystron OFF	15.3e-4	1.3732e-03 8.8196e-05 1.0643e-04 2.6315e-05 5.5453e-05	6.6763e-04 (6.1524e-04 w/o 3734 Hz)
Klystron ON	16.1e-4	1.5134e-03 7.5403e-05 7.4083e-05 2.3624e-05 1.2948e-04	5.2351e-04 (4.5304e-04, w/o 3734 Hz)

“Klystron ON” - “Line” - “Mumetal box” = 1.8e-4 0e

60 Hz issues...

- Can we *really* synchronize to 60 Hz?
- Even if we can, what about
 - Finite length of the bunch (1 msec)?
 - Finite wavelength of 60 Hz ($10 \text{ km} / 5000 \text{ km} = 2\text{e-}4$)

60 Hz – bunch length issue



Tune timing such that the bunch goes near the max or min of the 60-Hz signal. Then the middle of the bunch will see field change:

$$1 - \cos(2\pi f t / 2) \sim 0.5 (\pi f t)^2 = 0.017$$

This is the effective level of 60 Hz signal with phase lock-in (reduced to 2% of the original amplitude).

If we control 60 Hz, we cannot control other harmonics. Worst case for 3d harmonic: $2\pi(3f)t \Rightarrow A_3/A_1 = \pi f t / 12 = 0.016$

We have 0.05 to 0.07.

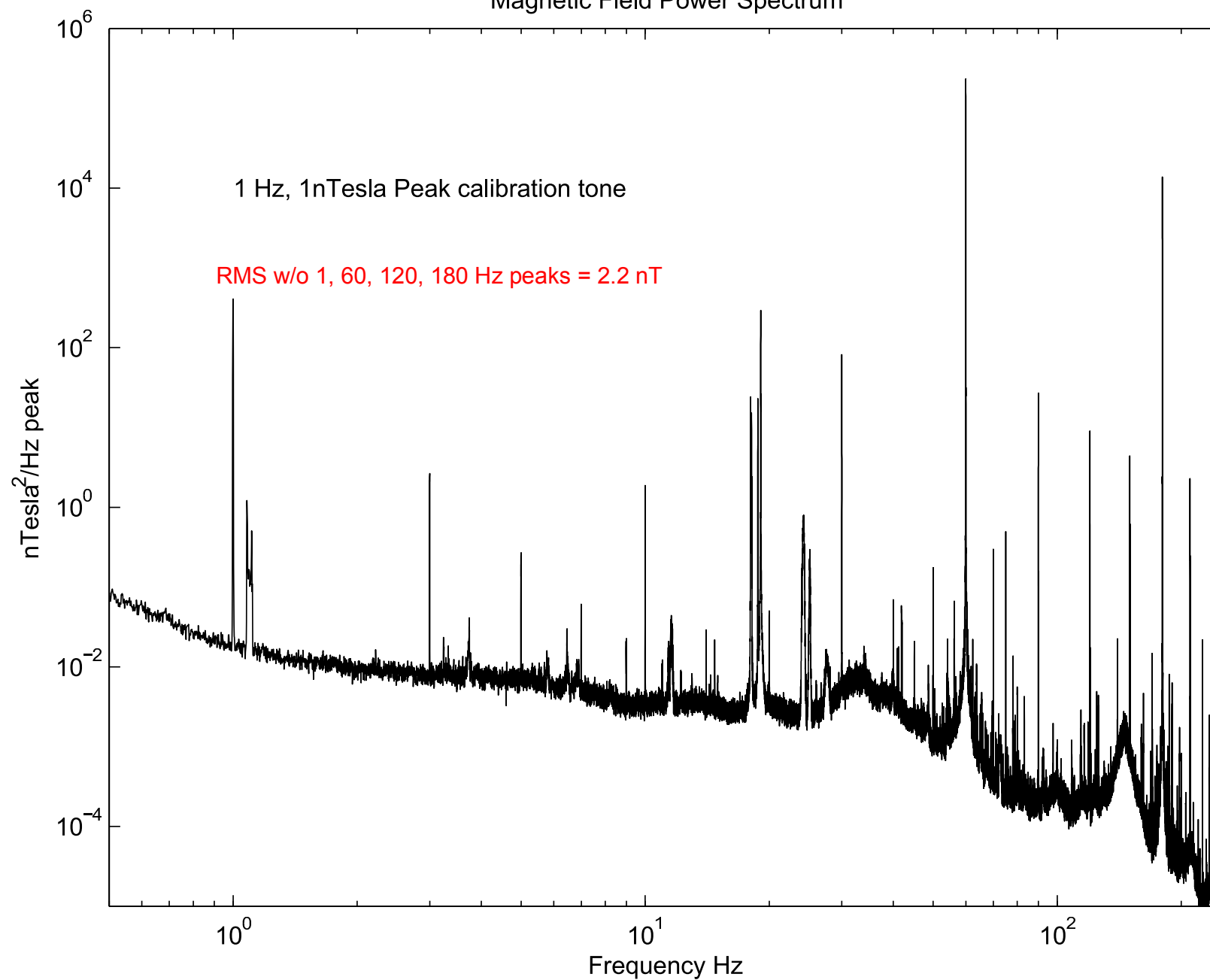
Similarly for 5th: $A_5/A_1 = \pi f t / 20 = 0.01$. **We have 0.04 to 0.09**

Future work

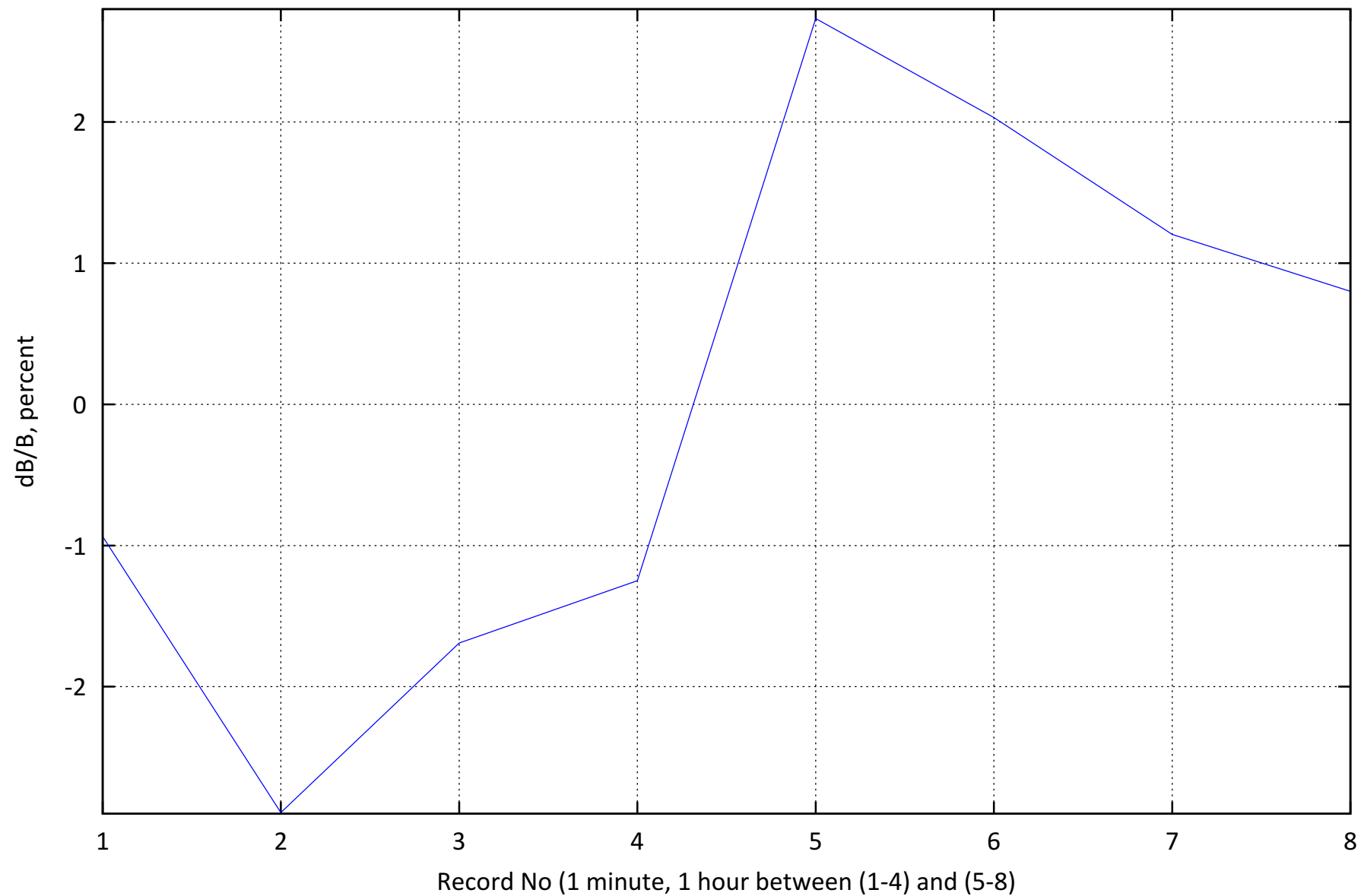
- 60 Hz is a challenge. Need to understand *finer* effects.
- We need more data!
 - Different sites; different locations on the same site.
 - Improve measurement techniques.
- Defensive design of RTML line:
 - Design in *provisions* for magnetic shielding?

Appendix

Magnetic Field Power Spectrum



Field RMS stability in mu-metal box



60-Hz amplitude stability

