

Supersymmetry Without Prejudice

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- The MSSM has ~ 140 parameters
- Most studies/searches incorporate simplified version
 - mSUGRA with 5 parameters
- Does mSUGRA adequately describe all possible SUSY effects? **-No!-**
- The LHC is turning on, era of speculation will end, and we need to be ready for all possible signals
- Ready to determine underlying physics from LHC data
- Study the consequences of a more general set of MSSM models

SUSY Parameter Determination @ ILC

Berger, Gainer, Hewett, Lillie, Rizzo
arXiv:0711.1374, 0712.2965

Analysis:

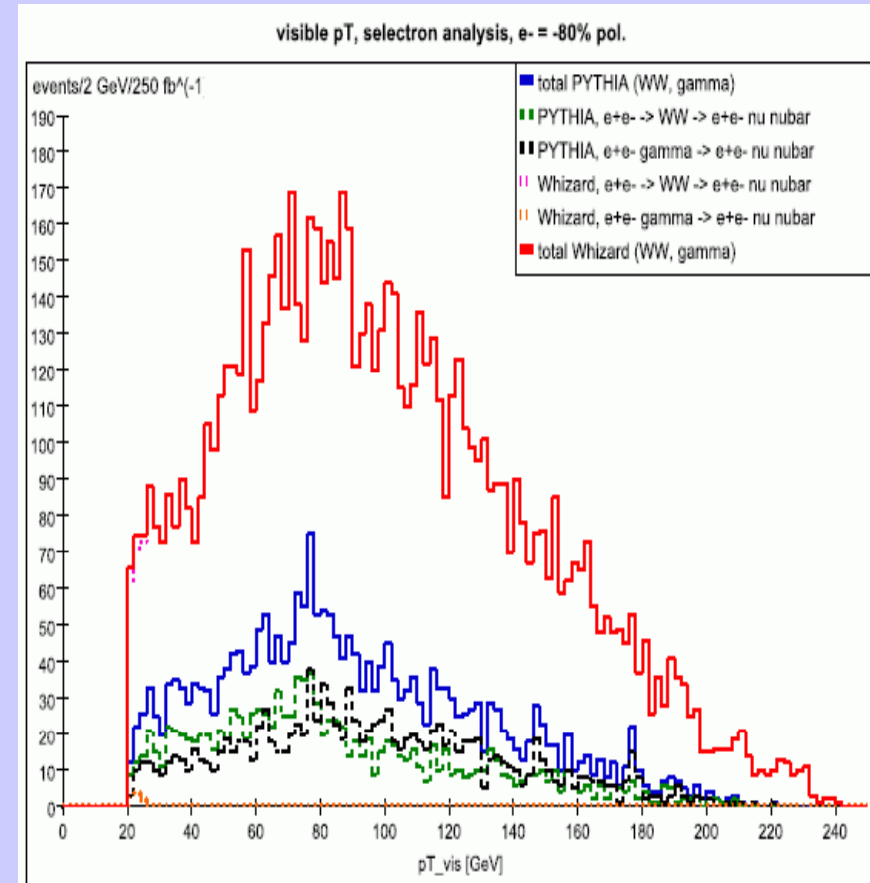
- Study 242 generic models which appear in 162 pairings found to be indistinguishable @ LHC (Arkani-Hamed et al)
- 12 simultaneous SUSY channels (Pythia & CompHEP)
- Full SM background,
 - 1016 processes w/ full matrix elements (Whizard)
- ISR, ILC-design Beamstrahlung, Beam energy spread
- SiD fast detector simulation
- Analyze 500 fb^{-1} “data” at 500 GeV with 80% P_{e^-} and appropriate cuts
- 1 CPU century to complete analysis!

Several iterations necessary to find best (model independent) kinematic cuts!

General Conclusions

- Generic SUSY signal smaller than SPS1 a
- Many cuts designed for SPS1 a(LEP II) kill generic SUSY signal
- Larger SM background than previous studies
- Forward detector coverage critical
- Some difficult cases:
 - close stau-LSP mass,
 - $\chi_1^\pm \rightarrow W^* \chi_1^0 \rightarrow jj \chi_1^0$,
 - Neutralino Production

Random SUSY signal is not a piece of cake!

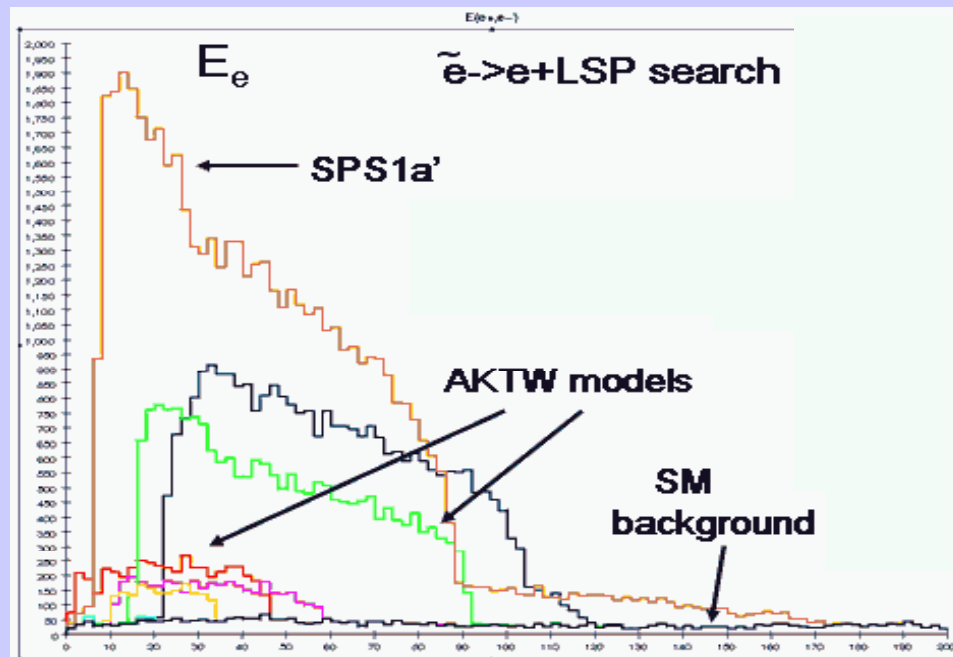
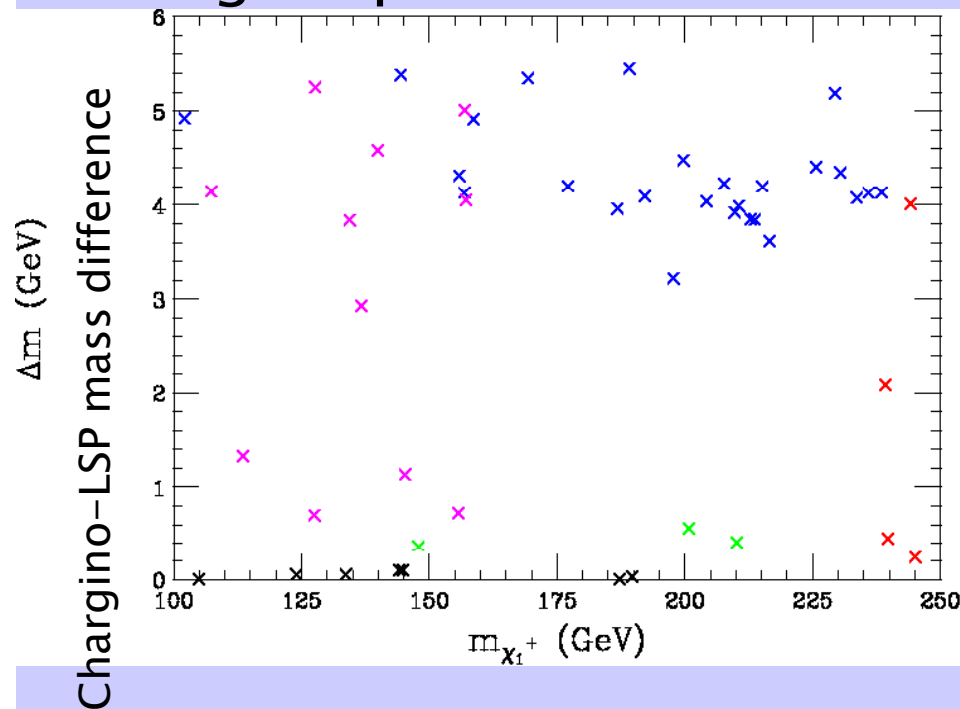


Sample Results

Selectron Production

Generic signal
is much
smaller than
SPS1a'

Chargino production



Many times charginos have small mass splitting with the LSP requiring many different searches: stable particles, photon tagging, soft jets, or a combination. Four are missed due to tiny phase space

Overview of Results

Visibility: ILC can see

- 78/85 AKTW models w/ at least one charged sparticle
- 17/96 models w/ neutral sparticles only
- 82/161 models w/ any accessible sparticle
- 82/242 of all models

Particle	Number Visible
\tilde{e}_L	8/9
\tilde{e}_R	12/15
$\tilde{\mu}_L$	9/9
$\tilde{\mu}_R$	12/15
$\tilde{\tau}_{1,2}$	21/28
$\tilde{\nu}_{e,\mu}$	0/11
$\tilde{\nu}_\tau$	0/18
$\tilde{\chi}_1^\pm$	49/53
$\tilde{\chi}_1^0$	17/180
$\tilde{\chi}_2^0$	5/46

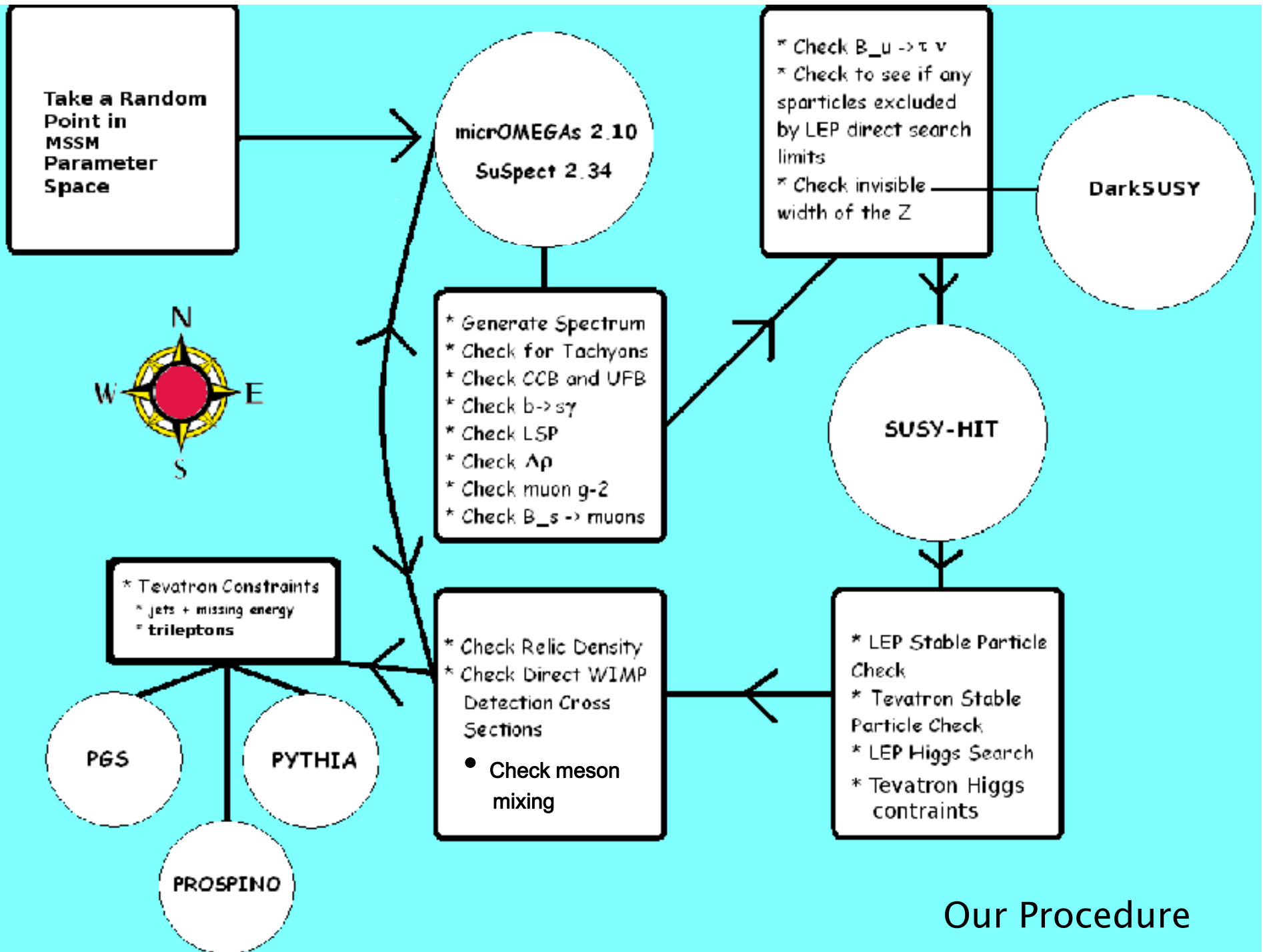
Model Distinguishability:

- 57(63)/72 pairs w/ at least one charged sparticle at $5(3)\sigma$
- 0/90 model pairs with only neutral spartners accessible
- 57(63)/162 of all model pairs

MSSM Model Generation

Berger, Gainer, Hewett, Rizzo
In progress

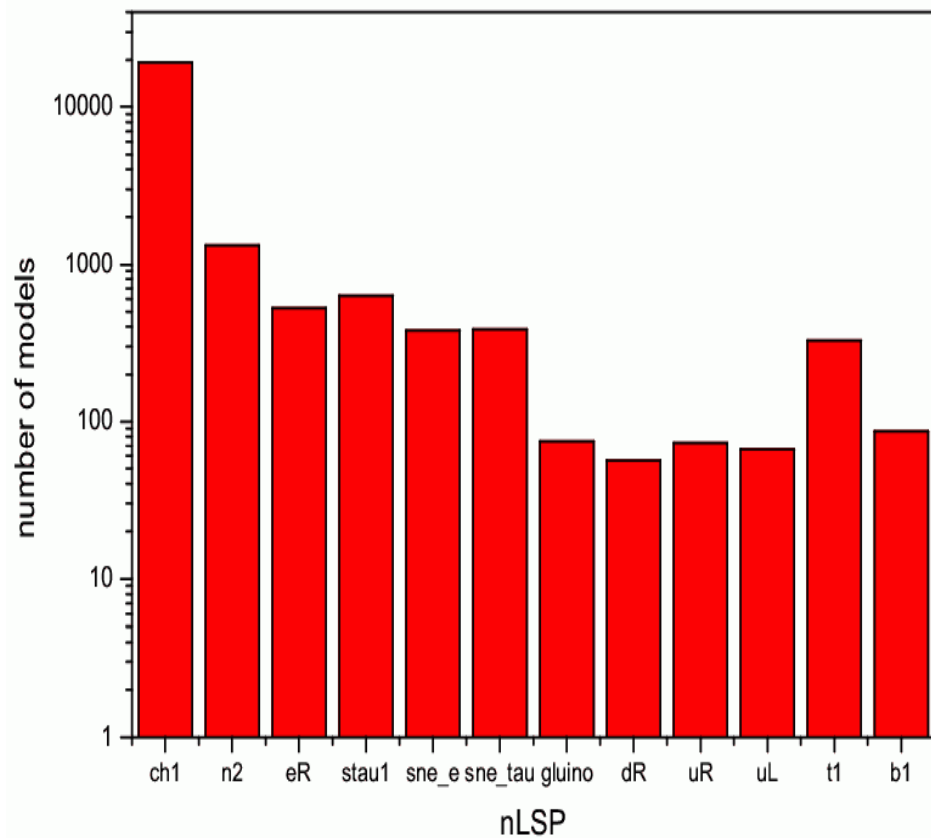
- Goal: Generate 1000's of MSSM models, consistent with all known data, not tied to any specific SUSY breaking mechanism, and study them
- Scan 19-parameter CP-conserving MSSM (the PMSSM), accessible to LHC
- Expect to populate new regions of parameter space with different or potentially difficult signatures
- Using this model set, we will perform **model independent** collider studies (LHC & ILC), make predictions for direct & indirect **DM** searches (GLAST), and more
- We will make this model set public



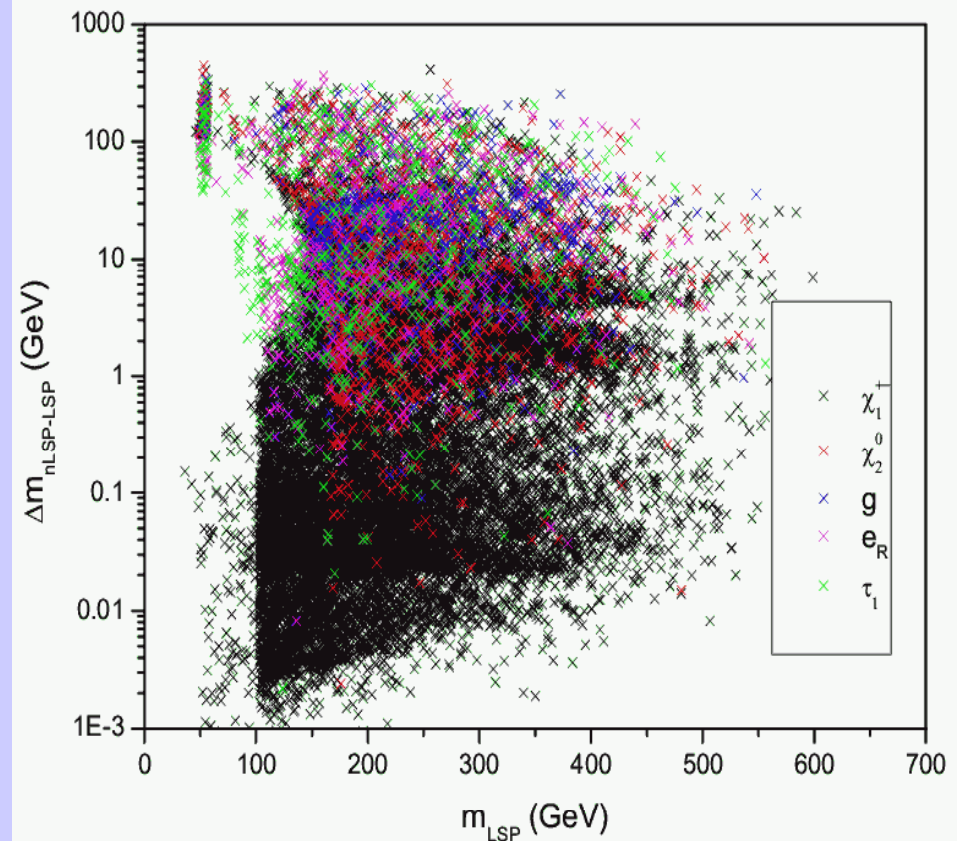
Preliminary Results: Model Properties

23,652 models have passed the constraints thus far....(more to come!)

NLSP

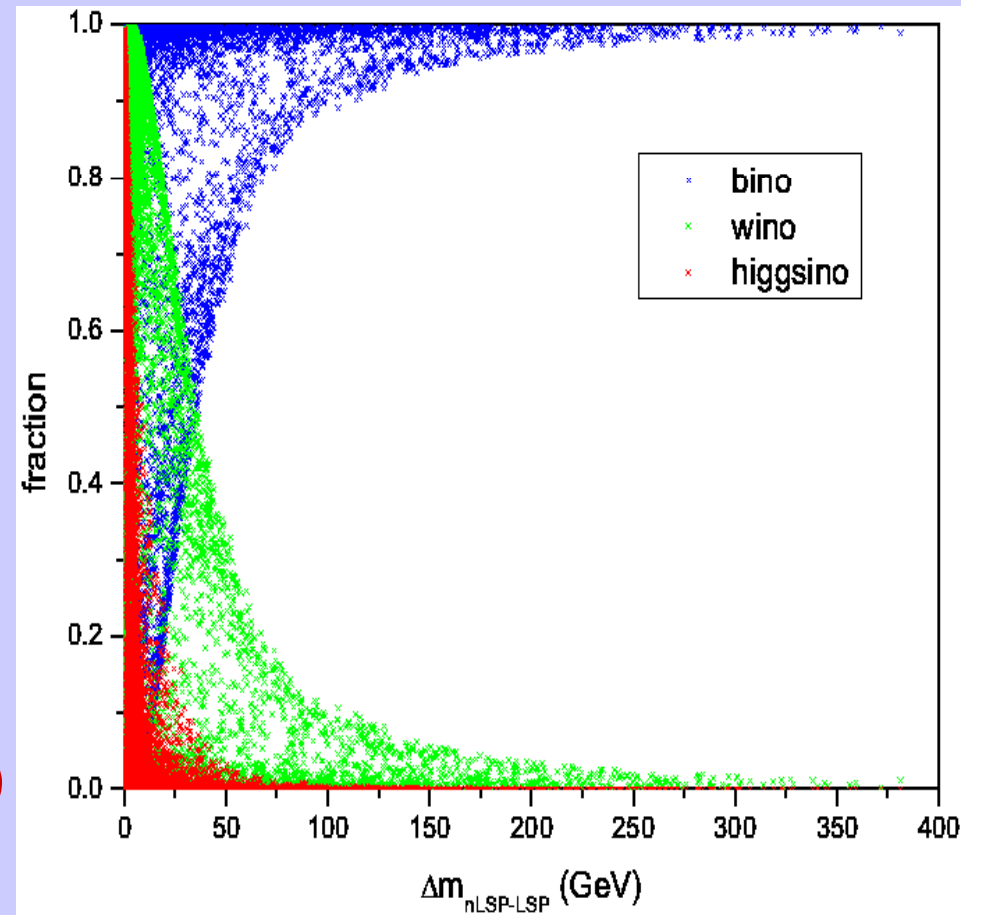
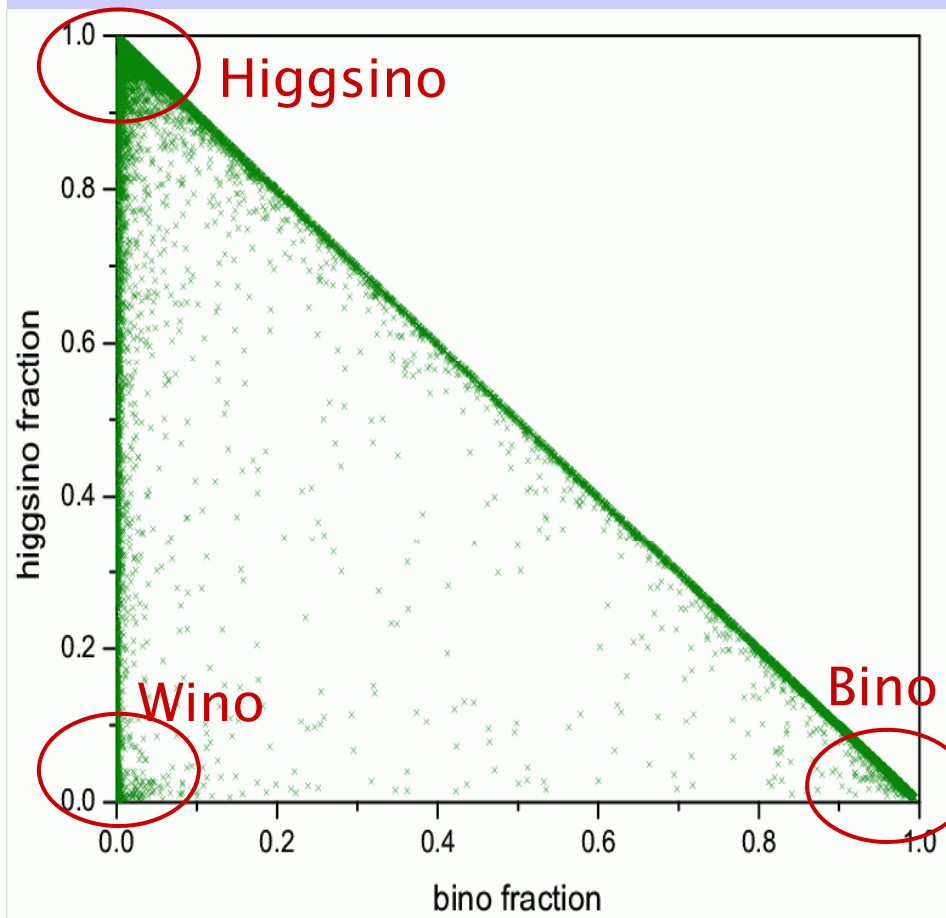


NLSP - LSP mass difference

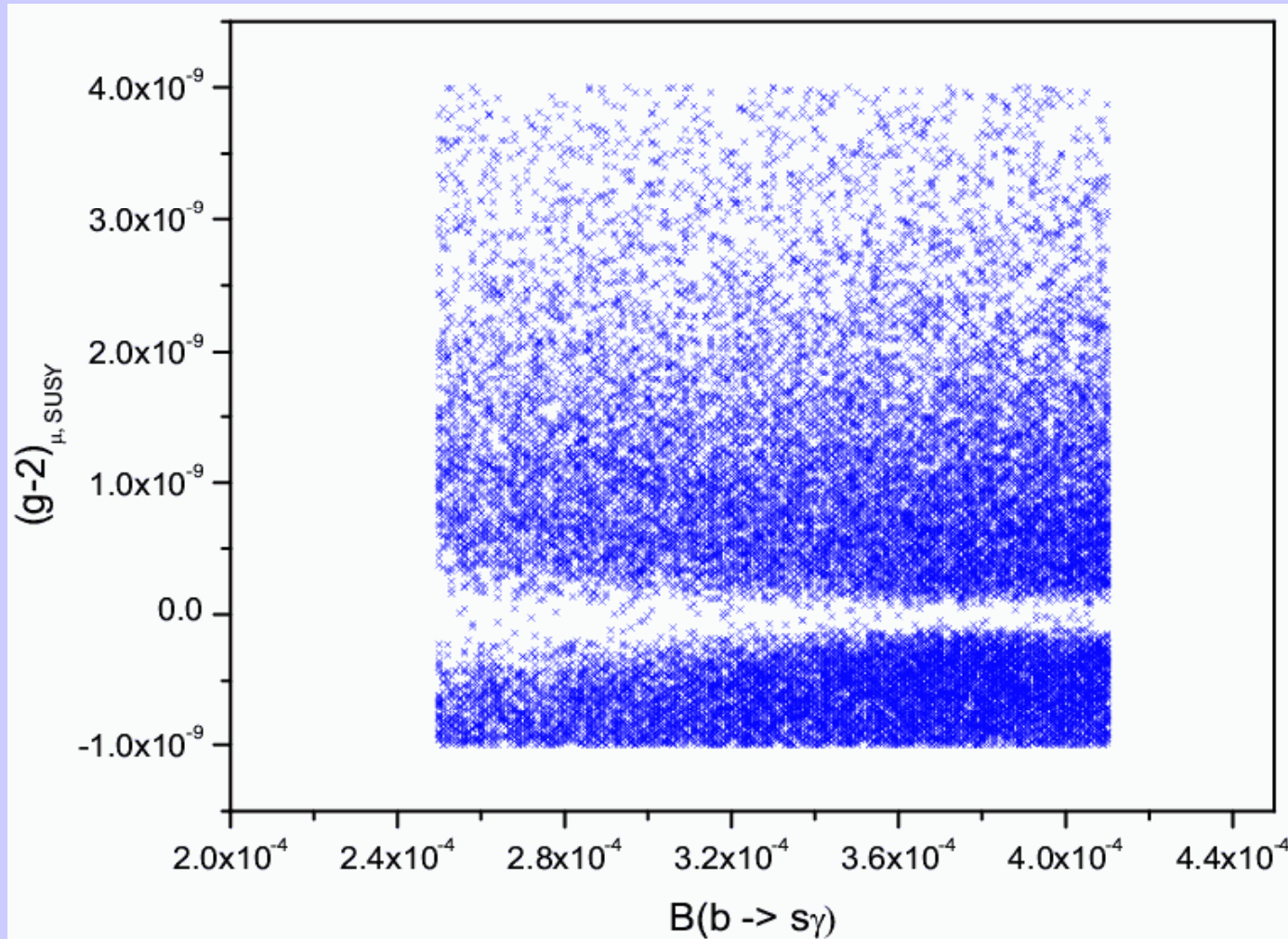


Many models with tiny mass differences

Preliminary Results: LSP Fractional Content



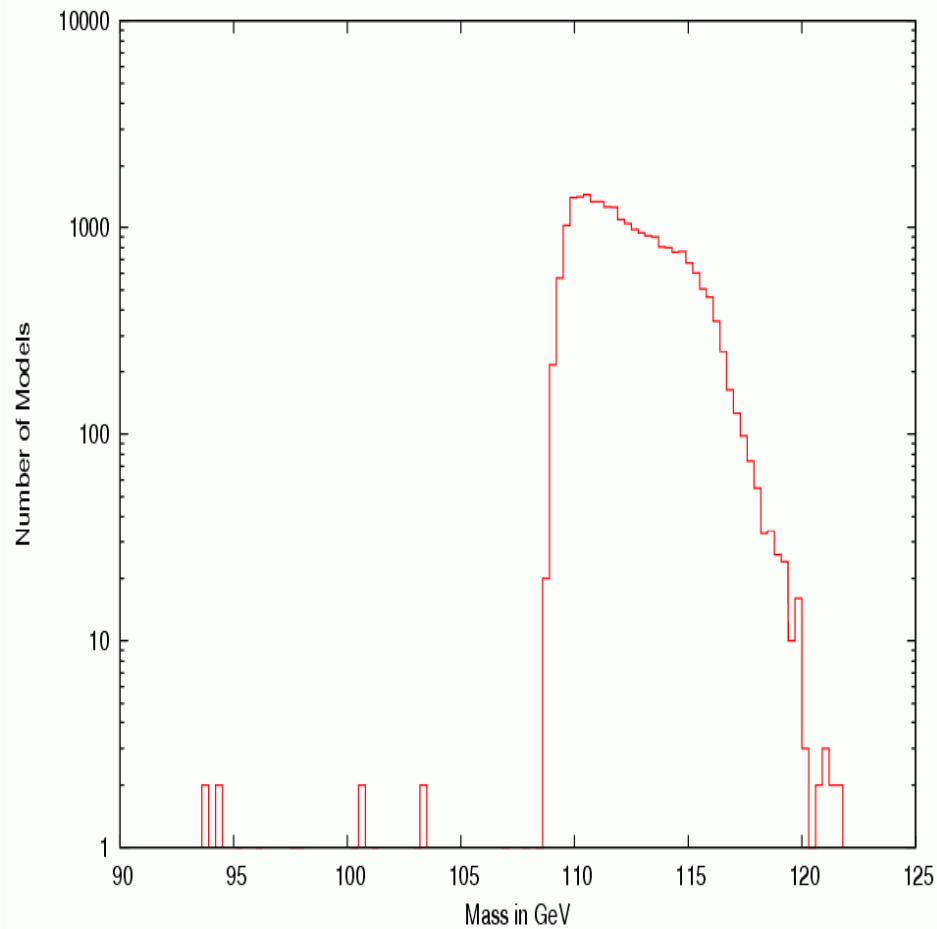
Preliminary Results: Exp't Constraints



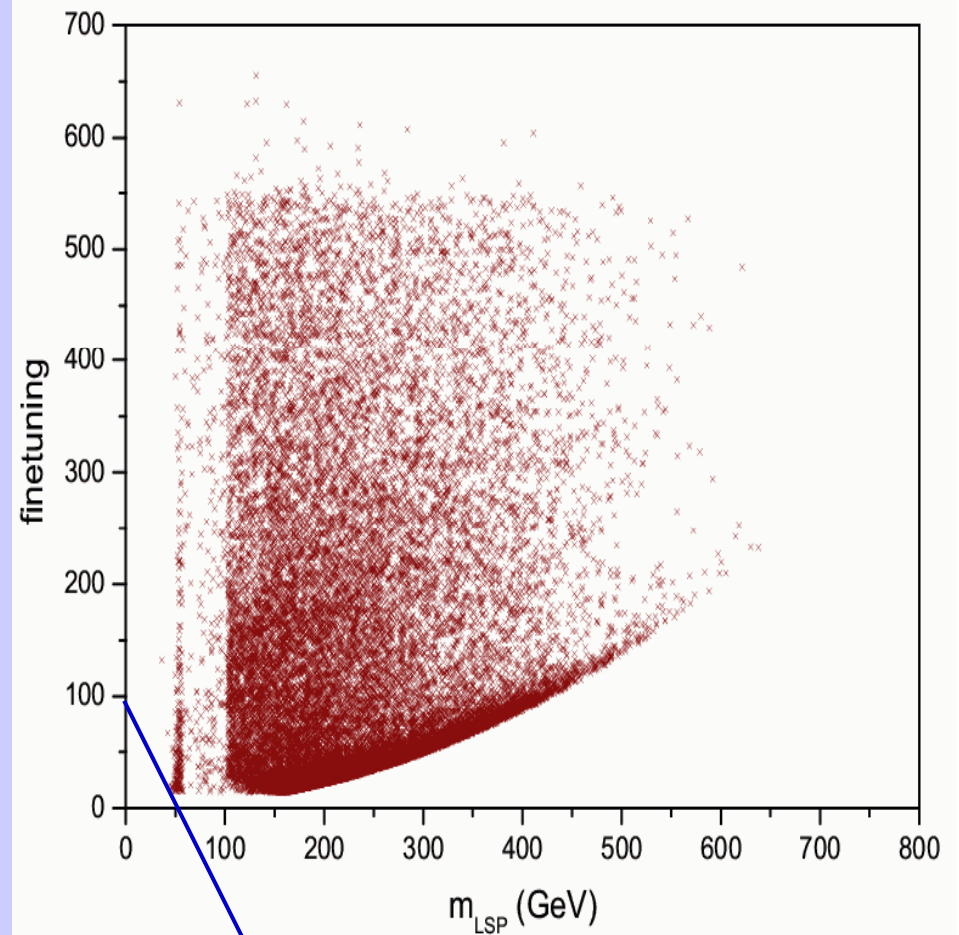
No correlation

Preliminary Results: Model Properties

Higgs mass prediction

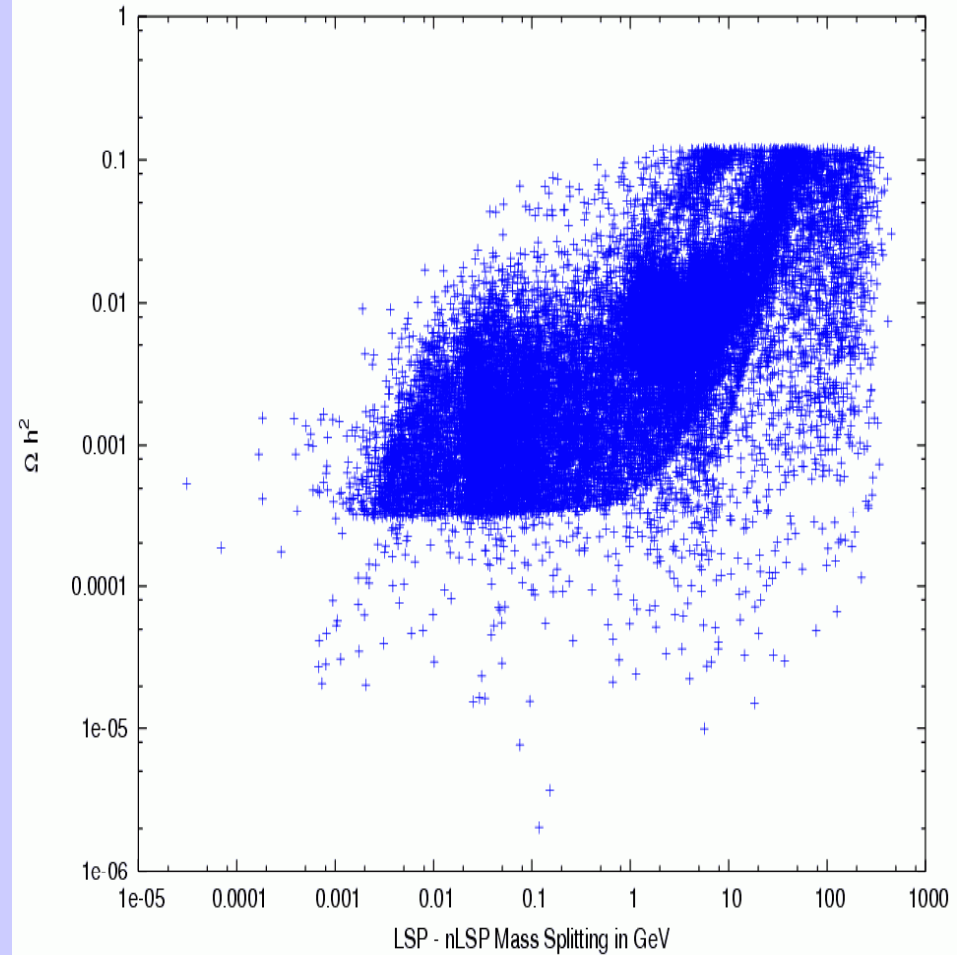
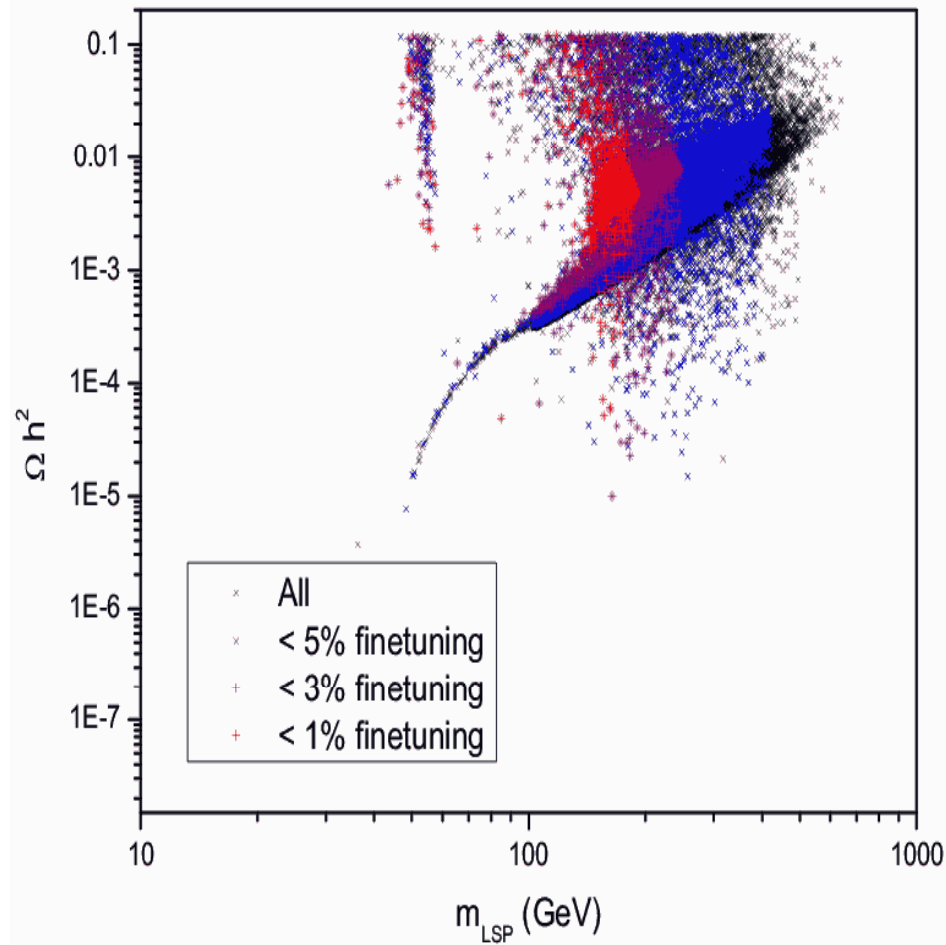


Fine-tuning



1% fine-tuned

Preliminary Results: Dark Matter



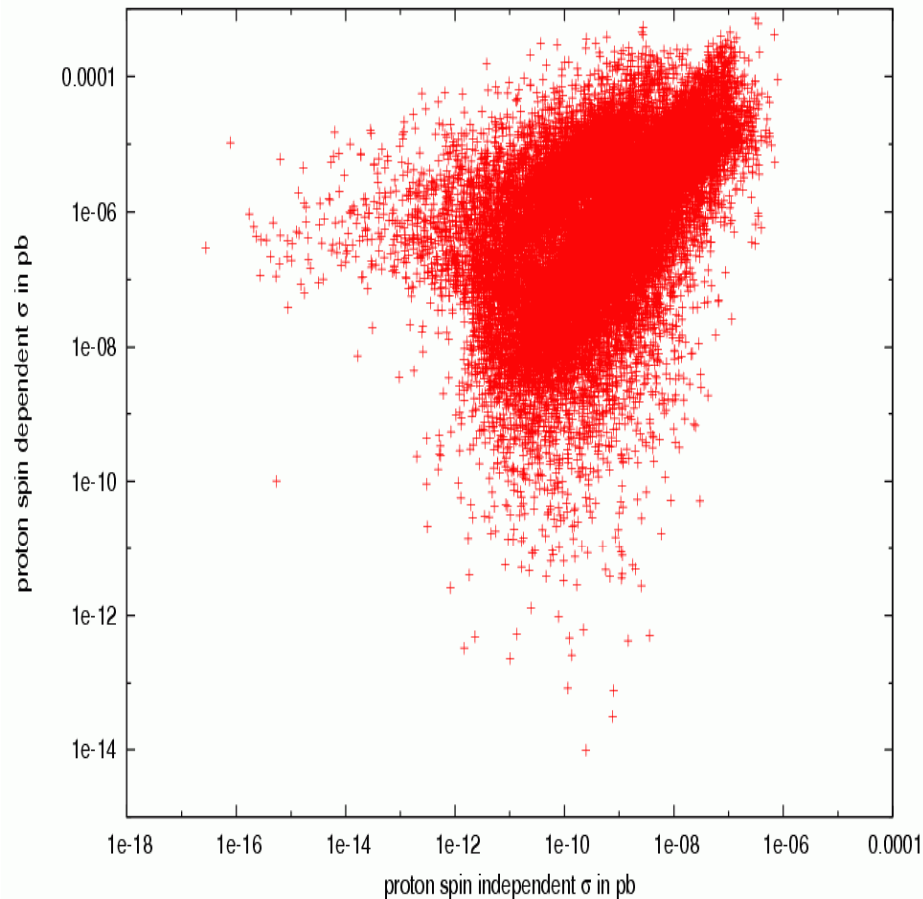
Preliminary Results: Dark Matter

Spin dependent vs independent Direct Detection cross sections

Barger et al
arXiv:0806.1962

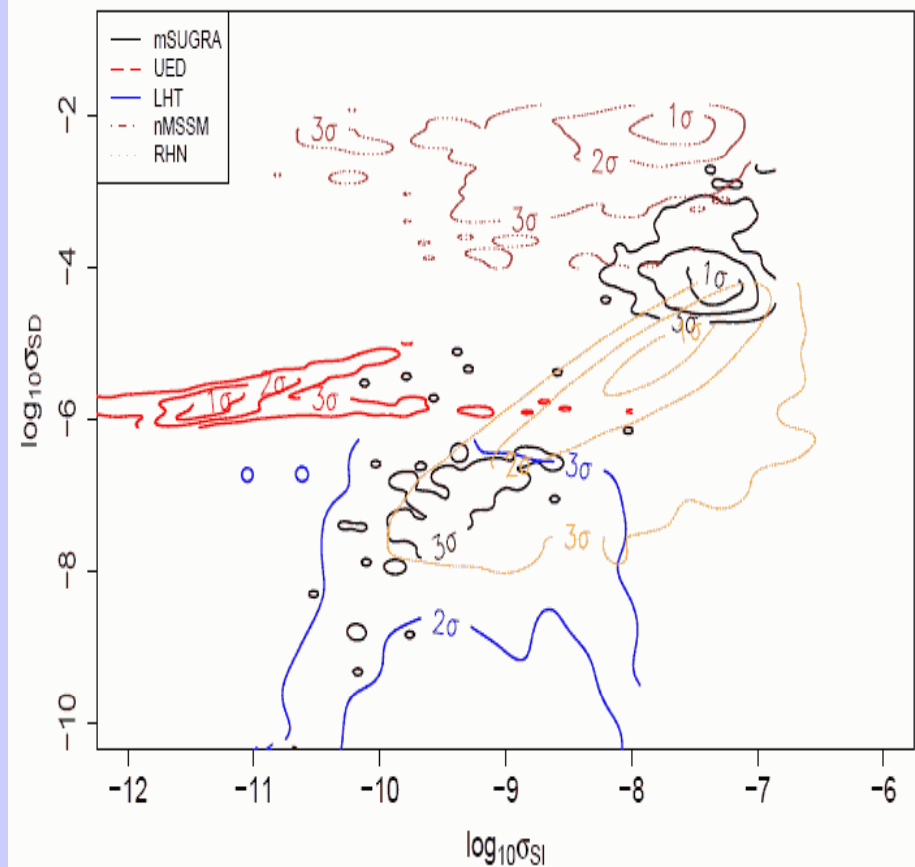
PMSSM

Spin Independent and Spin Dependent WIMP - Proton Cross Sections



Various Models

SI vs. SD



Summary

- MSSM @ ILC is more difficult than previously thought
- More to come for the PMSSM analysis
- The general MSSM can behave very differently than mSUGRA
- The LHC era is starting and we need to be prepared!