

Effect of L^* on Collimation Depth

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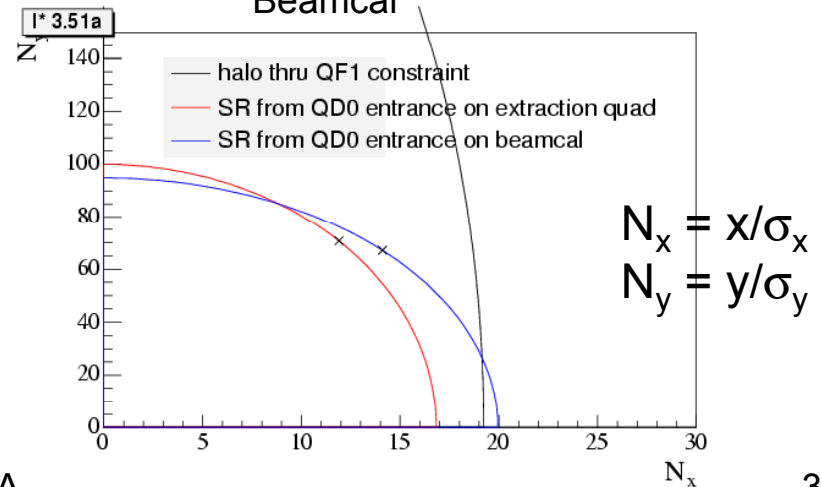
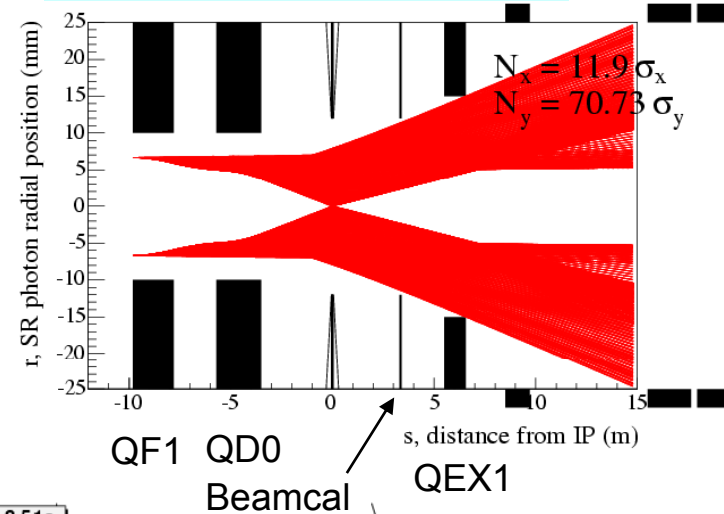
Introduction

- IR design directly determines BDS collimation depth
- In period where there is matrix of options
 - Different beam parameter sets
 - Different L^* s
 - Different detectors
- Is collimation depth reasonable over the whole space?
- What issues do we need to consider in calculation?

Quick Method Overview

- Consider halo synch radiation in FD
- Ray tracing technique
 - Semi-analytical linear optics calculation (O. Napoly)
- Gives collimated halo size at FD entrance which allows SR clearance
- Solution is not unique, ellipse in x, y space
- Important constraints
 - Extraction quads
 - Beamcal
 - QF1 quad acceptance

14 mrad RDR baseline

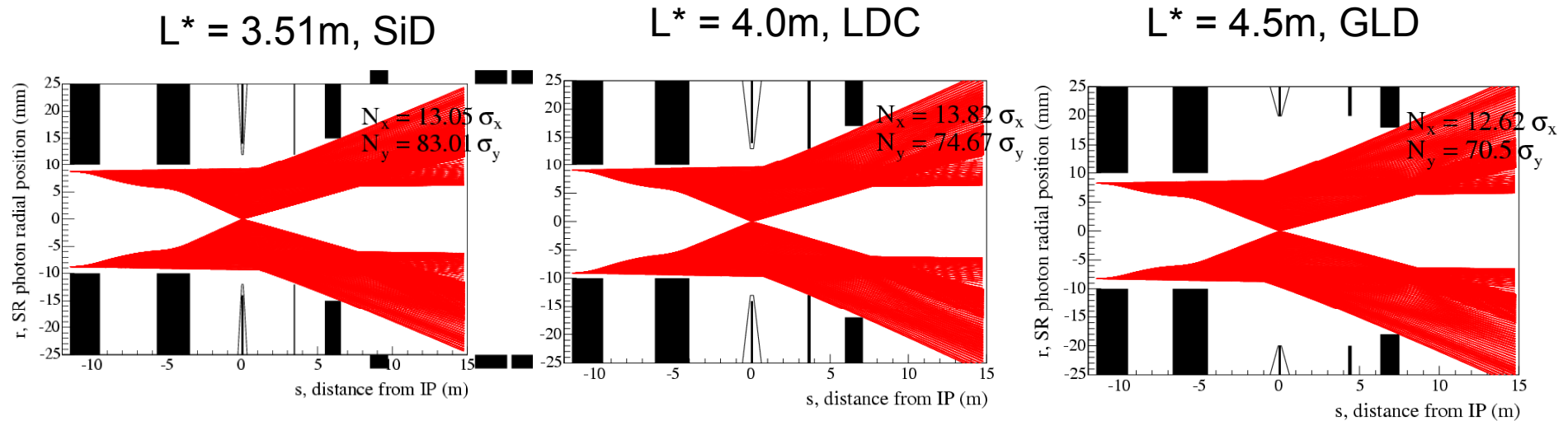


Present IR Parameter Space

- Push-pull FD+extraction designs for $L^*=3.51, 4.0, 4.5\text{m}$ (A. Seryi, Y. Nososhkov)
 - Fixed positions for QF1 and first extraction quad
 - Push-pull QD0 and QEX1 for the 3 different L^* s/detectors
 - Ex. quad positions and apertures increase with L^* (15, 17, 26mm)
- Detector beamcal 'hole' aperture from Detector Outline Documents*
- Assume nominal parameter set

Concept	Beamcal hole r, z (mm)
SID	13.5, 2950
LDC	13, 3750
GLD	20, 4500

Nominal Collimation Depths



- Naively expect collimation depths to tighten as L^* increases
- Dependence on L^* is not severe
- Wider SR fan in large L^* partially compensated by wider extraction apertures

Preliminary Conclusions

- L^* does not severely affect collimation depth
- Constraints of BeamCal, extraction apertures, QF1 acceptance all fairly close
 - Loosening one constraint does not help
 - Limited scope for loosening collimation depth by IR design
- Effect on wakefields estimate
 - RDR emittance grows 0.08% x and 4.4% y (for $\frac{1}{2} \sigma$ beam jitter, spoilers and absorber w'fields)
 - Emittance growth increases at least with the square of the collimation aperture
 - So modest changes in collimation depth become significant, e.g. $N_x=80 \rightarrow 70$ gives 30% increase in emittance growth
 - None-uniqueness of collimation depths could offset this effect (trade N_x for N_y)

Parameter Sets

- High lumi parameter sets (and others)
 - β^* are $\times 2$ smaller than nominal
 - IP divergence is $\times \sqrt{2}$ bigger
 - Collimation depths $\times \sim \sqrt{2}$ tighter
 - Wakefield emittance growth $\times \sim 2$ bigger

IR Beam Orbit

- Detector field correction schemes (anti-solenoids, DID, Anti-DID) perturb the beam orbit and direction of the SR rays
- Max orbit perturbations of the order $\sim 100 \mu\text{m}$, $100 \mu\text{rad}$ (my guesses)
- Could lead to $\sim 1 \text{ mm}$ deviations in SR rays at apertures

Other Issues

- Margins – how much SR can we tolerate on apertures?
- Realistic beams and IR geometry
 - Energy spread, jitter, halo population
 - Magnet and mask misalignment, beampipe thickness
- Is it possible (or worthwhile) to include precise estimates of all effects – or only consider worst-case scenarios/biggest effects?

Reference

Reference Information

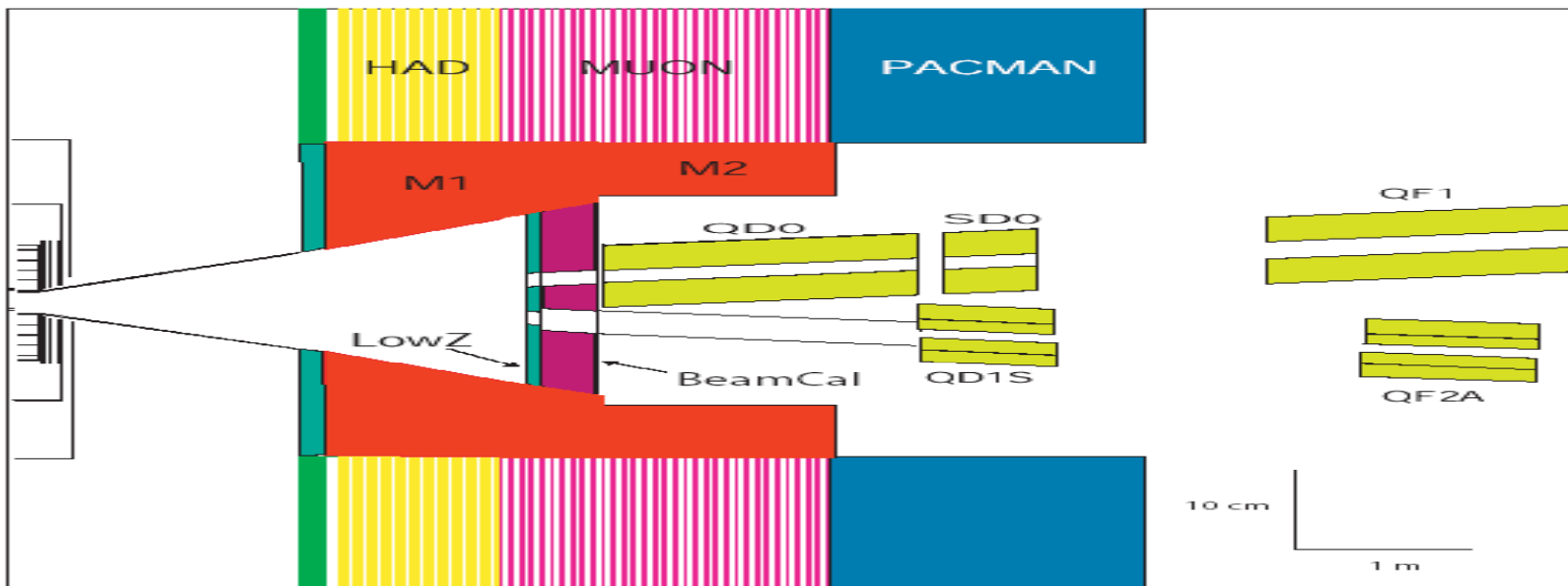
- Collimation half gaps 1.3 mm in x 0.7 mm in y
- Emittance growth much larger in y.
 - Spoilers gaps smaller, beta functions 30% larger
 - Phase relationships of y spoilers to IP
 - Energy spoiler has vertical gap and large beta function (compared to x)

BeamCal Info

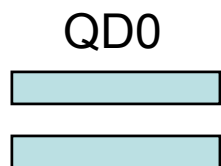
- SiD $r=13.5$ mm at 2.95 m
- PTO for slide from SLAC ILC BDS weekly meeting 08 April 07 T. Maruyama

14 mrad crossing geometry

14 mrad crossing geometry in Geant 3 and FLUKA

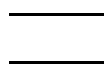


Apertures:



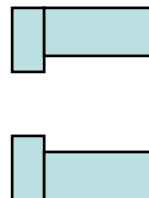
$R=1.0\text{ cm}@z=-3.51\text{ m}$

Beampipe@IP



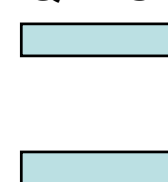
$1.2\text{ cm}@0.0\text{m}$

Low Z



1.35 cm
 $@2.85-2.95\text{m}$

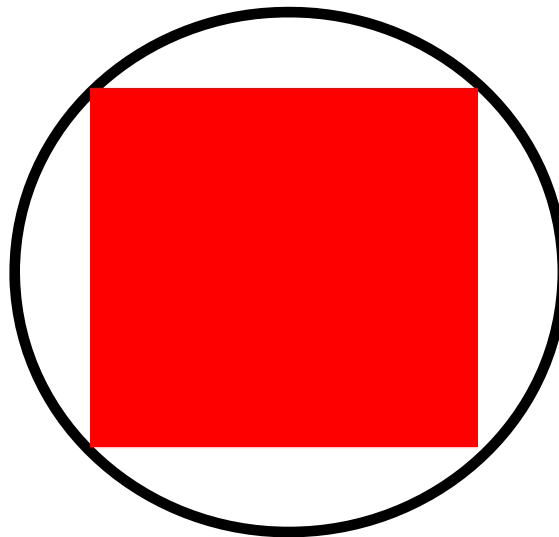
QD1S



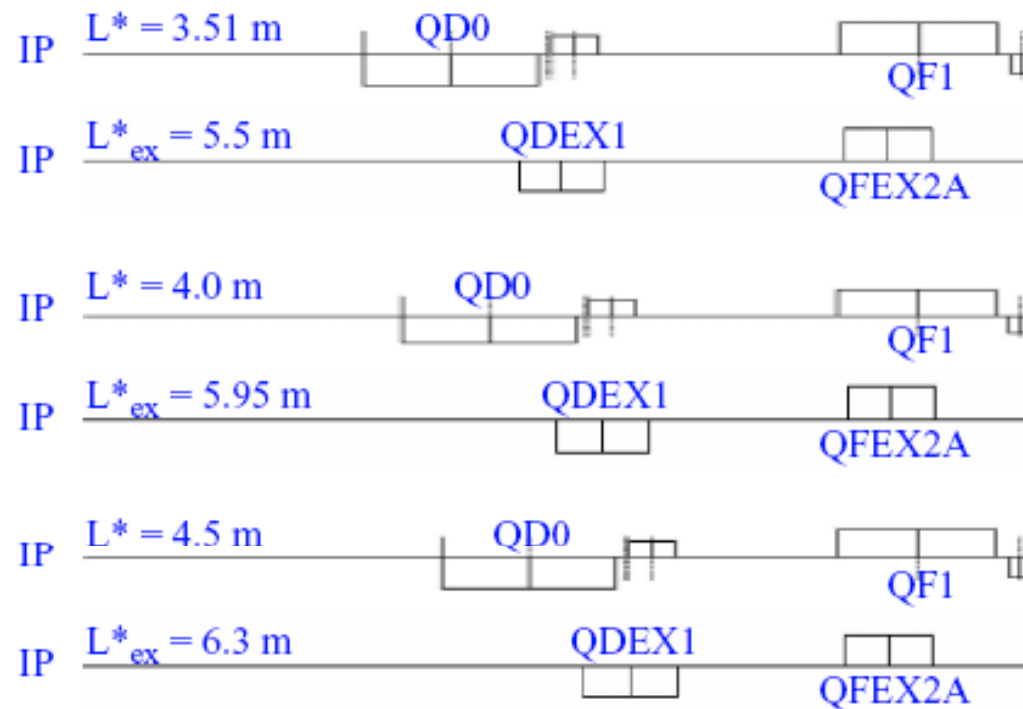
1.5 cm
 $@5.5-6.56\text{m}_{13}$

Why choose just one collimation depth ...

- ... when there is an ellipse of solutions?
- The one chosen is so that the SR fan is a square fitting inside the circular aperture constraint



Push Pull Schemes



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