Precision Measurement of the Stop Mass at the Linear Collider

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In Collaboration with Ayres Freitas, Michael Schmitt, André Sopczak Publication in Preparation

Introduction

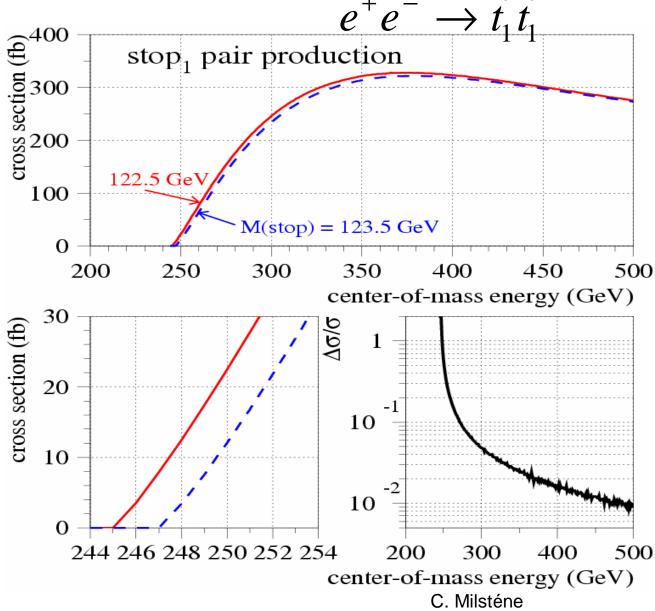
• We have previously studied the light stop, with a small mass difference to the neutralino, in an attempt to understand EW baryogenesis the asymmetry matter anti-matter and the role of the stop in dark matter annihilation.

Phys. rev. D 72,115008(2005)

M. Carena, A. Finch, A. Freitas, C. Milstene, H. Nowak, A. Sopczak
 The mass precision measurement reached was δm~1.2GeV.
 This analysis aims at the minimization of the systematics while using more realistic data, stop hadronization/fragmentation included. We will show that:

- The precision is improved in two ways:
 - a/ <u>The systematic uncertainties</u> are minimized by measuring the production cross-section at two energies \rightarrow cancellations.
 - b/ <u>The 2nd energy point chosen</u> at or close to the production energy threshold \rightarrow increased sensitivity to mass changes.
- The stop hadronization is included at production of the data → the c quark energy is spread out in the process of hadronization. As a result: the final number jets increases- the c-tagging is now <u>necessary</u> to identify the charm jets (bench-marking for the vertex detector)
- Two approaches are used, a cut based analysis, a multi-parameters optimization analysis IDA
- The polarization improves further the signal to background ratio

Cross-Section Precision In Production $e^+e^- \rightarrow \tilde{t_1}\tilde{t_1}$



Cross-sections [fb] calculated up to NLO In MC software by Freitas et al EPJ C21(2001)361, EPJ C34(2004)487

The Method

$$\sigma = \frac{N - B}{\varepsilon L}$$
$$Y(M_x \sqrt{s_{th}}) = \frac{N_{th} - B_{th}}{N_{pk} - B_{pk}} = \frac{\sigma(\sqrt{s_{th}})}{\sigma(\sqrt{s_{pk}})}$$

 σ the cross-section [fb]

N the number of selected data events

B number of estimated background events

s Square of the energy in center of Mass

 N_{th} , B_{th} , s_{th} at or close to production threshold

N_{pk}, B_{pk}, s_{pk}, at peak value

ε total efficiency & acceptance

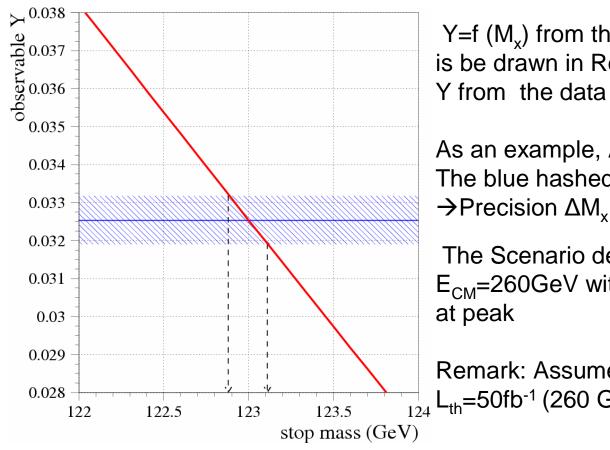
L Integrated luminosity

M_x: Mass to be determined with high precision.

Y ratio of cross-section σ_{th} and $\sigma_{pk} \rightarrow$ Allows Reduction of systematic uncertainty as well as uncertainties from L measurement.

Remark: yield close to threshold is very sensitive to $M_x \rightarrow$ choice of N_{th} and $B_{th...}$

Determination of the Stop Mass



Y=f (M_x) from the theoretical cross-section is be drawn in Red (NLO) Y from the data the blue line.

As an example, Assume 2% precision for Y, The blue hashed region \rightarrow one obtains \rightarrow Precision $\Delta M_x \sim \pm 0.1\%$ the 2 vertical arrows

The Scenario depicted: E_{CM} =260GeV with σ =9.2 fb and σ =77fb at peak

Remark: Assumed luminosities L_{124} L_{th}=50fb⁻¹ (260 GeV), L_{pk}=500fb⁻¹(500 GeV)

Theoretical Motivation

• <u>Electroweak Baryogenesis:</u> Sakharov Requirements:

 $e^+e^- \rightarrow \widetilde{t_1} \overline{\widetilde{t_1}} \rightarrow c \widetilde{\chi}_0^1 \overline{c} \widetilde{\chi}_0^1$

- 1- Baryon Number Violation (SM Anomalous process)
- 2- C & CP violation (SM-Quark CKM mixing)
- 3- Departure from Equilibrium (SM-at EW phase transition) *Limitations of SM:*
- 2)Not Enough CP violation & 3) $\rightarrow M_{Higgs}$ <40 GeV ,LEP Bound M_{Higgs} >114.4 GeV
- \rightarrow <u>Supersymmetry</u> with light scalar top, below the top mass: $m\tilde{t}_1 < mt$
- Dark Matter

The Supersymmetric Lightest particle (LSP), in the MSSM, the neutralino X_{1}^{0} is a candidate

However, the annihilation cross-section $\sigma_a (X_{1,}^0, X_1^0)$ too small

But for $m\tilde{t}_1 - m X_1^0 \sim 15-30$ GeV, there is co-annihilation between the \tilde{t}_1 and the $X_1^0 \rightarrow \sigma_a (X_1^0, \tilde{t}_1) + \sigma_a (X_1^0, X_1^0)$ consistent with dark matter.

 $e^+e^- \rightarrow \widetilde{t_1}\widetilde{t_1} \rightarrow \widetilde{\gamma_0}$

A scan in the super-symmetry parameter space (hep-ph/0403224v2-2004) C. Balazs, M. Carena, C. Wagner) Baryogenesis \rightarrow (mt̃1 <mtop && mt̃1 > 120 GeV) ;Higgs involved in the symmetry breaking mechanism mHiggs \leq 114.4 GeV \rightarrow Our points mt̃₁=122.5 GeV; mX₀¹=107.2 GeV ; Δ m=15.3 GeV

•Stop Hadronization \rightarrow the final state jets smeared :

due to Radiation + Fragmentation

Events Final State :

•Soft Multi-jets in the final state

•Stop Hadronization \rightarrow the final state jets smeared :

due to gluon radiation + fragmentation

•At ECM=260 GeV mostly 2 jets, carry the charm.

•At ECM=500 GeV 2jets \rightarrow 2,3,4 jets (more energy available in the CM)

→the Charm tagging (*T. Kuhl*) a <u>necessary tool</u>

to identify the charm jets (Vertex bench-marking)

•Analysis uses N-tuple tool incorporating jet finding algorithm (*T. Kuhl*)

Simulation Characteristics

Signal and Background generated with: Pythia (6.129)
 Simdet (4-0-3)– Circe(1.0)

- Hadronisation of the c quark and the $\,\tilde{t}$ from the Lund string fragmentation Pythia uses Peterson fragmentation

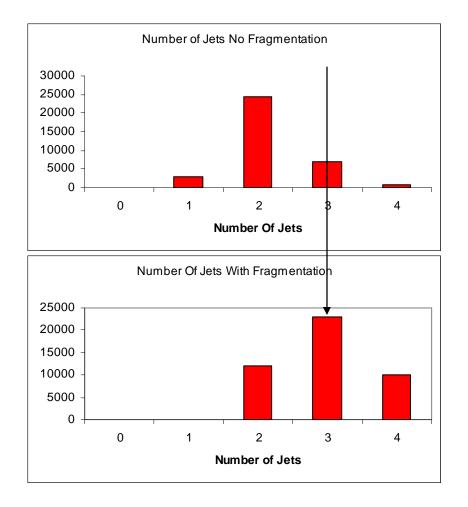
(Peterson et al PR D27:105)

- The t̃ fragmentation is simulated using Torbjorn 's code //http://www.thep.lu.se/torbjorn/pythia/main73.f

The \tilde{t}_1 quark is **set stable** until **after fragmentation** where it is Allowed to **decay again** as described in (*Kraan, EPJ C37:91*)

• Signal and Background are generated in each channel for the given luminosity in conjunction to the cross-sections

Jet Multiplicity – Without/With Fragmentation



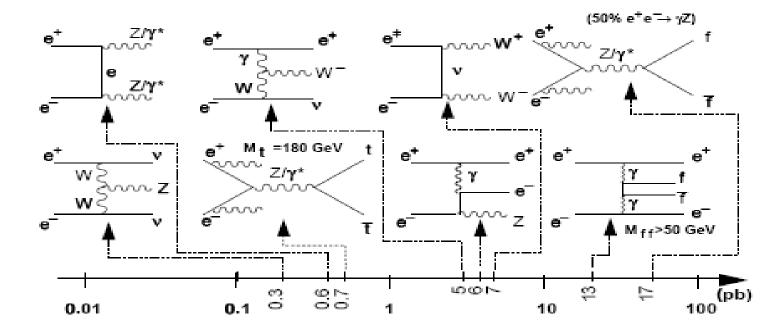
- •Stop fragmentation simulated using Torbjorn code //http://www.thep.lu.se/torbjorn/pythia/mai n73.f
- •The stop fragmentation parameter is set relative to the bottom fragmentation parameter

$$\epsilon \tilde{t} = \epsilon b^* m_b^2 / m \tilde{t}^2$$

And ε_b=-0.0050+ /- 0.0015 following (OPAL,EPJ C6:225)

- •The jet Multiplicity <u>without Fragmentation</u> Upper figure
- ~ 70% 2 jets
- •The jet Multiplicity <u>with t Fragmentation</u> Lower Figure
- ~ 50% 3 jets
- & bigger admixture of 4jets

Background- Channels @500 GeV



Z Phys. C 76 (1997) 549- A.Bartl, H. Eberl, S. Kraml, W.Majerotto, W.Porod, A. Sopczak

The cross-sections

Process	σ[pb]	at ECM=	260GeV	σ[pb]	at ECM=	500GeV
P(e-)/ P(e+)	0/0	-80%/+60%	+80%/-60%	0/0	-80%/+60%	+80%/-60%
$\tilde{t}_1 \tilde{t}_1^*$	0.032	0.017	0.077	0.118	0.072	0.276
WW	16.9	48.6	1.77	8.6	24.5	0.77
ZZ	1.12	2.28	0.99	0.49	1.02	0.44
Wenu	1.73	3.04	0.50	6.14	10.6	1.82
eeZ	5.1	6.0	4.3	7.5	8.5	6.2
qq, qq ≠ tt	49.5	92.7	53.1	13.1	25.4	14.9
tt	0.0	0.0	0.0	0.55	1.13	0.50
2γ (p _t > 5 GeV)	786			936		

Table 1

A. Freitas et al EPJ C21(2001)361, EPJ C34(2004)487 and GRACE and COMPHEP -Next to leading order, assuming a stop mixing angle (0.01)

Pre-Selection Cuts

- A short list of the sequential cuts applied as a pre-selection first, allowed larger samples to be produced
- The pre-selection cuts are the same at the 500 and 260 GeV unless listed in parenthesis for 500 GeV

Pre-selection: 260GeV ;(500 GeV)

- 4<Number of Charged tracks<50
- Pt> 5 GeV
- $\cos\theta_{\text{Thrust}} < 0.8$
- |P_I /P_{tot}|<0.9
- E_{vis} < 0.40 ECM; (E_{vis} < 0.76 ECM)
- M(inv)<200 GeV

The cuts were refined further at Selection as shown next

Selection Cuts at E_{CM} =260, 500 GeV

Variable	ECM 260 GeV	ECM 500 GeV
Number of jets	N _{jets} =2	N _{jets} ≥ 2 & E _n <25 GeV n=3,4
Transverse Momentum p_t Thrust T $cos\theta_{Thrust}$ Visible Energy E_{vis} Acoplanarity Φ_{acop} Invariant mass of jet pair m_{jj} Charm tagging likelihood P_c	$\begin{array}{l} p_t > 10 \ \text{GeV} \\ - \\ \cos\theta_{Thrust} < 0.7 \\ E_{vis} < 0.175 \ ^*\text{ECM} \\ \cos(acop) < 0.9 \\ m_{jj} < 25.5 \ \text{GeV}; \ 90 \ \text{GeV} \ >m_{jj} \\ P_c > 40\% \end{array}$	$\begin{array}{ll} p_t &> 12 \ GeV \\ T &> 0.8 \\ cos\theta_{Thrust} < 0.7 \\ E_{vis} < 0.4 \ ^{*}ECM \\ cos(acop) < 0.9 \\ m_{jj} < 60GeV; \ 90GeV \ ^{m_{jj}} \\ P_c &> 40\% \end{array}$

Table 2

In order to optimize the cancellation of the systematics we aim to have a selection as similar as possible at the two energies. (cancellation in $Y=(N_{th}-\sigma_{th})/(N_{pk}-\sigma_{pk}))$ The two-photons background did require a 5GeV pt cut.

Events Generated and After Sequential cuts

	L=50fb ⁻¹ at ECM=260GeV		L= 500fb ⁻¹ at ECM=500GeV			
P (e-)/ P(e+)	Generated	0/0	+80%/-60%	Generated	0/0	+80%/-60%
$\tilde{t}_1 \tilde{t}_1^*$	50000	382	921 <u>(24%eff.)</u>	50000	11300	26430 <u>(19%eff</u> .)
WW	180000	<5	<1	210000	102	9
ZZ	30000	<2	<2	30000	250	224
Wenu	210000	36	4	210000	10102	2994
eeZ	210000	<1	<1	210000	<18	<15
qq, q≠t	350000	<7	<8	350000	19	22
tt	-	0	0	180000	21	19
2-Photons	1.6 10 ⁶	12	12	8.5x10 ⁶	120	120

Table 3

Preliminary Results

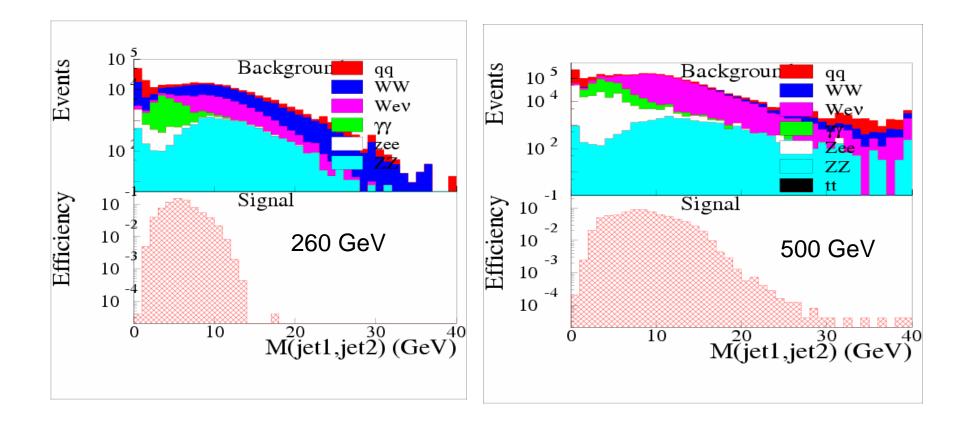
0/0 polarization beam \rightarrow Unambiguous discovery +80%/-60% polarization \rightarrow Precision Measurement

Remark: \tilde{t}_1 fragmentation \rightarrow the separation from the Wenu more difficult

Iterative Discriminant Analysis (IDA)

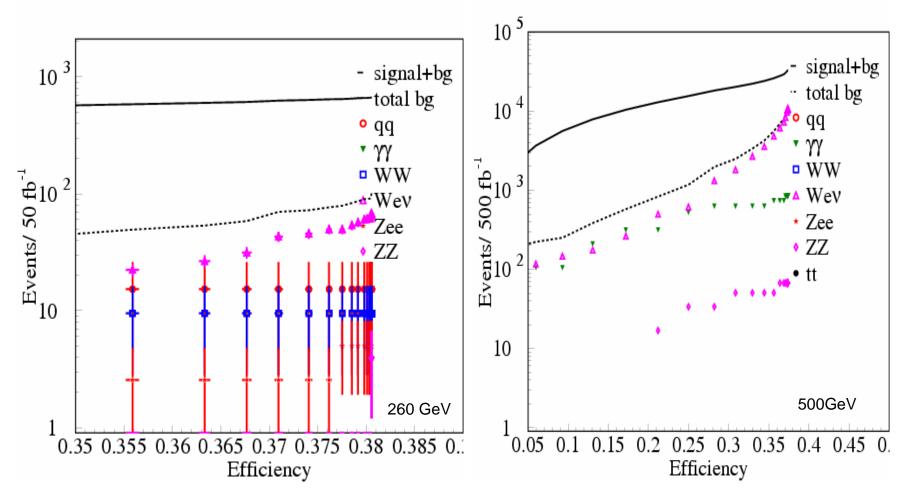
- Improves even more the precision in the \tilde{t}_1 mass measurement an Iterative Discriminant Analysis (IDA) is used. (modified Fisher Disc. Analysis)
- IDA combines the kinematic variables in parallel. The same variables and simulated events are used than in the cut based analysis . A non linear discriminant function followed by iterations are enhancing the separation between signal and background.
- Both the signal and background have been divided in two equally sized samples, one sample is used for training, the other as data.
- Two IDA steps have been performed, with a cut after the 1st IDA iteration keeping 99% of the signal efficiency.
- The performance is shown in the two next figures at 260 and 500 GeV.

Invariant Mass Di-Jets 1 Step Before Final IDA



C. Milsténe

IDA Performance



Work in Progress

Events Generated and After IDA Selection

	L=50fb ⁻¹ at ECM=260GeV		L= 500fb ⁻¹ at ECM=500GeV		
P (e-)/ P(e+)	0/0 +80%/	′-60%	0/0	+80%/-60%	
$\tilde{t}_1 \tilde{t}_1^*$	610 147	0 (38%eff.)	21240	49700	<u>(36%eff</u> .)
WW	19 2		<41	<4	
ZZ	7 7		67	60	
Wenu	68 39		10640	3155	
eeZ	10 8		<36	<30	
qq, q≠t	30 32		<38	<43	
tt	0 0		<3	<3	
2-Photons	<25 <25		840	840	

Table 4:

Preliminary Results- (In Progress)

The efficiencies improves from 24% ,19% cut based \rightarrow 38% ,36% IDA, while the background is of the same order of magnitude.

Systematic Uncertainty in Kinematics Cuts Variables

	Error on	
Variable	variable	Error on Y
p _t	2%	0.28%
cosθ _{Thrust}	1.8%	0.18%
E _{vis}	2%	0
Φ _{acop}	1%	0.08%
m _{jj}	4%	0.61%

Table 5

•All cuts are applied to hadronic and jet observables → Calibration quantities are jet energy scale & jet angle.

•Based on LEP, we assume 2% calibration error for jets, 1 deg for jet angle

•Effect on signal efficiency: Partial cancellation between 260 and 500 GeV

•We assume cancellation in total luminosity in Y between 260&500GeV

Effect of Stop and Charm Fragmentation

Comparison of the signal generated with and without gluon radiation \rightarrow The signal efficiency changes due to jet number cut is 2.5% \rightarrow We assume an error of 1% for the number of jets Charm fragmentation parameters assumed as precise as for LEP/OPAL $\rightarrow \epsilon_c = -0.0031 \pm 0.0011$ Stop fragmentation is set relative to bottom fragmentation, $\epsilon_{\tilde{t}1} = \epsilon_b (m_b/m_t)^2$ $\epsilon_{\tilde{t}1} = -0.0050 \pm 0.0015$ They don't cancel between the 2 energies but are small Including the effects of the fragmentation at both energy points $\delta \epsilon_c = \pm 35\% \rightarrow \text{Error } \delta Y = \pm 1.2\% - 0.2\%$ $\delta \epsilon_{\tilde{t}1} = \pm 30\% \rightarrow \text{Error } \delta Y = \pm 0.4\% \pm 2.4\%$ $\rightarrow \text{contribute an error O(few%)}$

Theoretical Uncertainties

- Precise cross-section calculations are needed
- t1 production receives large corrections from QCD gluon exchange Between the final state t1 (bigger @Threshold) → Coulomb corr.
- NLO- QCD corrections ~100% @threshold down to 10% at high energies are included here
- NNLO-QCD corrections are expected of to be same order than NLO based on the results for the top quark. The missing higher order correction ~7% @260GeV, 2.5% @500 GeV
- It is expected that theoretical uncertainties can be brought down by a factor 2
- Here we assume an uncertainty of 3.5% @260GeV and 1% @500 GeV
- The EW corrections : NLO ~several %, the NNLO ~1%
- Combined $\rightarrow \sim 4\%$ @260 GeV and 1.5% @500GeV $\rightarrow \delta Y=5.5\%$

Combined Statistic and systematic Errors

Error source for Y	Cut-based Analysis
Statistical	4.1%
Detector Effects	1.15%
Jet number	1%
Charm Fragmentation	1.2%
Stop Fragmentation	2.4%
Charm tagging algorithm	<0.5%
Sum of Experimental Errors	5.2%
Theory for signal σ	5.5%
Theory for background σ	0.5%
Total error δY	7.2%

For IDA the determination of *systematic uncertainties* in progress.

Table 6

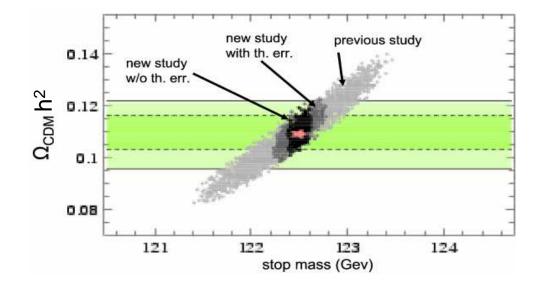
Results

Combining the statistical and systematic errors Table 6(*) $\delta Y=7.2\% \rightarrow \delta m_{\tilde{t}1} \sim 0.3 \text{ GeV} - a \text{ factor 4 better (Phys. rev. D 72,115008(2005))}$ (dominated by the theory, expected to improve for signal and background) $\delta Y=5.2\% \rightarrow \delta m_{\tilde{t}1} \sim 0.21 \text{ GeV}$ (cut based experimental errors alone) $\delta Y=4.2\% \rightarrow \delta m_{\tilde{t}1} \sim 0.17 \text{ GeV}$ (experimental errors & IDA) (expected)

→ Improvements in dark matter relic density due to improvement in $\delta m_{\tilde{t}_1}$ is shown in the next figure.

Other limiting factors start to interplay, e.g. the precision on the neutralino mass $\delta m_{\chi_1}^0 \sim 0.3 \text{ GeV}$, (hep-ph/0608255, M.Carena, A.Freitas)

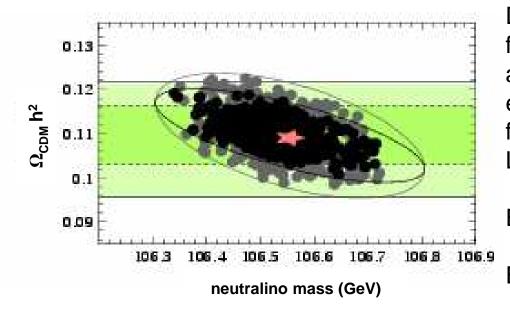
Dark Matter Relic Abundance=f (m $_{\tilde{t}1}$)



Dark Matter relic density accounting The estimated experimental errors For stop, Chargino, neutralino and Higgs sector –(scan over 1σ) versus m_{$\tilde{t}1$} for $\delta m_{\tilde{t}1}$ =1.2 GeV light gray dot Previous study $\delta m_{\tilde{t}1}$ =0.3 GeV dark gray dot Now this study $\delta m_{\tilde{t}1}$ =0.17GeV black dots Expected this study with IDA

 $\begin{array}{lll} \delta m_{\,\tilde{t}1} = 0.3 \; GeV {\begin{subarray}{c} \begin{subarray}{c} \Delta GeV {\begin{subarray}{c} \begin{subarray}{c} \begin{subarray}{c} \begin{subarray}{c} \Delta GeV {\begin{subarray}{c} \begin{subarray}{c} \Delta GeV {\begin{subarray}{c} \begin{subarray}{c} \begin{subarray}{$

Relic Abundance as Function of m_{X1}^{0}



Dark Matter relic density as a function of the neutralino mass accounting for the estimated experimental errors as before but as function of the Lightest neutralino mass $m_{\chi 0}^{1}$ Gray dots for $\delta m_{\tilde{t}1}=0.3$ This study Errors from Experiment+theory Black dots for $\delta m_{\tilde{t}1}=0.17$ This Study Experiment. Err. and IDA

 $δm_{\tilde{t}1}$ = 0.3 GeV→Ω_{CDM} h² = 0.109+0.0013-0.010 Exp. Err.+ Th. Err. $δm_{\tilde{t}1}$ = 0.17 GeV→Ω_{CDM} h² = 0.109+0.0011-0.009 Exp. Err. IDA WMAP: Ω_{CDM} h² = 0.1106+0.0056-0.0075

Conclusion

- More realistic data were produced including hadronization/fragmentation
- The precision, however, improved by a factor three on our previous analysis with $\delta m_{\tilde{t}1} = 0.3 \text{ GeV}$
- This method <u>could be applied to other particles</u> e.g. to measure the Higgs mass
- The method improves the precision to the mass determination in two ways a/ by reducing the systematics in Y- <u>cancellation</u> between the two energy points.
 b/ by choosing the energy at threshold, Y extremely <u>sensitive to the mass</u>
- The polarization separates the right-handed signal \tilde{t}_1 from background.
- Due to hadronization and fragmentation the <u>c-tagging</u> was a <u>necessary tool</u> to identify the charm jets at E_{CM}=500 GeV (benchmark for the vertex detector)
- Systematics in progress for the IDA a multi-parameters analysis, expected improvement to $\delta m_{\tilde{t}1} = 0.17 \text{ GeV}$
- Progress in the theoretical calculations is expected and partly accounted for
- With that precision we become limited by other factors.
- With this mass precision, the calculated relic density is in accordance with WMAP and SLOAN,

 $\delta m_{\tilde{t}1} = 0.15 \text{ GeV} \rightarrow \Omega \text{CDM h}2 = 0.109 + 0.0011 - 0.009$ WMAP: $\Omega \text{CDM h}^2 = 0.1106 + 0.0056 - 0.0075$

Backup slides

A Sample Parameter Point

- $m_{\tilde{U}3}^2 = -99^2 \, \text{GeV}^2$
- A_t = -1050 GeV
- $M_1 = 112.6 \text{ GeV}$
- $M_2 = 225 \text{ GeV}$
- |µ| = 320 GeV
- Φµ = 0.2
- $\tan \beta = 5$

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<u>Which gives:</u>

m\tilde{t}_1 = 122.5 \text{ GeV}; m\tilde{t}_2 = 4203 \text{ GeV};

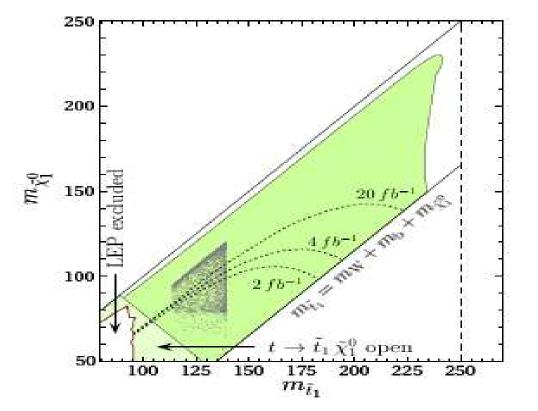
m\tilde{x}_1^0 = 107.2 \text{ GeV}; m\tilde{x}_1^+ = 194.3 \text{ GeV}; m\tilde{x}_2^0 = 196.1 \text{ GeV}

m\tilde{x}_3^0 = 325.0 \text{ GeV}; m\tilde{x}_2^+ = 359.3 \text{ GeV}

\cos\theta \tilde{t} = 0.0105 \sim \tilde{t} \text{ right-handed}

→ \Delta m = 15.2 \text{ GeV}
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Stop Discovery Reach Snowmass 2005



From Simulations: strong green region:

$$e^+e^- \rightarrow \widetilde{t_1}\overline{\widetilde{t_1}} \rightarrow c\widetilde{\chi}_0^1 \overline{c}\widetilde{\chi}_0^1$$

And Significance: $(S/\sqrt{(S+B)}) > 5$ Background B Signal S= $\epsilon\sigma L$ For ϵ , Signal efficiency For σ , Theoretical cross-section <u>dark gray region</u>: Consistent with DM And Baryogenesis

Fig 4a-<u>Luminosity: 500 fb⁻¹</u> Fig 4b-<u>Lumi. 500 fb⁻¹, 50 fb⁻¹, 10 fb⁻¹</u> E_{cm}=500 GeV

C-Tagging — The Data Samples

• Neural Network (NN):

data used: 255000 stops, Mstop=120-220; Dm=5,10, 20 GeV

240000 Wev, the most resilient background