

# Precise Prediction of $M_W$ in the MSSM

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Vancouver, 07/2006

based on collaboration with

*W. Hollik, D. Stöckinger, A.M. Weber and G. Weiglein*

1. Motivation
2. Calculation of  $M_W$  in the MSSM
3. Numerical results
4. Remaining theoretical uncertainties
5. Conclusions

## 1. Motivation

$M_W$  is a precision observable

Experimental situation:

today: LEP2, Tevatron  $\Rightarrow \delta M_W^{\text{exp}} = 32 \text{ MeV}$   
Tevatron (8 if $-1$ )  $\Rightarrow \delta M_W^{\text{exp}} = 20 \text{ MeV}$   
LHC  $\Rightarrow \delta M_W^{\text{exp}} = 15 \text{ MeV}$   
ILC  $\Rightarrow \delta M_W^{\text{exp}} = 10 \text{ MeV}$   
GigaZ  $\Rightarrow \delta M_W^{\text{exp}} = 7 \text{ MeV}$

Prediction in the SM:

$$\delta M_W^{\text{theory,SM,today}} \approx \pm 4 \text{ MeV}$$

$$\delta M_W^{\text{theory,SM,future}} \approx \pm 2 \text{ MeV}$$

$$\delta m_t = 2.3 \text{ GeV} : \quad \delta M_W^{m_t, \text{para,today}} \approx \pm 14 \text{ MeV}$$

$$\delta m_t = 0.1 \text{ GeV} : \quad \delta M_W^{m_t, \text{para,future}} \approx \pm 1 \text{ MeV}$$

$$\delta(\Delta\alpha_{\text{had}}) : \quad \delta M_W^{\Delta\alpha, \text{para,today}} \approx \pm 6.5 \text{ MeV}$$

$$\delta(\Delta\alpha_{\text{had}}) : \quad \delta M_W^{\Delta\alpha, \text{para,future}} \approx \pm 1 \text{ MeV}$$

## 1. Motivation

$M_W$  is a precision observable

Experimental situation:

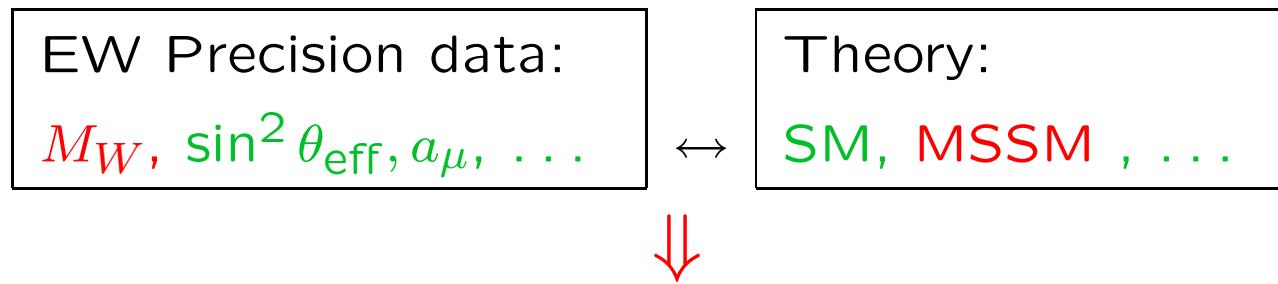
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Prediction in the SM:

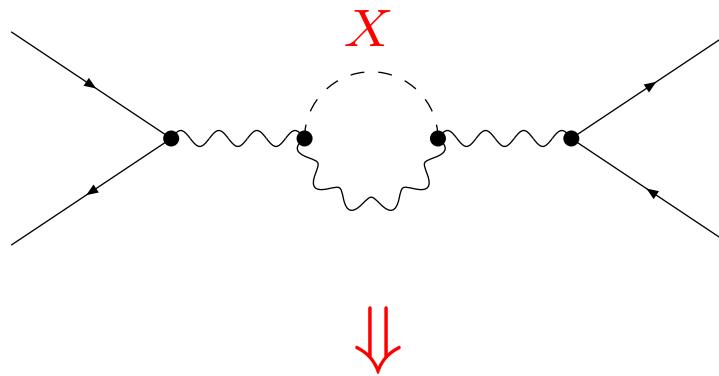
$\delta m_t = 2.3 \text{ GeV} :$	$\delta M_W^{m_t, \text{para}, \text{today}}$	$\approx \pm 14 \text{ MeV}$	<b>SUSY?</b>
$\delta m_t = 0.1 \text{ GeV} :$	$\delta M_W^{m_t, \text{para}, \text{future}}$	$\approx \pm 1 \text{ MeV}$	
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## Precision Observables (POs):

Comparison of electro-weak precision observables with theory:



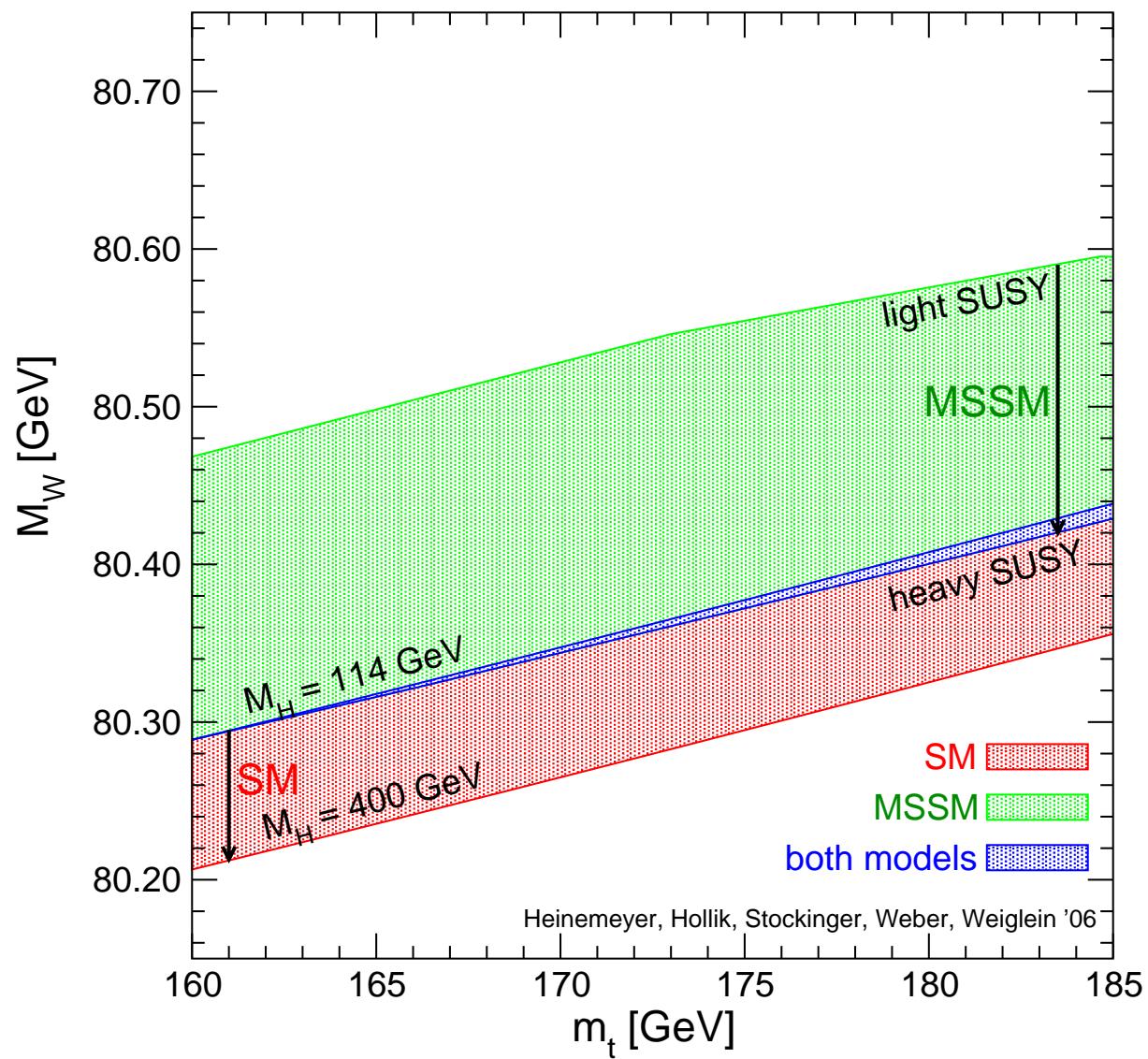
Test of theory at quantum level: **Sensitivity to loop corrections**



Very high accuracy of measurements and theoretical predictions needed

- Which model fits better?
- Does the prediction of a model contradict the experimental data?

Prediction for  $M_W$  in the SM and the MSSM :  
[S.H., W. Hollik, D. Stockinger, A.M. Weber, G. Weiglein '06]



MSSM band:

scan over  
SUSY masses

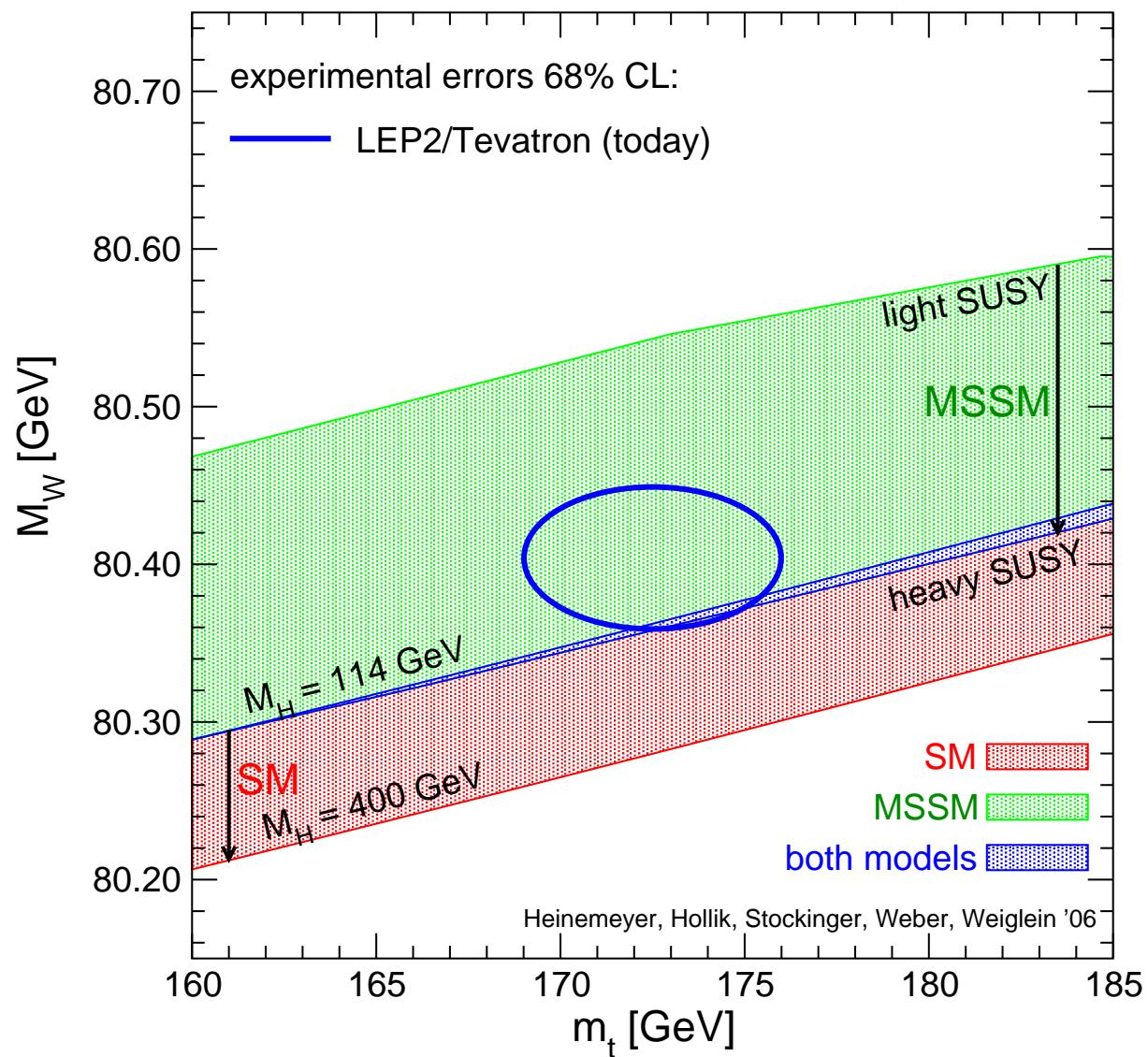
overlap:

SM is MSSM-like  
MSSM is SM-like

SM band:

variation of  $M_H^{\text{SM}}$

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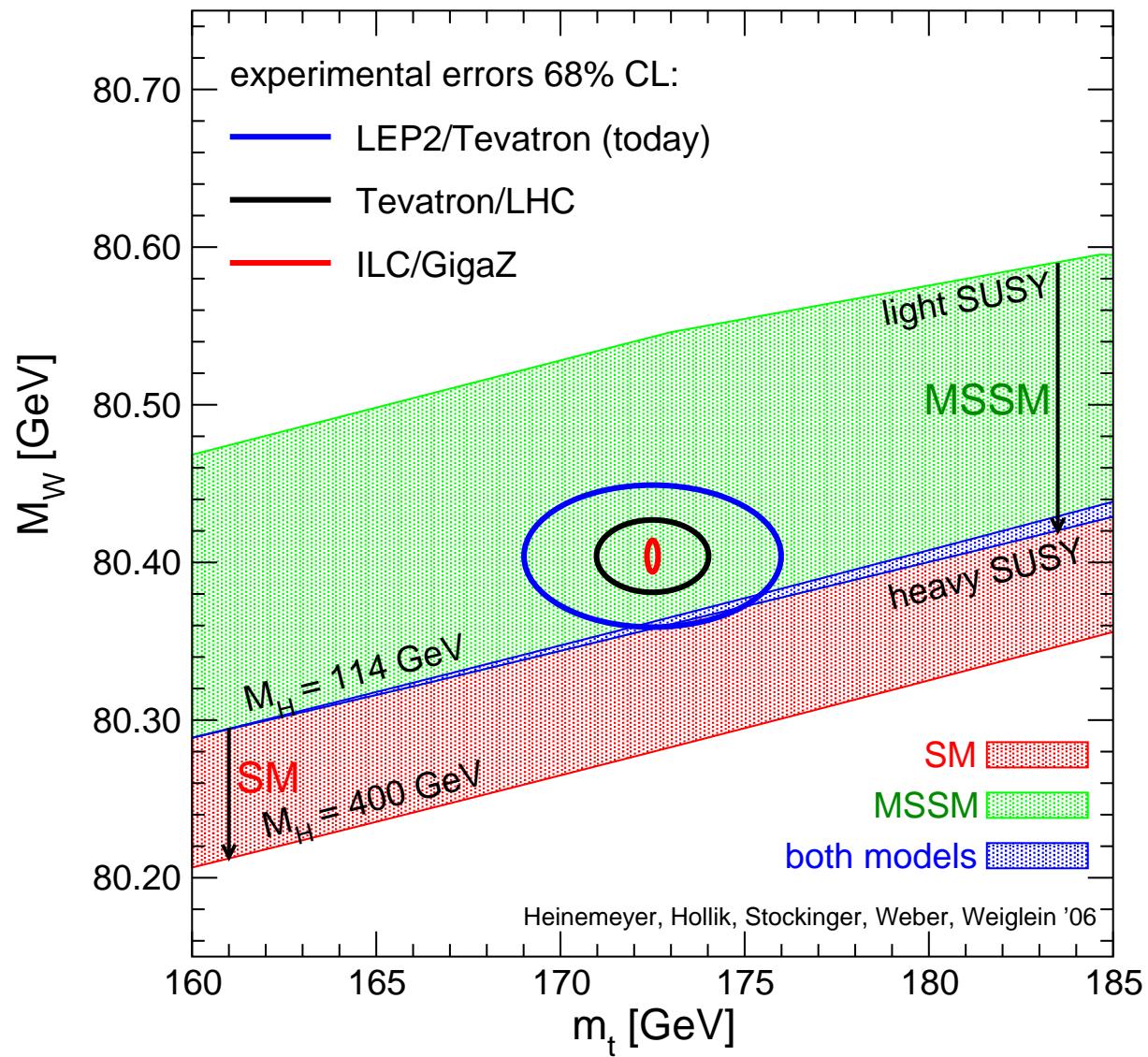
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variation of  $M_H^{\text{SM}}$

# Prediction for $M_W$ in the SM and the MSSM :

[S.H., W. Hollik, D. Stockinger, A.M. Weber, G. Weiglein '06]



MSSM band:

scan over  
SUSY masses

overlap:

SM is MSSM-like  
MSSM is SM-like

SM band:

variation of  $M_H^{\text{SM}}$

## 2. Calculation of $M_W$ in the MSSM

Theoretical prediction for  $M_W$  in terms of  $M_Z, \alpha, G_\mu, \Delta r$ :

$$M_W^2 \left( 1 - \frac{M_W^2}{M_Z^2} \right) = \frac{\pi \alpha}{\sqrt{2} G_\mu} (1 + \Delta r)$$

⇓  
loop corrections

Evaluate  $\Delta r$  from  $\mu$  decay  $\Rightarrow M_W$

One-loop result for  $M_W$  in the SM:

[A. Sirlin '80] , [W. Marciano, A. Sirlin '80]

$$\begin{aligned} \Delta r_{\text{1-loop}} &= \Delta \alpha - \frac{c_W^2}{s_W^2} \Delta \rho + \Delta r_{\text{rem}}(M_H) \\ &\sim \log \frac{M_Z}{m_f} \quad \sim m_t^2 \\ &\sim 6\% \quad \sim 3.3\% \quad \sim 1\% \end{aligned}$$

## Combination of SM and MSSM result

- use best available SM result
- add all available MSSM corrections
- subtract double counting

⇒ decoupling limit ok ( $M_{\text{SUSY}} \rightarrow \infty$ )

... but some corrections included only via their SM part

⇒ best solution

## Status of SM calculation

- Most recently full electroweak two-loop result available  
*[Awramik, Czakon, Freitas, Hollik, Onishenko, Veretin, Walter, Weiglein '03, '04]*
- Most recent electroweak three-loop corrections via  $\Delta\rho$   
*[Faisst, Kühn, Seidensticker, Veretin '04]*
- Compact parametric formula used  
*[Awramik, Czakon, Freitas, Weiglein '04]*

## Status of MSSM calculations of $M_W$

- MSSM,  $\Delta r$ : full one-loop corrections  
[P. Chankowski, A. Dabelstein, W. Hollik, W. Mösle, S. Pokorski, J. Rosiek '94]  
[D. Garcia, J. Solà '94]
- MSSM,  $\Delta \rho$ : leading  $\mathcal{O}(\alpha \alpha_s)$  corrections  
[A. Djouadi, P. Gambino, S.H., W. Hollik, C. Jünger, G. Weiglein '97]
- MSSM,  $\Delta r$ : leading gluonic  $\mathcal{O}(\alpha \alpha_s)$  corr.  
[S.H. '98] [S.H., W. Hollik, G. Weiglein '04]
- MSSM,  $\Delta \rho$ : leading  $\mathcal{O}(\alpha_t^2, \alpha_t \alpha_b, \alpha_b^2)$  corrections  
[S.H., G. Weiglein '01, '03]  
[J. Haestier, S.H., D. Stöckinger, G. Weiglein '05]

### Most recent:

- MSSM,  $\Delta r$ : one-loop with complex parameters  
[S.H., W. Hollik, D. Stöckinger, A.M. Weber, G. Weiglein '06]

### Most dominant:

- $\Delta \rho$ : one- or two-loop, contribution from  $\tilde{t}/\tilde{b}$  sector

## More details about our calculation:

- Complex phases in the squark sector enter only via shift in squark masses (explicit dependence drops out)

$$\begin{aligned}|X_t|^2 &= |A_t|^2 + |\mu \cot \beta|^2 - 2|A_t| \cdot |\mu| \cot \beta \cos(\phi_{A_t} + \phi_\mu) \\ |X_b|^2 &= |A_b|^2 + |\mu \tan \beta|^2 - 2|A_b| \cdot |\mu| \tan \beta \cos(\phi_{A_b} + \phi_\mu)\end{aligned}$$

Only some phase combinations are physical,  
other phases can be rotated away.

Examples for physical combinations:

$$\begin{aligned}\phi_{A_t} + \phi_\mu \\ \phi_{A_b} + \phi_\mu\end{aligned}$$

- Higgs mass dependence of the two-loop contributions is known to be very strong  
⇒ we use *FeynHiggs* ([www.feynhiggs.de](http://www.feynhiggs.de))
- All one-loop calculations have been performed with *FeynArts* and *FormCalc*  
[T. Hahn et al '00 - '05]

## Treatment of the phase dependence beyond one-loop order

Phase dependence at the two-loop level approximated by a simple interpolation based on:

full phase dependence at the one-loop level,  $M_W^{1L}(\phi)$ ,

two-loop results for real parameters,  $M_W^{\text{full}}(0)$ ,  $M_W^{\text{full}}(\pi)$

⇒ Two-loop result for complex phase  $\phi$ :

$$\begin{aligned} M_W^{\text{full}}(\phi) = & M_W^{1L}(\phi) + [M_W^{\text{full}}(0) - M_W^{1L}(0)] \times \frac{1 + \cos \phi}{2} \\ & + [M_W^{\text{full}}(\pi) - M_W^{1L}(\pi)] \times \frac{1 - \cos \phi}{2} \end{aligned}$$

### 3. Numerical results

#### 3A) phase dependence from squark sector

Complex phases in the squark sector enter only via shift in squark masses (explicit dependence drops out)  
⇒ phase dependence must be reflected in the squark masses

Only some phase combinations are physical,  
other phases can be rotated away.

Examples for physical combinations:

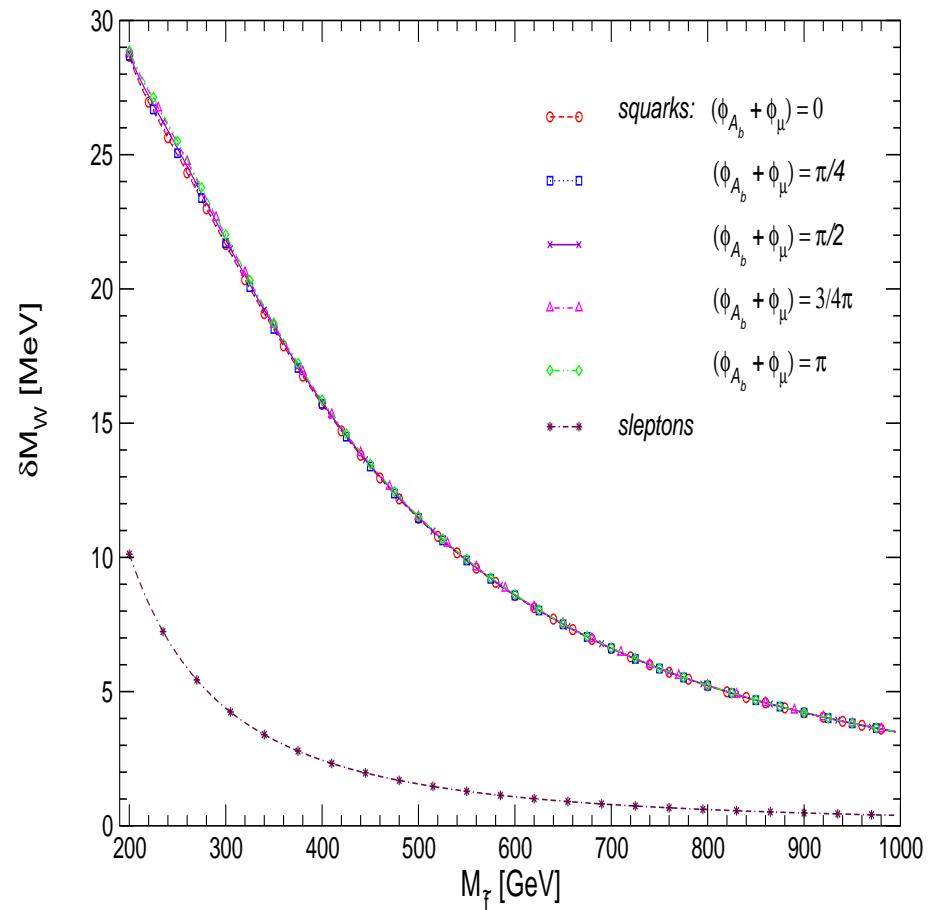
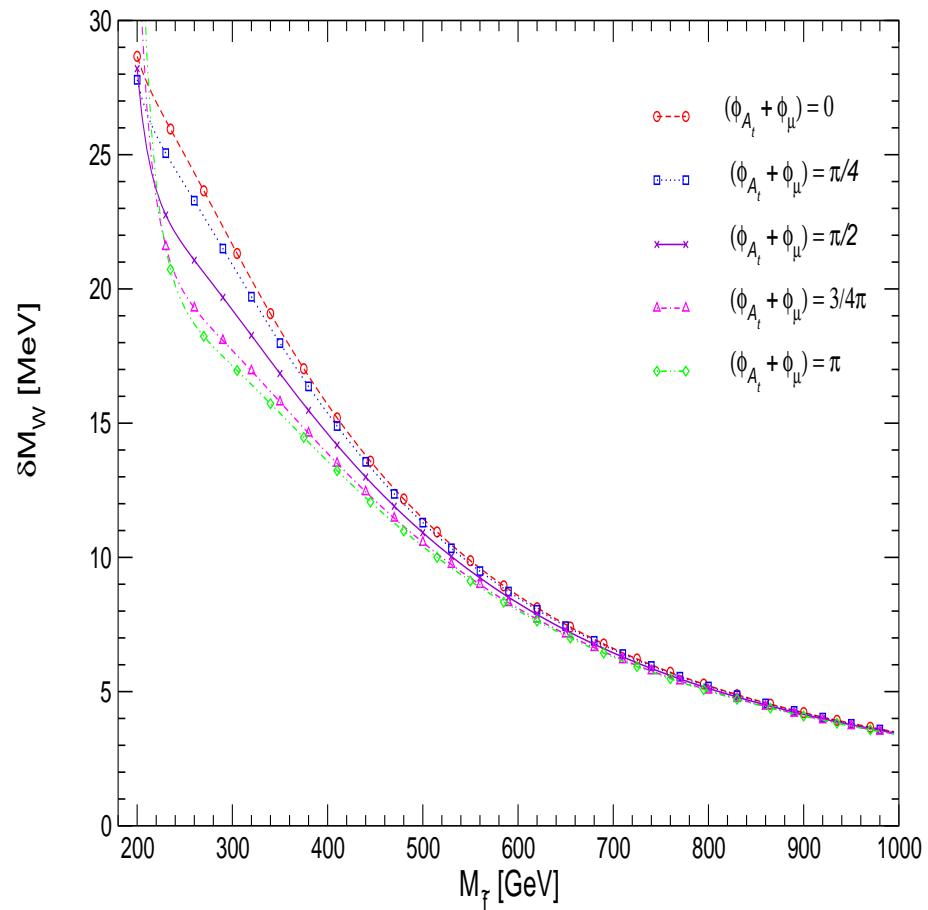
$$\begin{aligned}\phi_{A_t} + \phi_\mu \\ \phi_{A_b} + \phi_\mu\end{aligned}$$

$\delta M_W$  evaluated from

$$\delta M_W = -\frac{M_W}{2} \frac{s_W^2}{c_W^2 - s_W^2} \Delta r$$

## $\delta M_W$ dependence on $\phi_{A_t}$ and $\phi_{A_b}$ (I):

( $|A_{t,b}| = 350$  GeV,  $\mu = 300$  GeV,  $\tan \beta = 10$ )

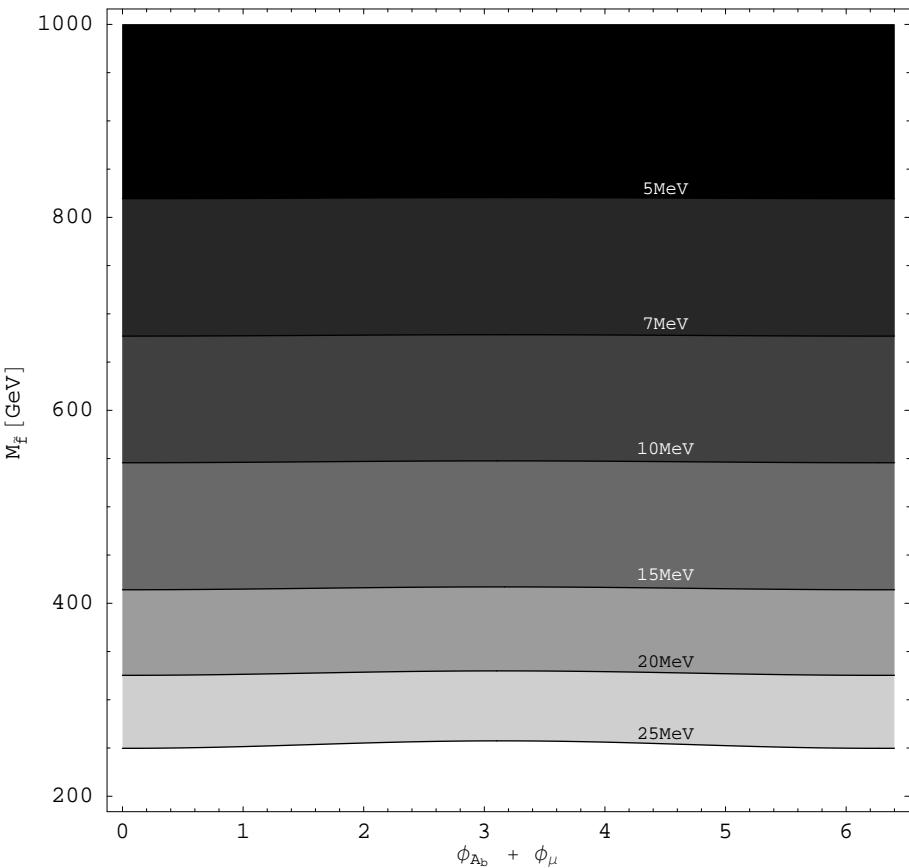
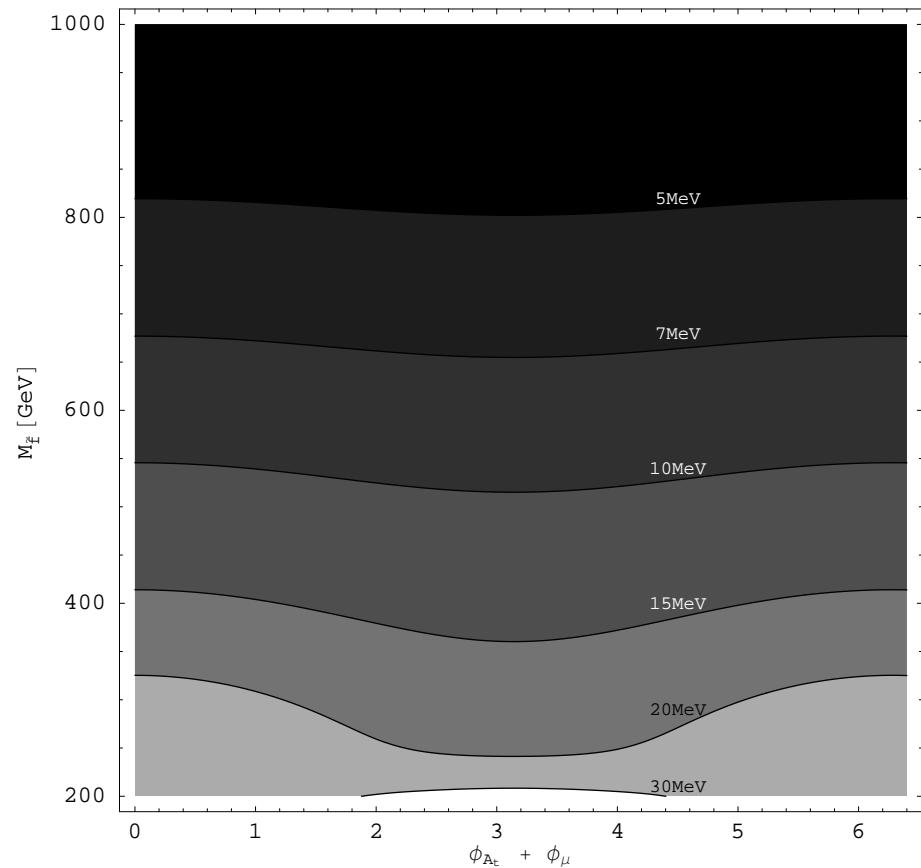


→ large squark contribution, decoupling with  $M_{\text{SUSY}}$

→  $\phi_{A_t}$  dependence, but no  $\phi_{A_b}$  effects

## $\delta M_W$ dependence on $\phi_{A_t}$ and $\phi_{A_b}$ (II):

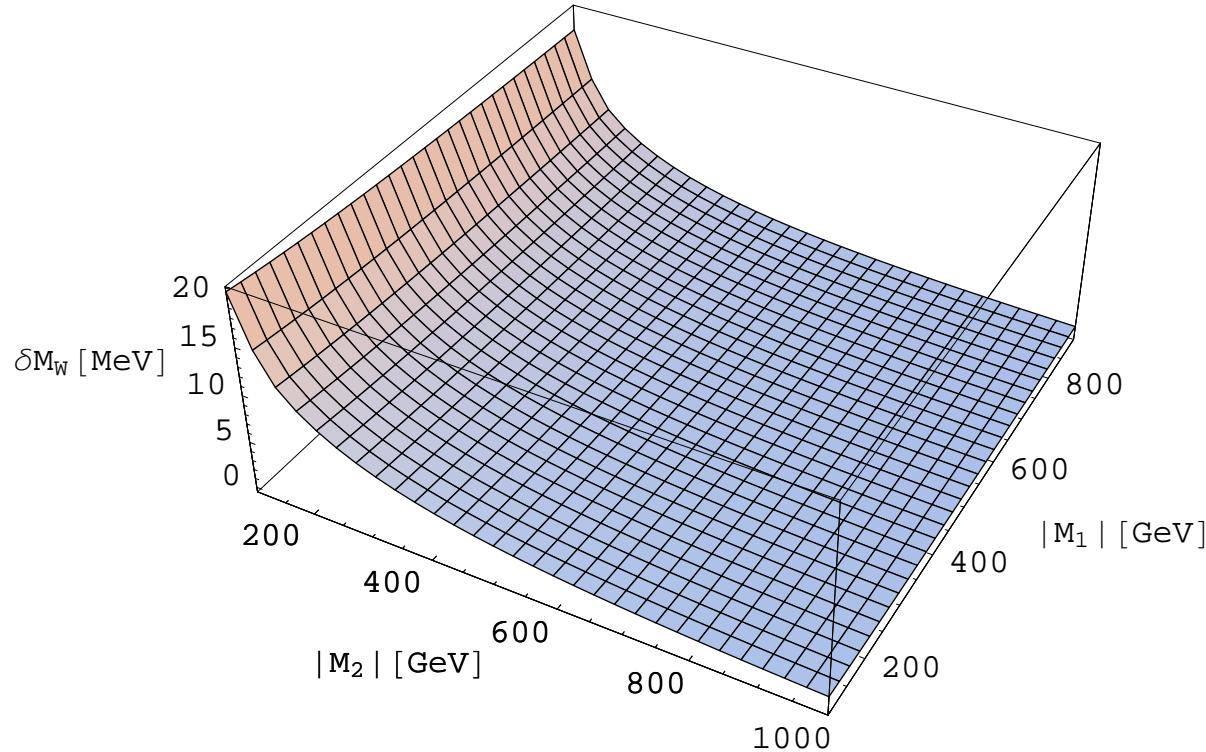
( $|A_{t,b}| = 350$  GeV,  $\mu = 300$  GeV,  $\tan \beta = 10$ )



→ large squark contribution, decoupling with  $M_{\text{SUSY}}$

→  $\phi_{A_t}$  dependence, but no  $\phi_{A_b}$  effects

### 3B) dependence on chargino/neutralino sector

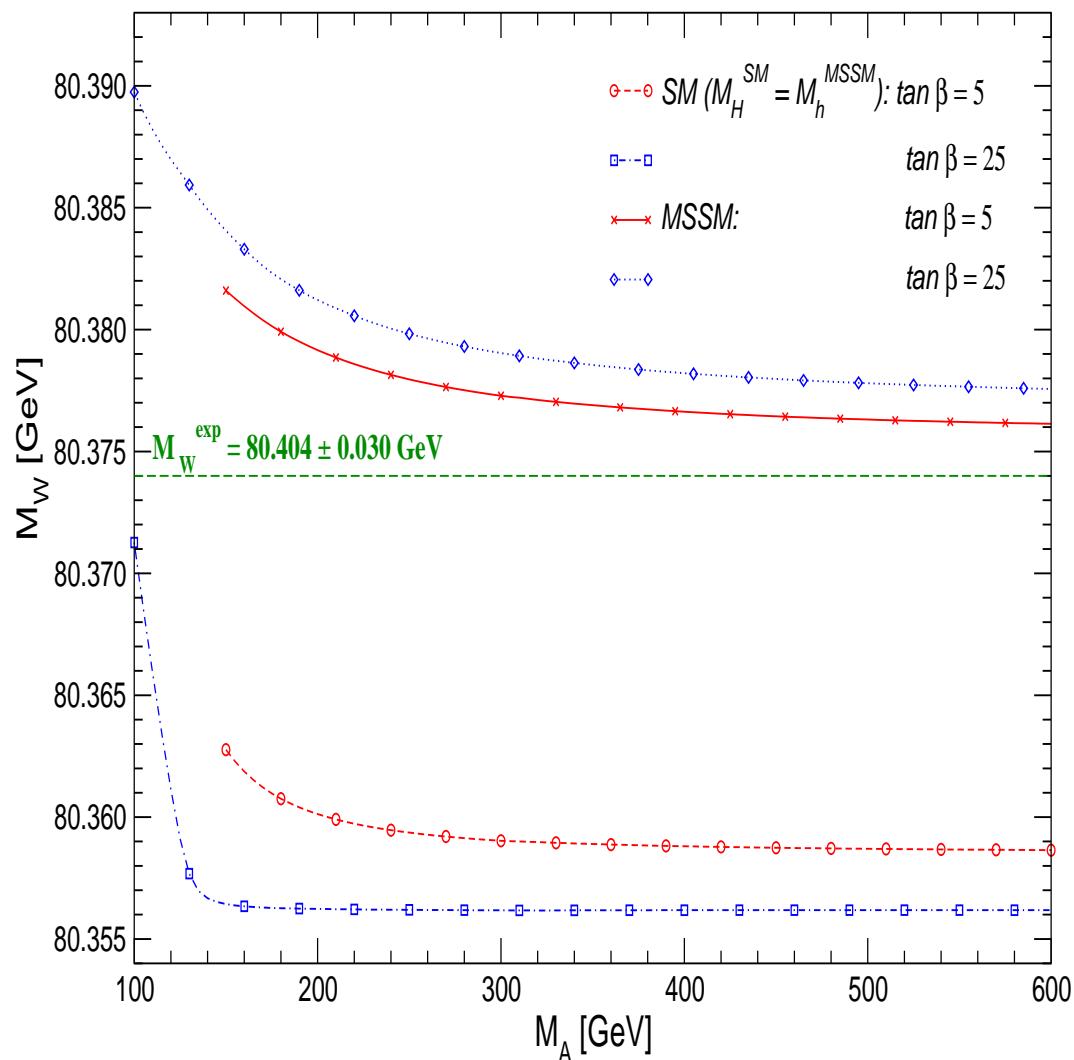


- hardly any  $M_1$  dependence
- up to **20 MeV** contribution via  $M_2$

other parameters:  
 $M_{\text{SUSY}} = 250 \text{ GeV}$   
 $\mu = 300 \text{ GeV}$ ,  
 $\phi_{M_{1,2}} = 0$   
 $\tan \beta = 10$

### 3C) Prediction of $M_W$

Comparison with SM:



other parameters:

$$M_{\text{SUSY}} = 600 \text{ GeV}$$

$$A_{t,b} = 1200 \text{ GeV}$$

$$\mu = M_2 = m_{\tilde{g}} = 300 \text{ GeV}, \\ \phi_x = 0$$

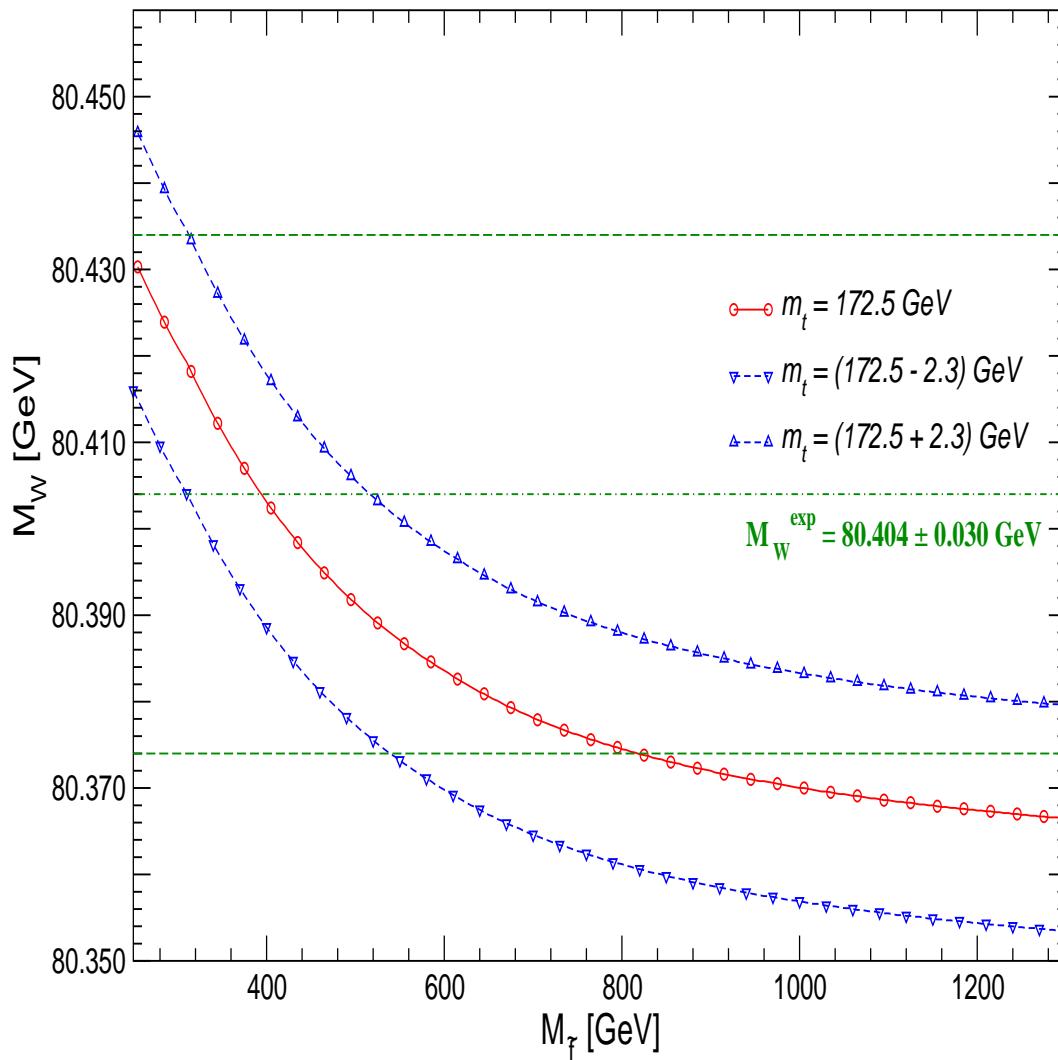
$$M_W^{\text{SUSY}} - M_W^{\text{SM}} \gtrsim 20 \text{ MeV} \\ (\text{typical difference})$$

Comparison with  $M_W^{\text{exp.}}$ :

MSSM agrees at  $1\sigma$

SM agrees at  $2\sigma$

## Prediction of $M_W$ : variation of $m_t$



other parameters:

$$A_{t,b} = 2 M_{\text{SUSY}}$$

$$\mu = M_2 = m_{\tilde{g}} = 300 \text{ GeV},$$

$$\phi_x = 0$$

$$\tan \beta = 10$$

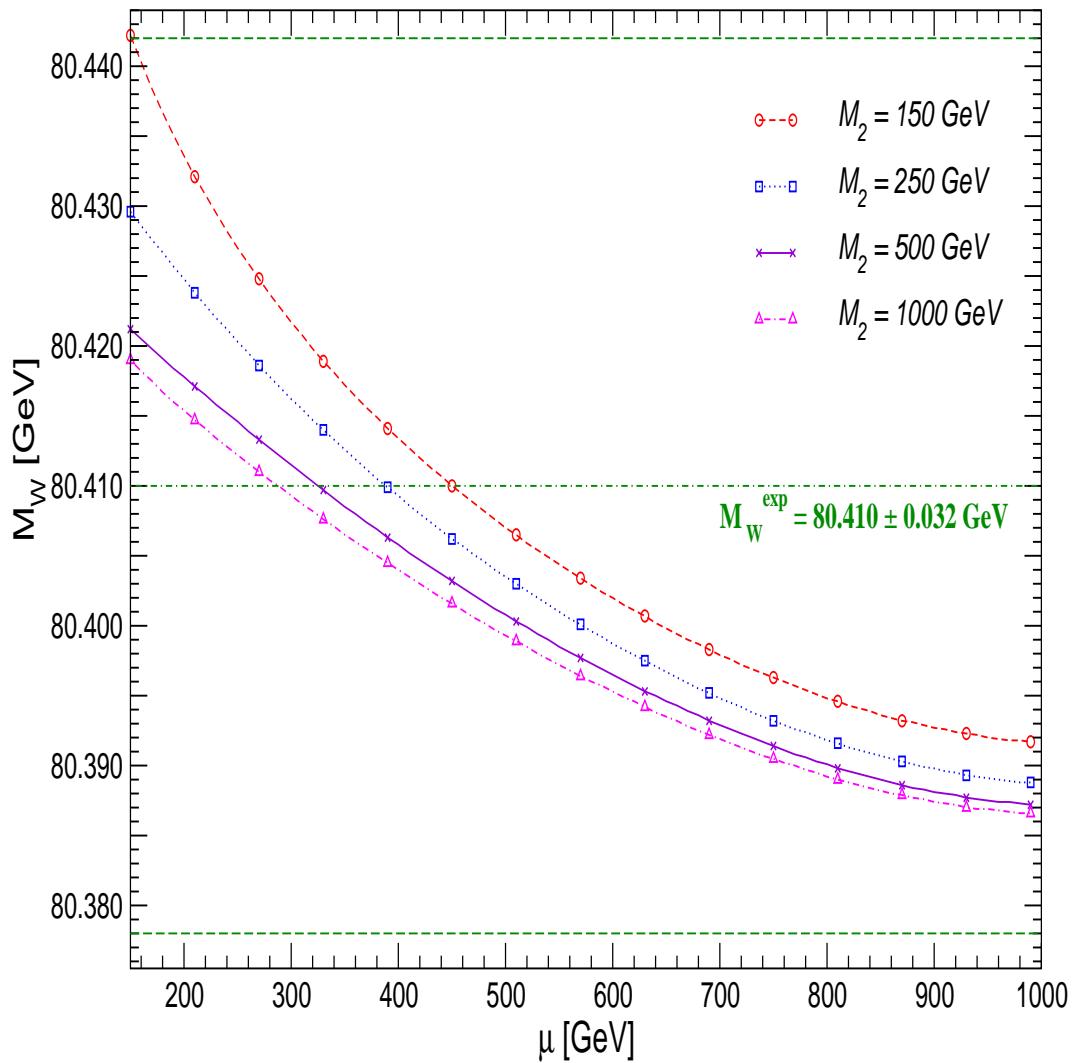
$$\delta m_t = \pm 2.3 \text{ GeV}$$

$$\Rightarrow \delta M_W \approx \pm 14 \text{ MeV}$$

lower  $m_t$  disfavored

$\Rightarrow$  largest parametric uncertainty

## Prediction of $M_W$ : variation of $\mu$ and $M_2$



other parameters:

$$M_{\text{SUSY}} = 300 \text{ GeV}$$

$$A_{t,b} = 2 M_{\text{SUSY}}$$

$$m_{\tilde{g}} = 300 \text{ GeV},$$

$$\phi_x = 0$$

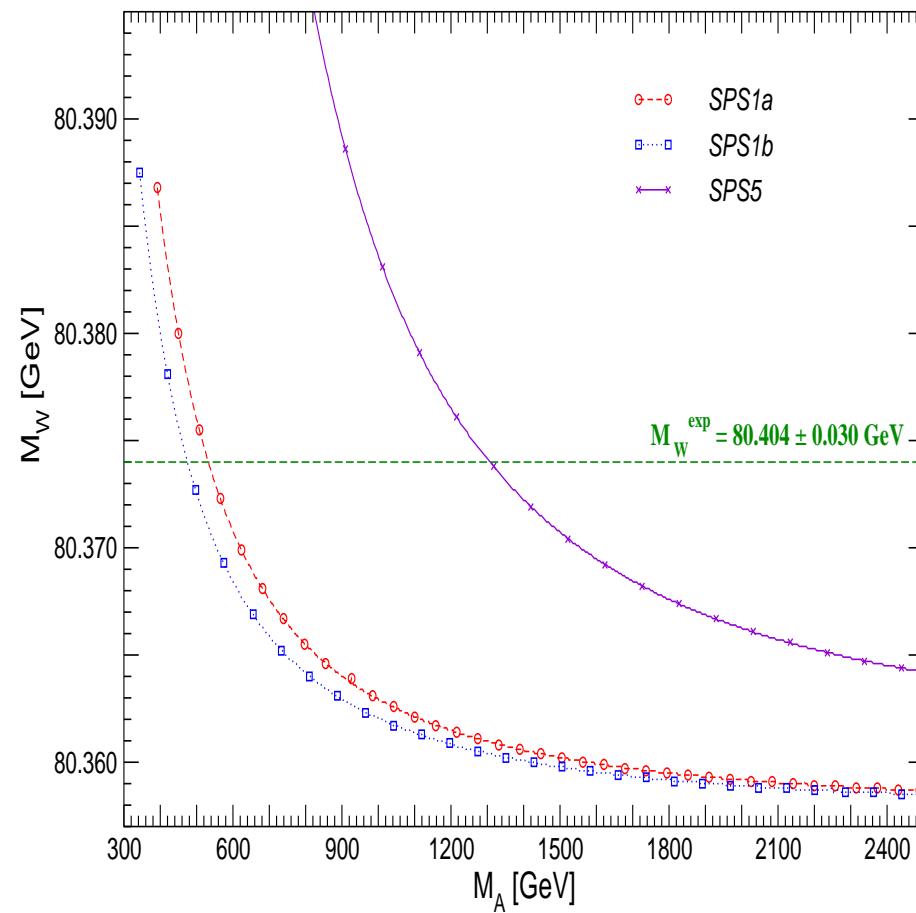
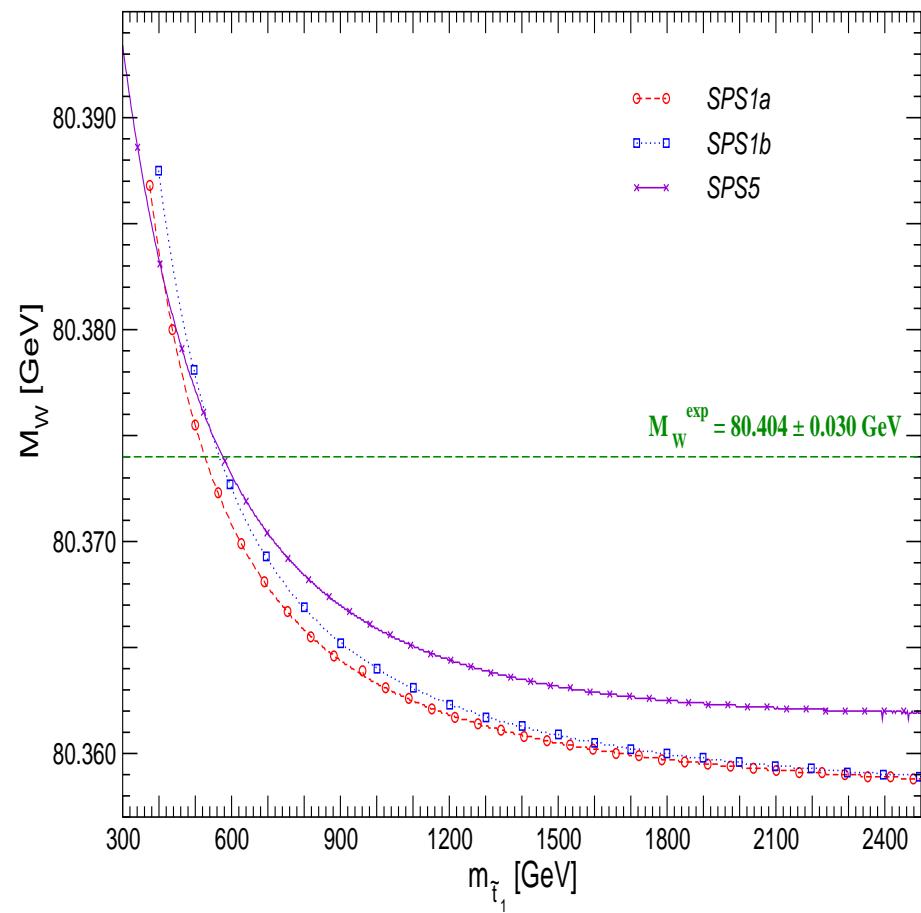
$$\tan \beta = 10$$

$$M_A = 1000 \text{ GeV}$$

⇒ large variation with  $\mu$

⇒ smaller variation with  $M_2$

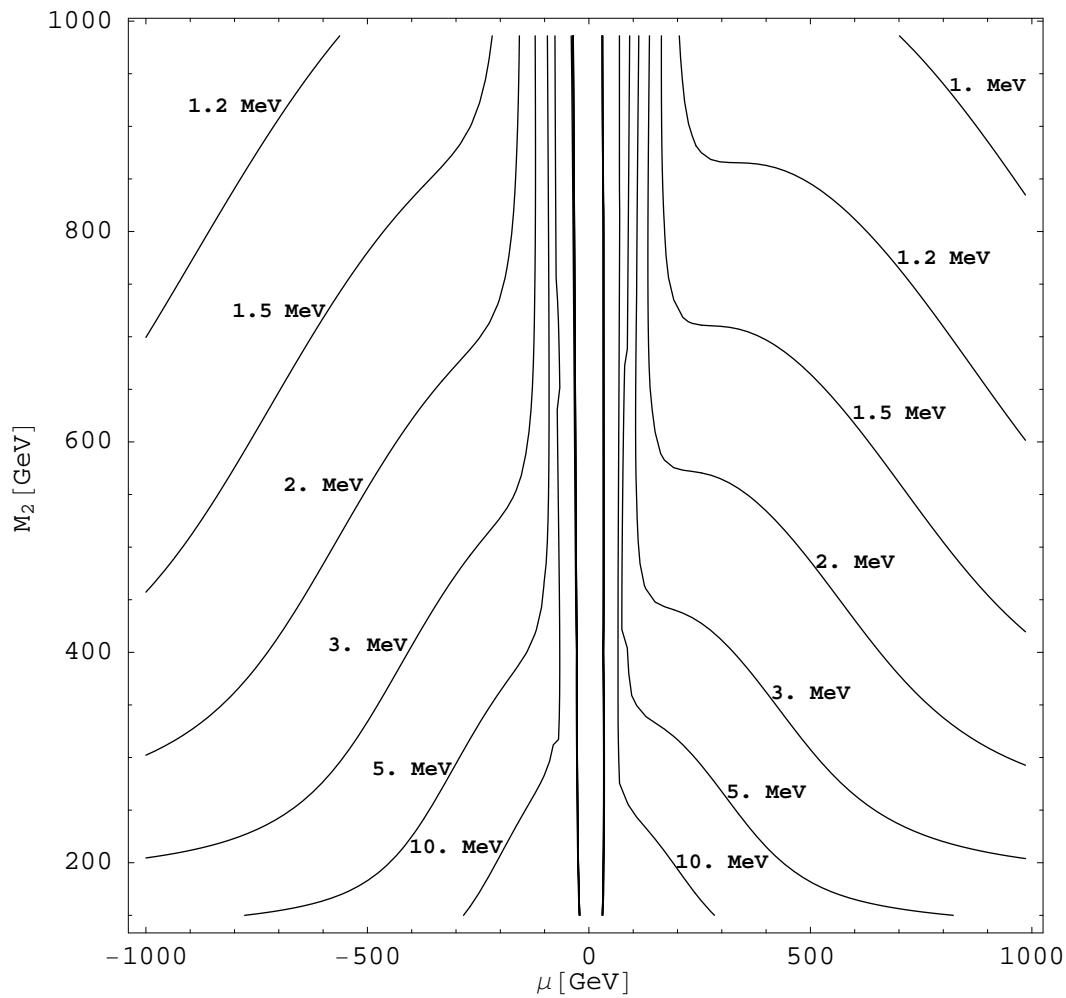
## Prediction of $M_W$ : typical scenarios: SPS 1a, 1b, 5



⇒ only “lower” masses show agreement at  $1\sigma$

## Prediction of $M_W$ : Split SUSY

[N. Arkani-Hamed, S. Dimopoulos '04] [G. Giudice, A. Romanino '04]



Difference to  $M_W^{\text{SM}}$   
with  $M_H^{\text{SM}} = M_h$

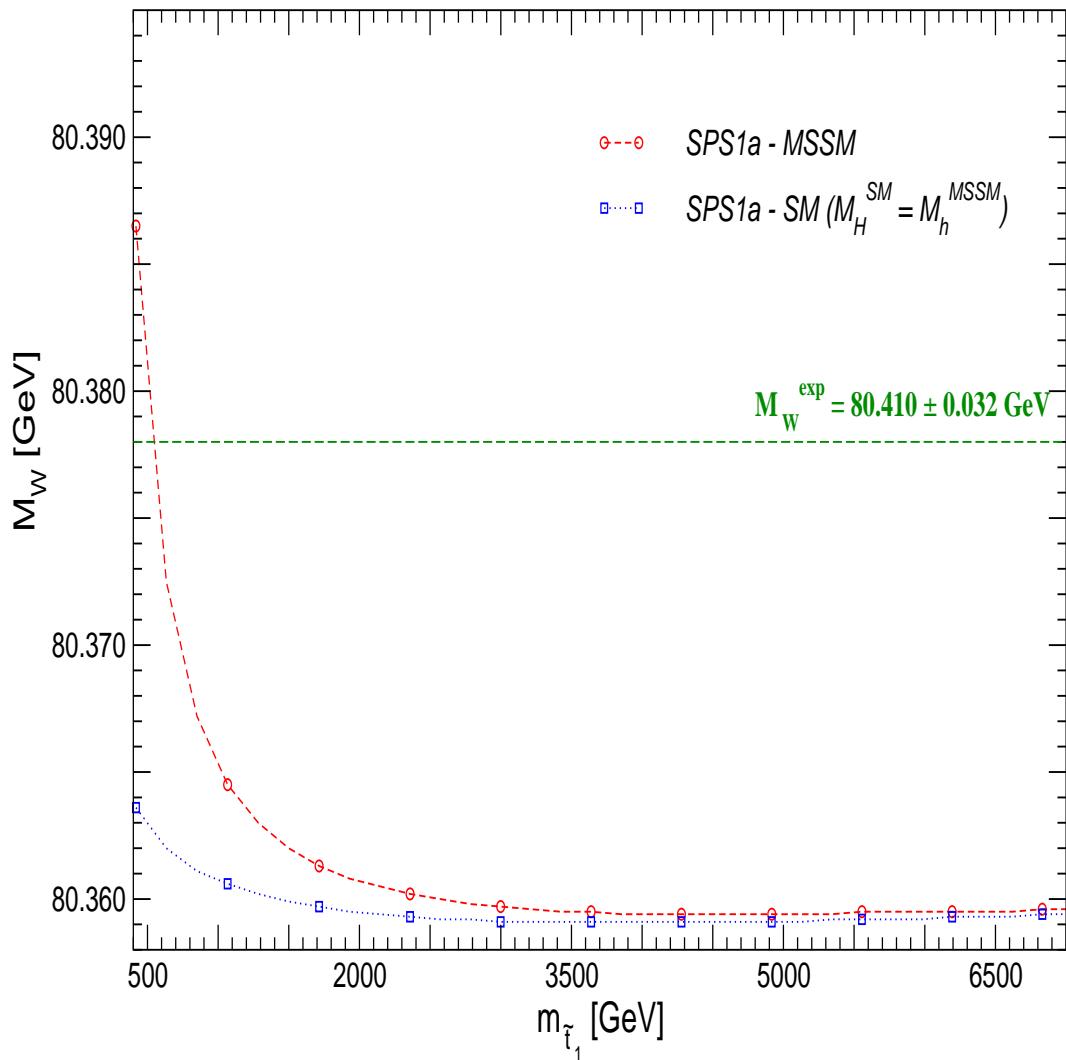
⇒ no deviation for  
Tevatron, LHC, ILC  
precision

⇒ GigaZ can confirm only  
very light masses

(as expected . . .)

Split SUSY disagrees with  
experiment at the  $2\sigma$  level

## Prediction of $M_W$ : Decoupling limit



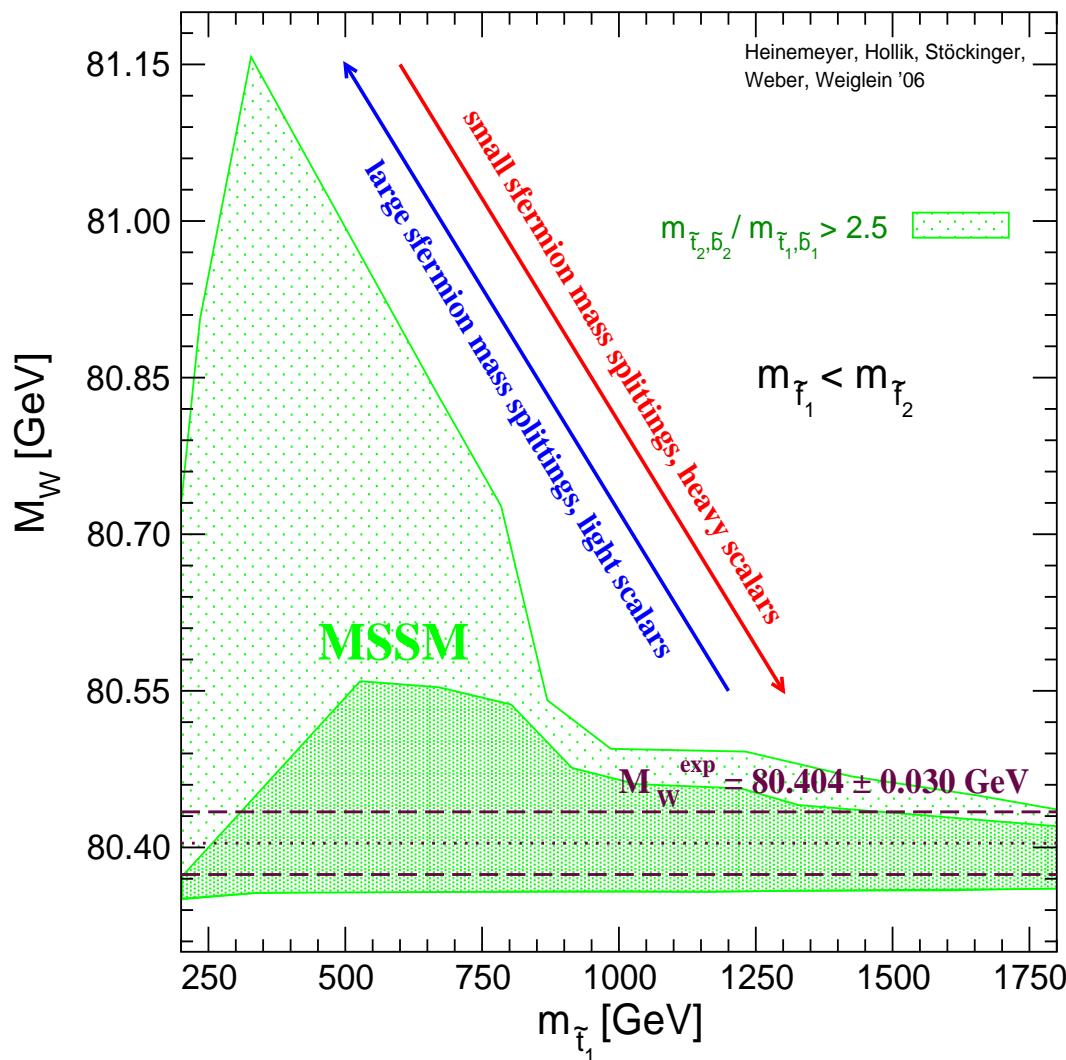
Compared to SM with  
 $M_H^{\text{SM}} = M_h$

$M_{\text{SUSY}} = 500 \text{ GeV}$   
 $\Rightarrow \delta M_W^{\text{MSSM-SM}} > 20 \text{ MeV}$

$M_{\text{SUSY}} = 1 \text{ TeV}$   
 $\Rightarrow \delta M_W^{\text{MSSM-SM}} < 10 \text{ MeV}$

$M_{\text{SUSY}} > 3 \text{ TeV}$   
 $\Rightarrow \delta M_W^{\text{MSSM-SM}} < 1 \text{ MeV}$

## Prediction of $M_W$ : Parameter scan



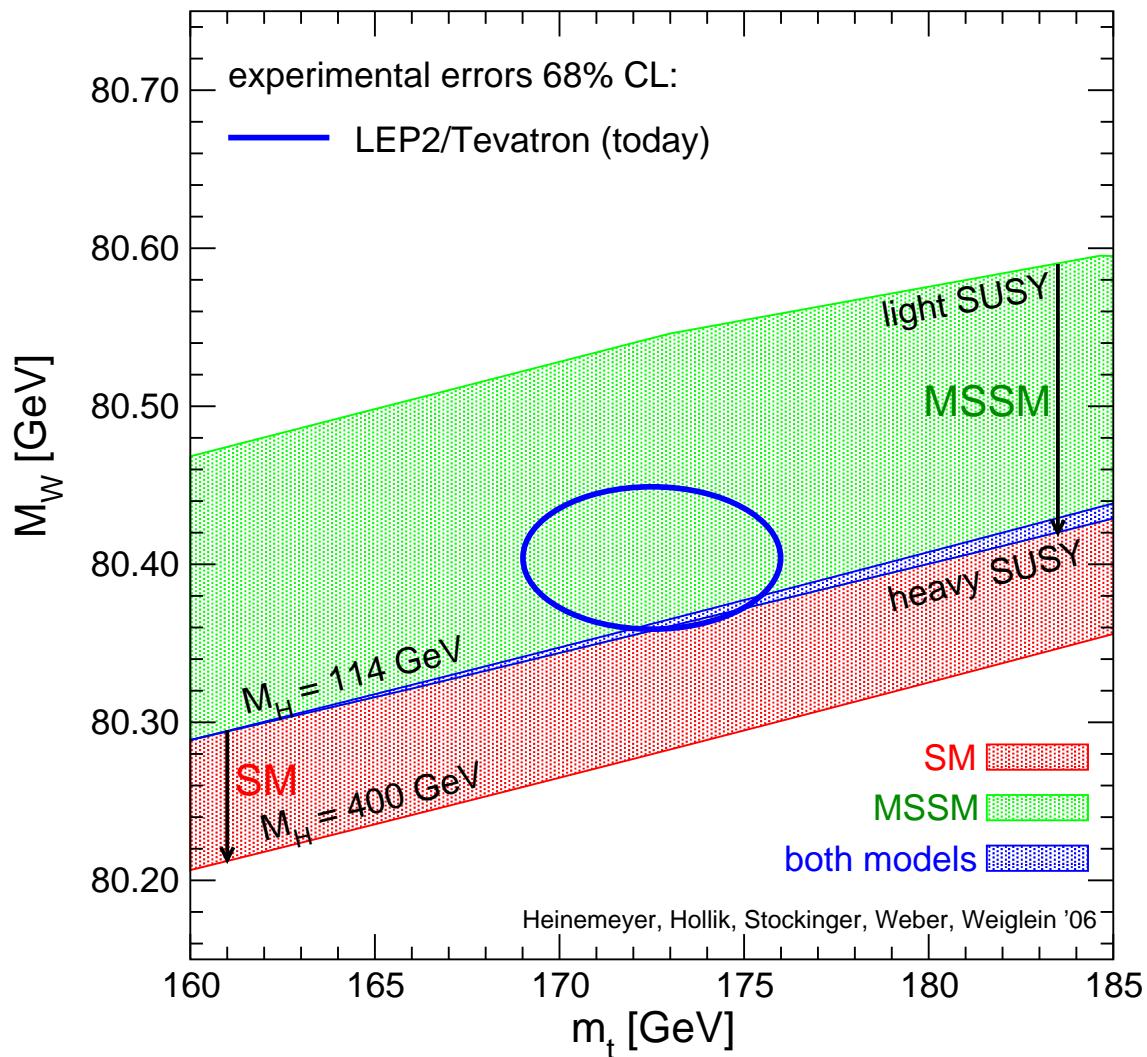
Scatter **all** relevant parameters **independently** over **wide** ranges:  
masses  $\lesssim 2$  TeV  
 $\tan \beta = 1.1 \dots 60$

Experimental constraints from LEP, Tevatron included

huge splitting  
 $\Rightarrow$  huge correction to  $M_W$   
(experimentally excluded)

## Impact of $M_W$ on preferred SUSY masses?

Prediction for  $M_W$  in the **SM** and the **MSSM**:



**MSSM band:**

scan over  
SUSY masses

**overlap:**

SM is MSSM-like  
MSSM is SM-like

**SM band:**

variation of  $M_H^{\text{SM}}$

⇒  $M_W$  is not enough

⇒ use more precision obs.

→ talk by S.H. later today :-)

## 4. Remaining theoretical (intrinsic) uncertainties

[J. Haestier, S.H., D. Stöckinger, G. Weiglein '05]

[S.H., W. Hollik, D. Stöckinger, A.M. Weber, G. Weiglein '06]

Estimate missing SUSY corrections order by order:

- $\mathcal{O}(\alpha_t^2, \alpha_t \alpha_b, \alpha_b^2)$ : beyond existing leading contributions
- $\mathcal{O}(\alpha \alpha_s)$ : beyond  $\Delta\rho$  approx.
- $\mathcal{O}(\alpha \alpha_s^2)$
- $\mathcal{O}(\alpha^2 \alpha_s)$
- $\mathcal{O}(\alpha^3)$
- missing phase dependence at two-loop

⇒ evaluate for  $M_{\text{SUSY}} = 300, 500, 1000 \text{ GeV}$

Combine with SM uncertainty:  $\delta M_W^{\text{SM,intr.}} = 4 \text{ MeV}$

$$\delta M_W^{\text{SUSY,intr.}} = 5 - 11 \text{ MeV}$$

(depending on  $M_{\text{SUSY}}$ )

## 5. Conclusions

- Precision observables
  - can give valuable information about the “true” Lagrangian
  - can provide bounds on SUSY parameter space  
⇒  $M_W$  prominent example
- Combination of
  - full SM result
  - all available MSSM corr. ⇒ best prediction of  $M_W$  in the MSSM  
⇒ remaining uncertainty:  $\delta M_W^{\text{intr.}} = 5 - 11 \text{ MeV}$  (depending on  $M_{\text{SUSY}}$ )
- Effects of certain sectors:
  - $\tilde{t}/\tilde{b}$  sector:  $\delta M_W \gtrsim 20 \text{ MeV}$ , slepton sector: much smaller
  - complex phases in  $\tilde{t}/\tilde{b}$  sector:  $\delta M_W \gtrsim 5 \text{ MeV}$
  - chargino/neutralino sector:  $\delta M_W \gtrsim 20 \text{ MeV}$
  - $\delta m_t = \pm 2.3 \text{ GeV} \Rightarrow \delta M_W \approx \pm 14 \text{ MeV} \Rightarrow$  largest parametric unc.
- Prediction of  $M_W$  in special scenarios:
  - SPS: only “lower” masses show agreement at  $1\sigma$
  - Split SUSY: basically no visible effect
  - Decoupling limit:  $M_{\text{SUSY}} > 3 \text{ TeV} \Rightarrow \delta M_W^{\text{MSSM-SM}} < 1 \text{ MeV}$
  - Parameter scan: very large  $M_W$  only for very large splitting in  $\tilde{t}/\tilde{b}$