

# Preliminary hadron analysis for the CALICE AHCAL

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**I. Event selection impact on the energy resolution and linearity**

**II. New method of software compensation for HCAL**



## Data and software

**DATA:** CERN 2007 test beam runs with complete CALICE setup

$\pi^-$  10, 12, 15, 18, 35, 80 GeV

$\pi^+$  30, 40, 50, 60, 80 GeV

the newest official reconstruction software as of April 2010

**MC:** QGSP BERT and LHEP physics lists

$\pi^-$  10 GeV

$\pi^+$  30, 50, 80 GeV (thanks to Lars Weuste)

official Mokka and digitization software as of April 2010

## Event selection stages

*Before any analysis:*

**Apply 0.5-MIP cut to both ECAL and HCAL hits**



**Identify muons using 2D histogram  $E_{ECAL}^{dep} + E_{HCAL}^{dep}$  vs.  $E_{TCMT}^{dep}$  (Vasily's method)**



**Identify trash and multiparticle events**



**Separate electrons from  $\pi^-$  (Čerenkov counter)  
Separate protons from  $\pi^+$  (Čerenkov counter)**

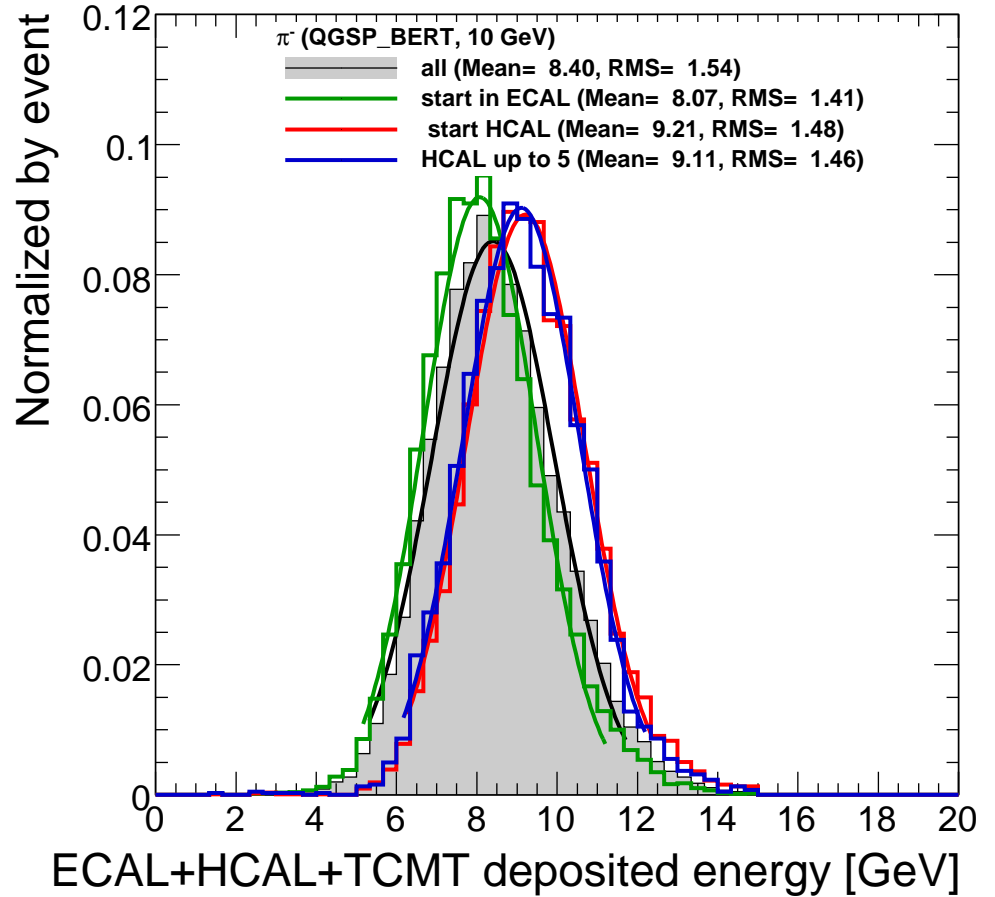
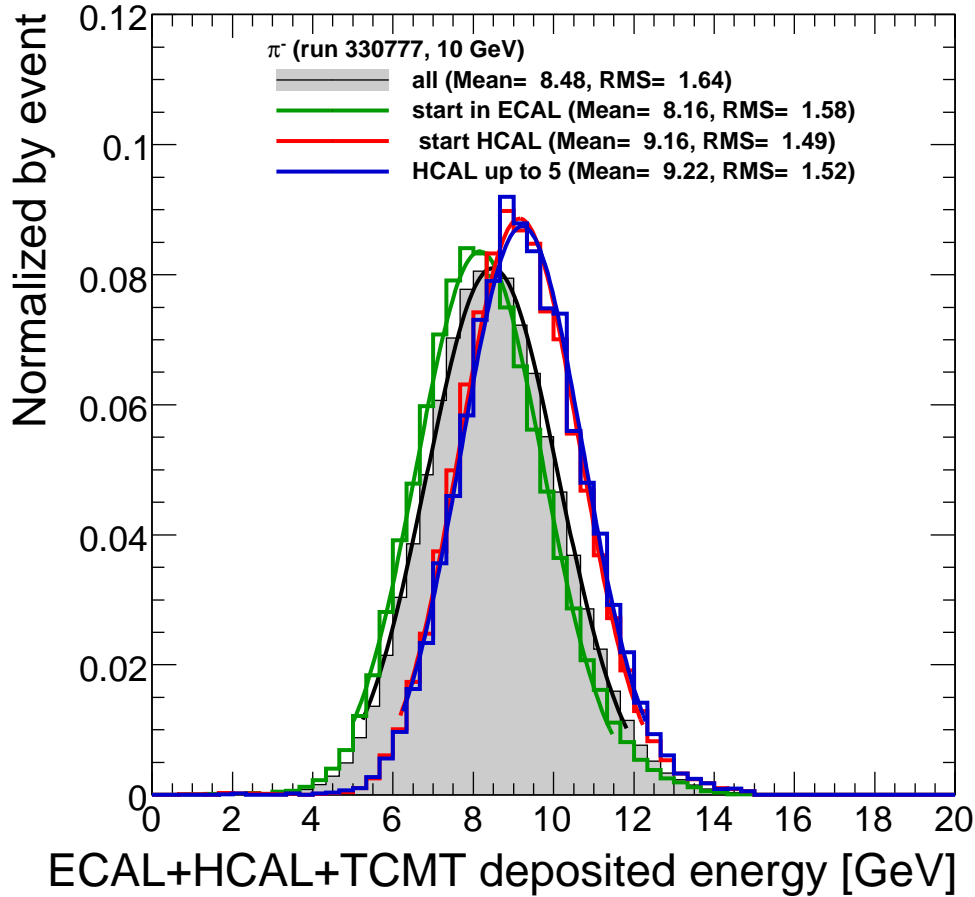
*For analysis:*

**Find shower start position (using PrimaryTrackFinder)**



**Select events by shower start position: in ECAL, in HCAL, in first 5 HCAL layers**

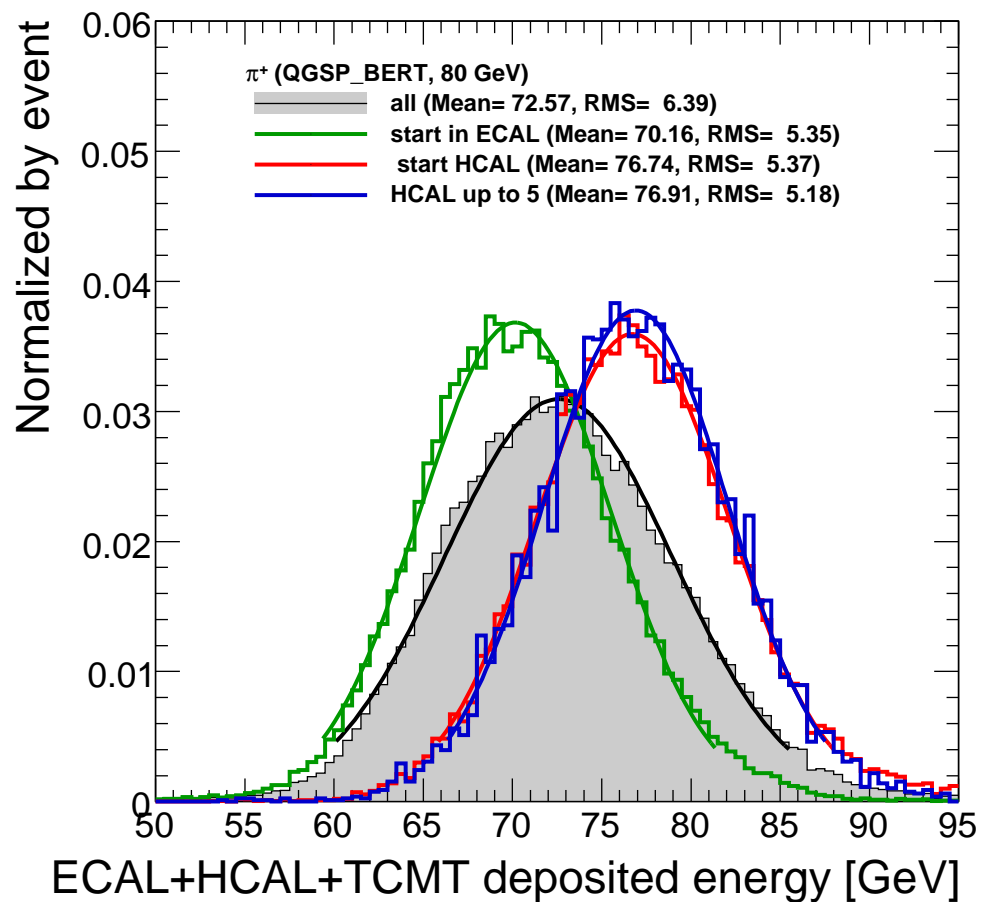
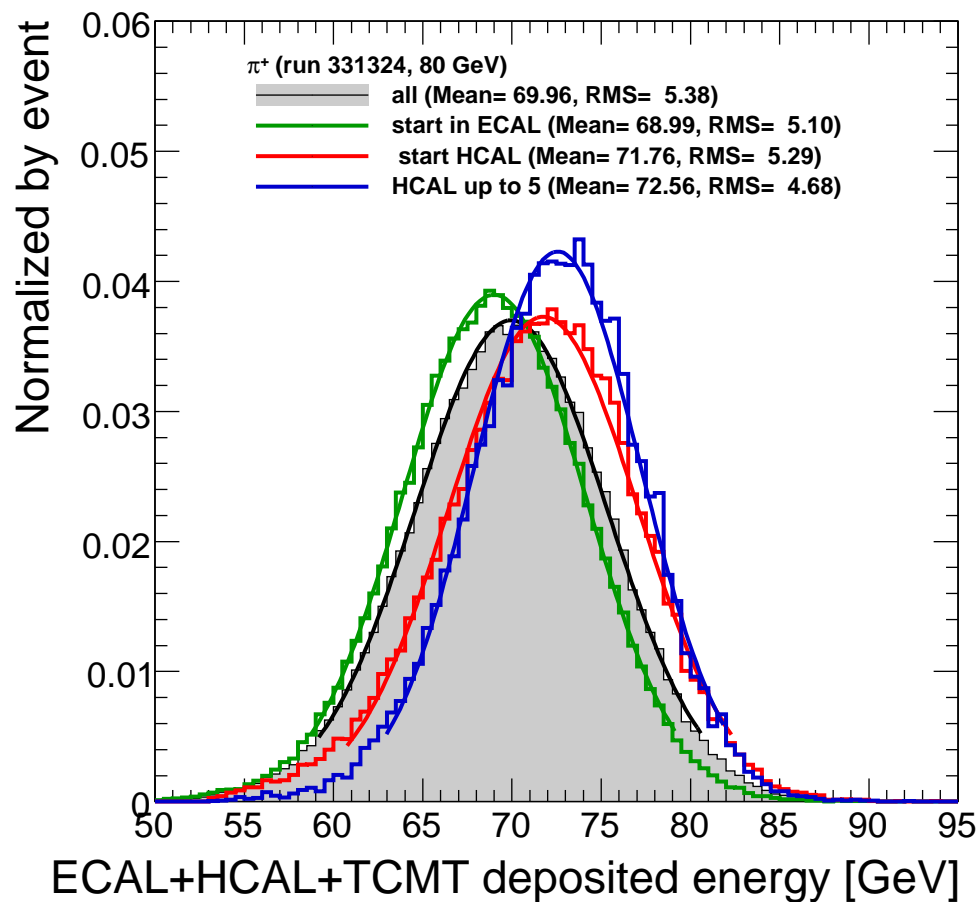
# Event selection: energy distributions for 10-GeV $\pi^-$



**ECAL:**  $\left(\frac{\text{GeV}}{\text{MIP}}\right)_{vis}^{ecal} = 0.000147$ ;  $S_{ecal} = 25.57$ ;  $\text{ECAL}_1:\text{ECAL}_2:\text{ECAL}_3 = 1:2:3$  (from CAN-008)

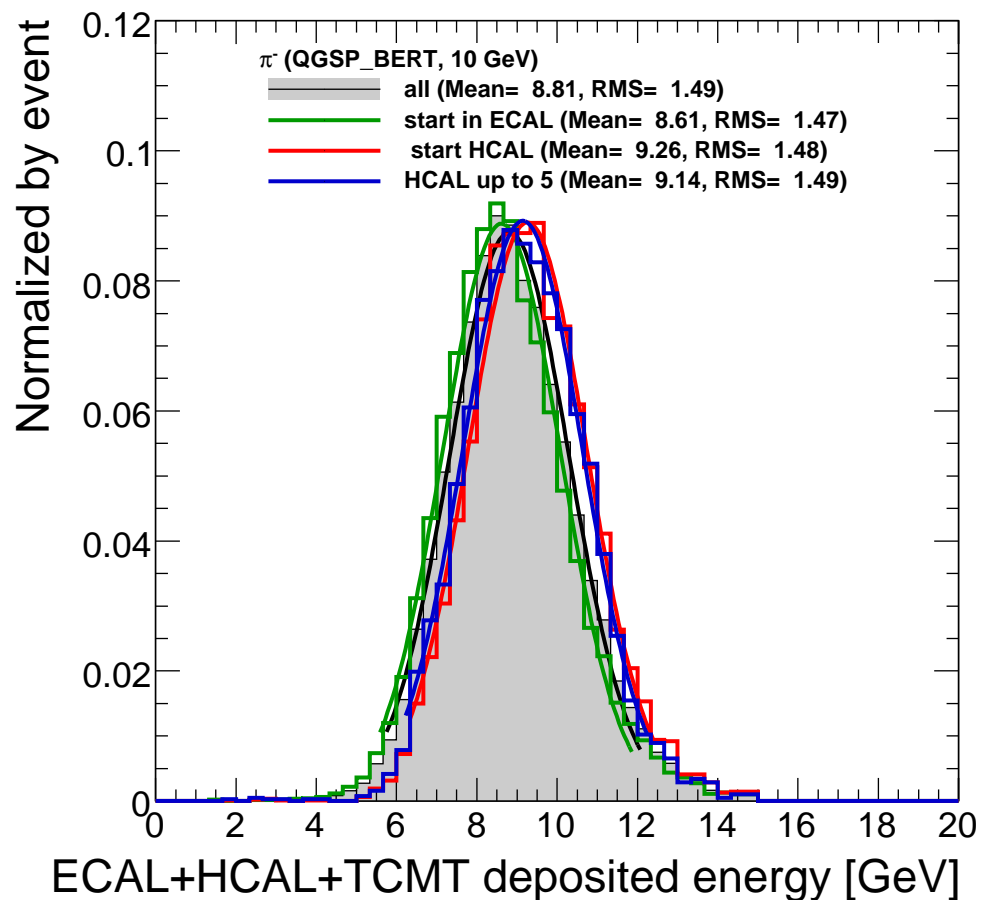
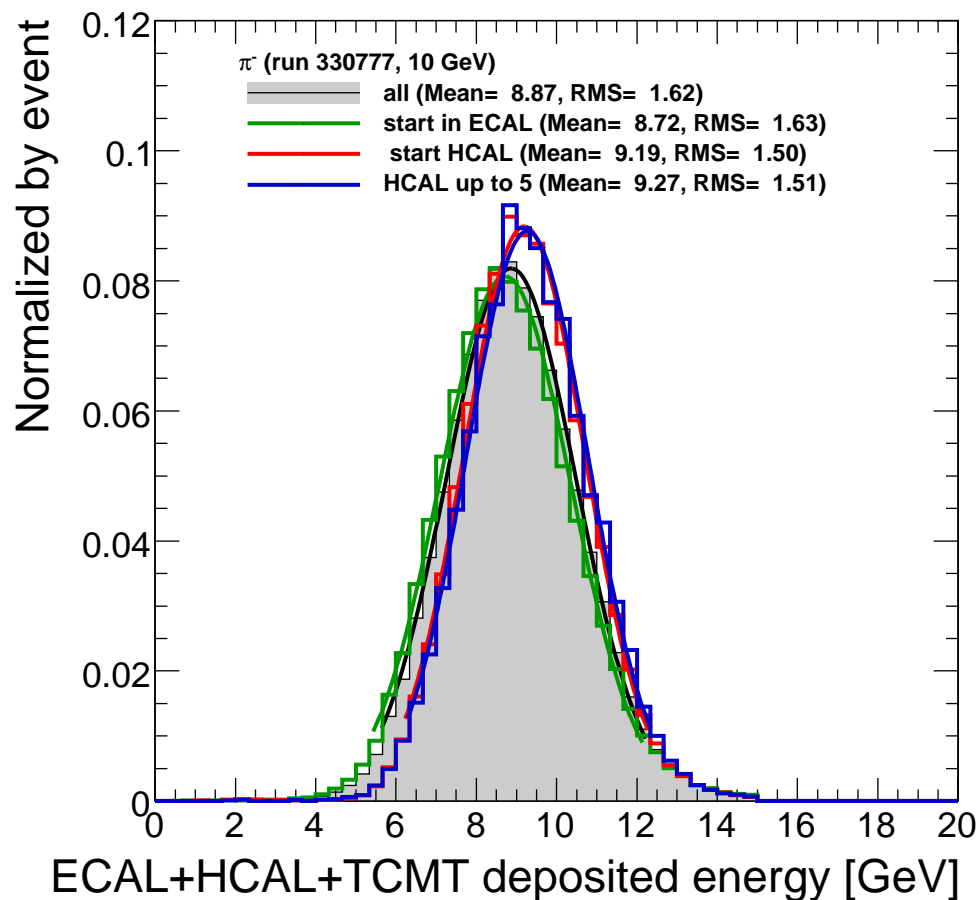
**HCAL:**  $\left(\frac{\text{GeV}}{\text{MIP}}\right)_{vis}^{hcal} = 0.000816$ ;  $S_{hcal} = 31.22$

## Event selection: energy distributions for 80-GeV $\pi^+$



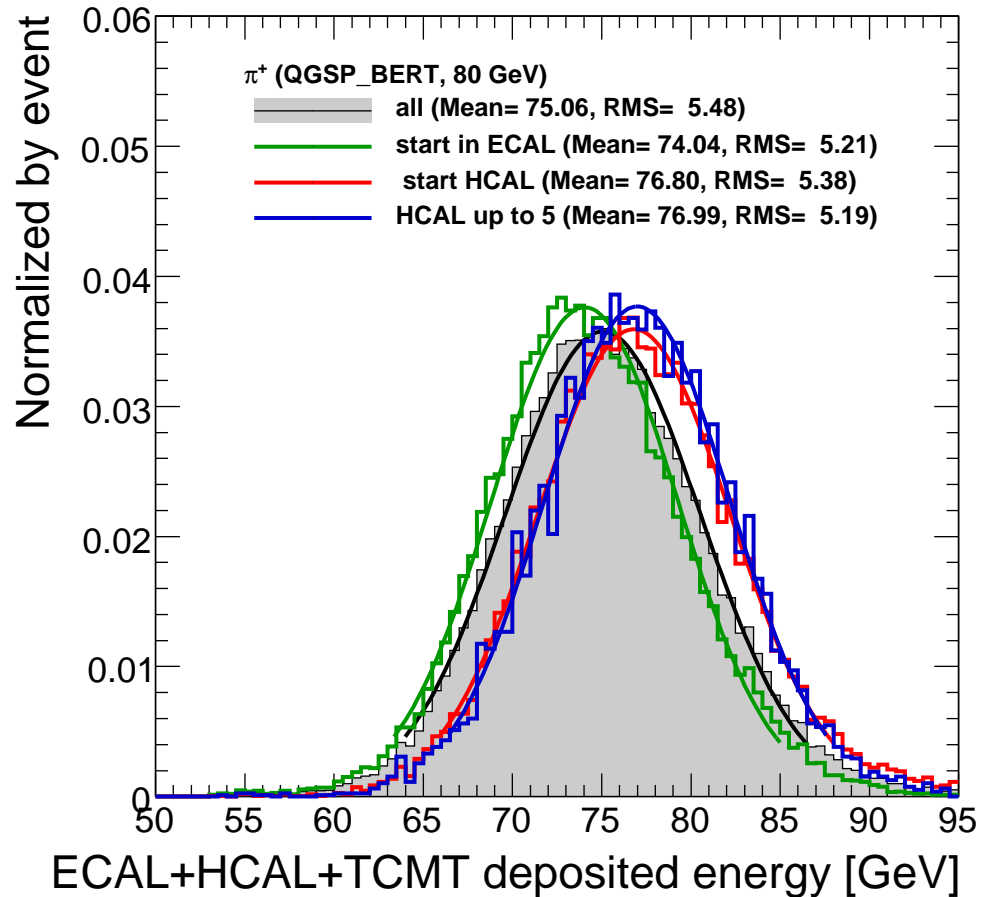
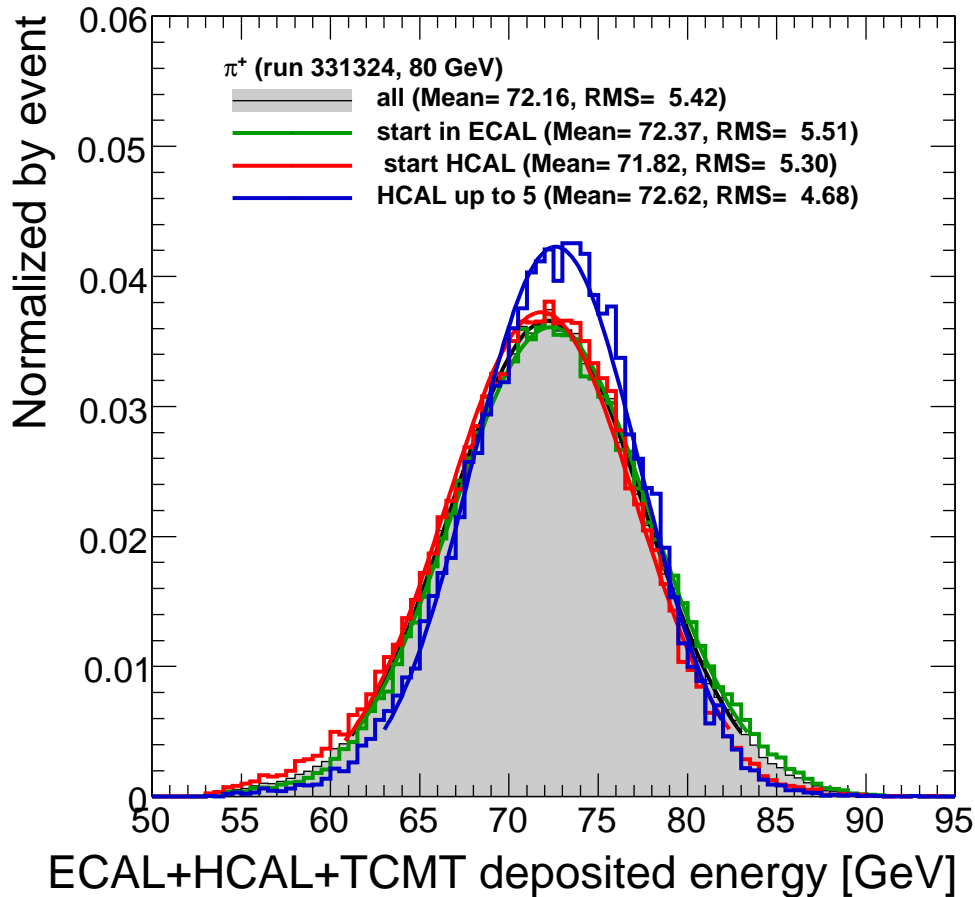
The same sampling factors and coefficients as in the previous slide.

## Event selection: energy distributions for 10-GeV $\pi^-$



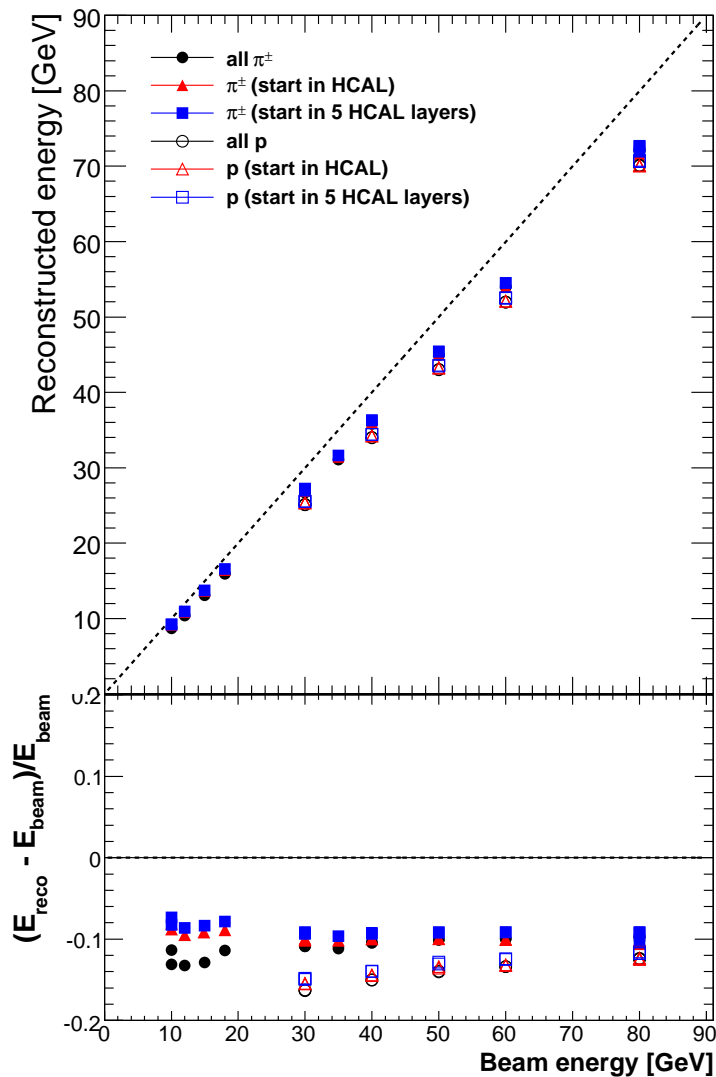
$S_{ecal} = 29.2$ . For data, the selection of events with shower start in HCAL improves energy resolution and shifts the mean value up by  $\sim 5\%$ . The shift predicted by QGSP\_BERT model is of the same order of magnitude.

## Event selection: energy distributions for 80-GeV $\pi^+$

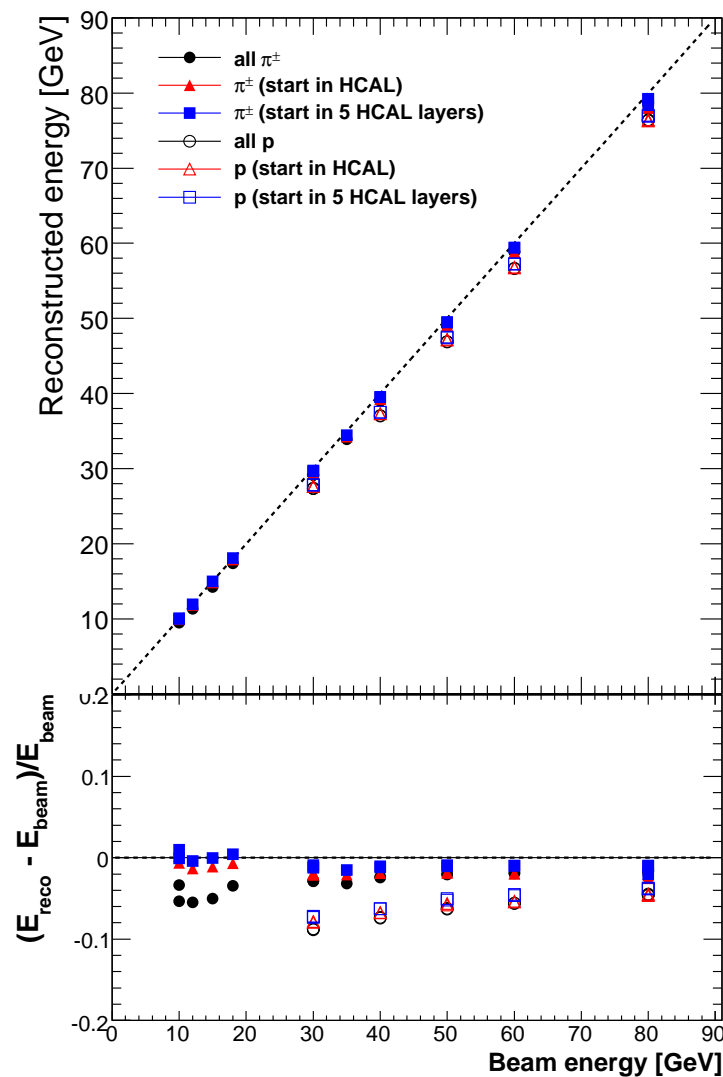


$S_{ecal} = 29.2$ . The mean value shift predicted by QGSP\_BERT is  $\sim 5\%$  as for 10 GeV while for 80-GeV data no such a shift is observed. The more significant RMS improvement can be seen for data than for QGSP\_BERT.

# Event selection: linearity



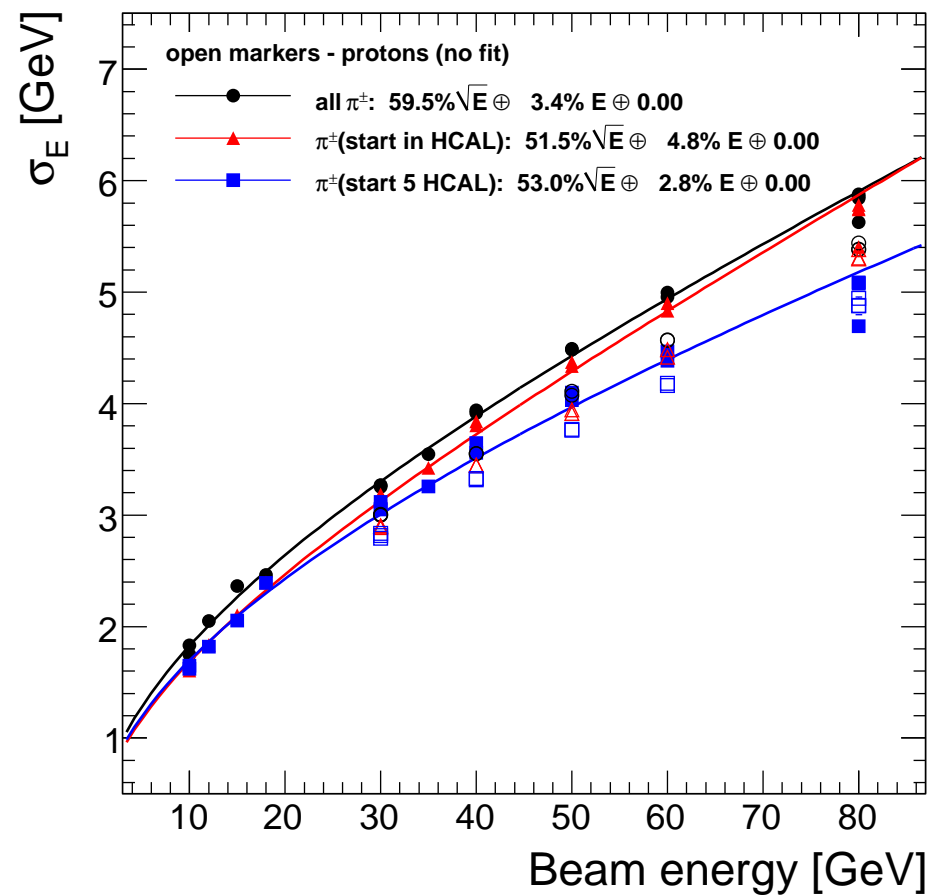
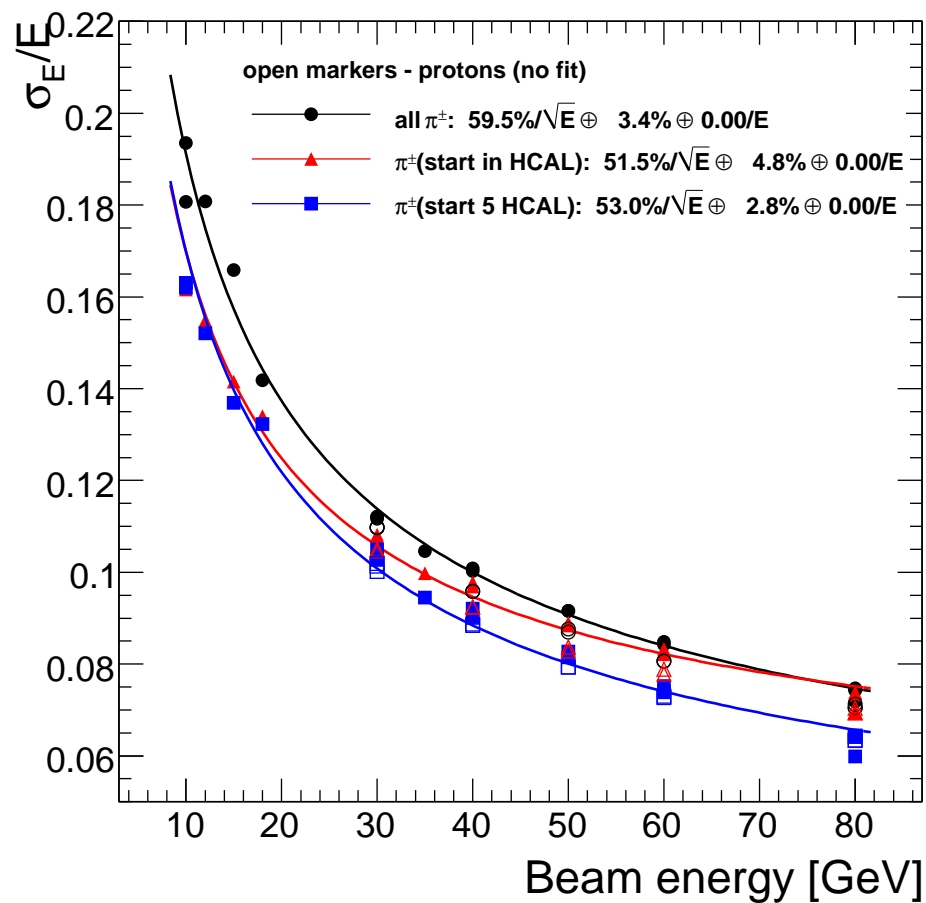
No scaling



$E_{reco}$  multiplied by 1.09 (e/ $\pi$  ratio)

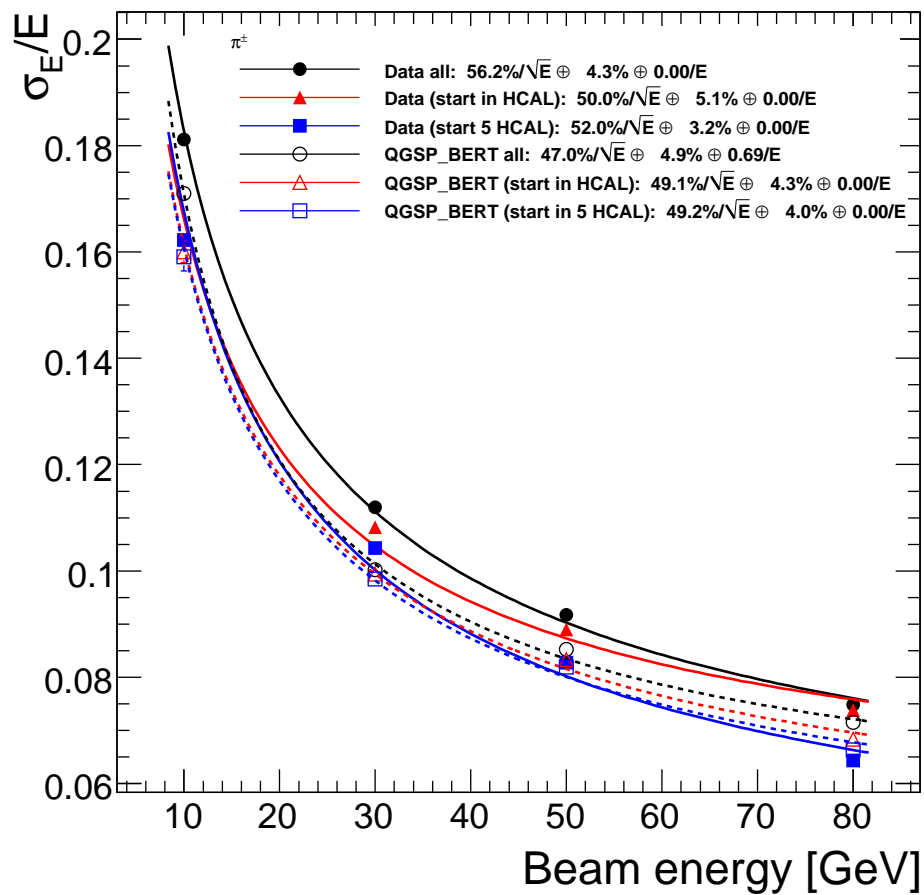


## Event selection: energy resolution for data

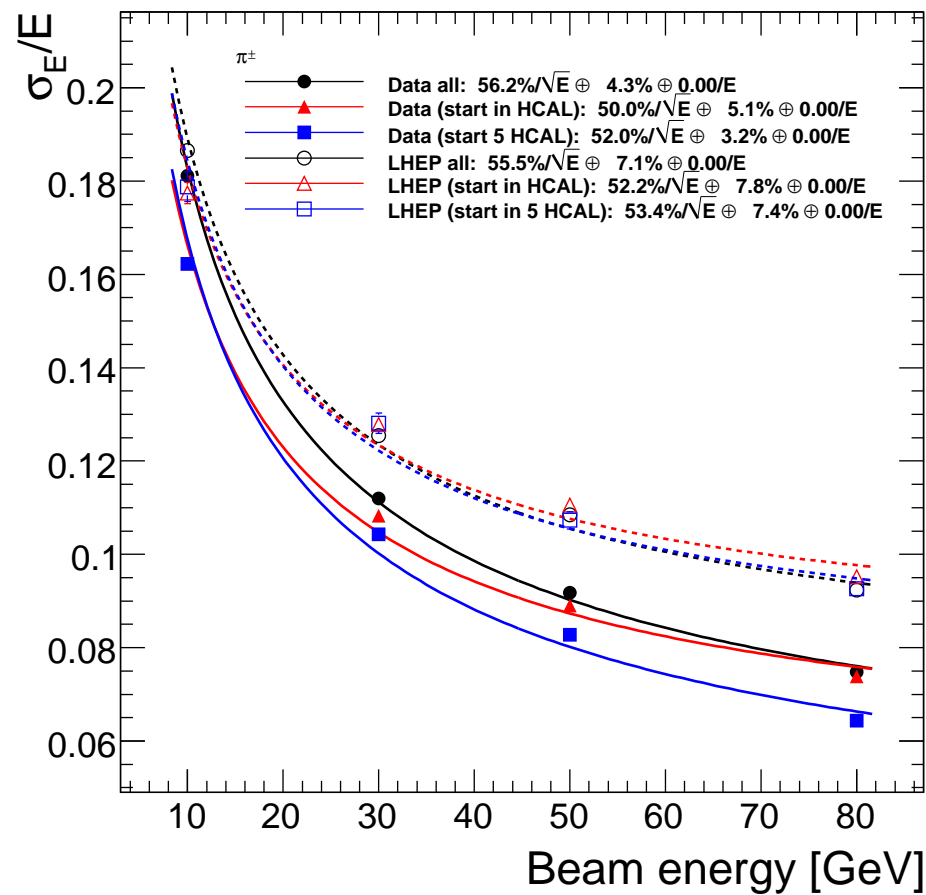


The increase of difference between start anywhere in HCAL and start in first 5 HCAL layers with increasing beam energy is due to higher probability of leakage for higher energies.

# Event selection: energy resolution for MC



QGSP\_BERT



LHEP

## Event selection: summary

calibration inconsistency (ECAL size?)  $\Rightarrow$  resolution  $\downarrow\downarrow$  by  $\sim 8\%$  @ low energies

shower leakage into TCMT  $\uparrow\uparrow \Rightarrow$  resolution  $\downarrow\downarrow$  by  $\sim 8\%$  @ high energies

How to avoid? Select events with shower start in the first 5 HCAL layers

$\downarrow\downarrow$

$$\frac{\sigma_E}{E} = \frac{(53.0 \pm 0.2)\%}{\sqrt{E/\text{GeV}}} \oplus (2.8 \pm 0.1)\% \oplus \frac{0.00 \pm 0.05}{E/\text{GeV}}$$

linearity  $\sim 2\%$  for  $\pi$ , while  $\sim 8\%$  for protons

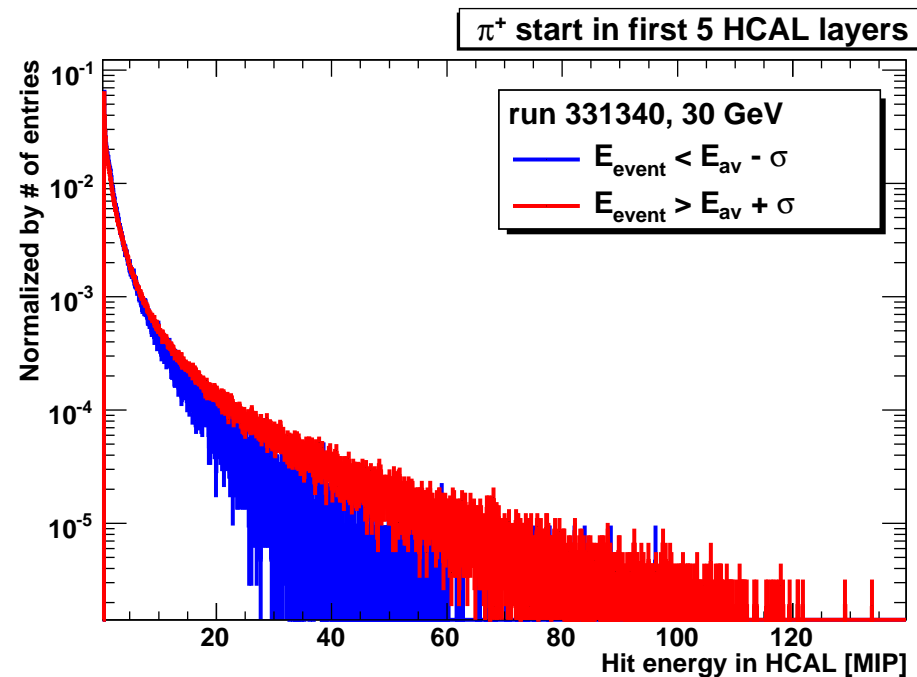
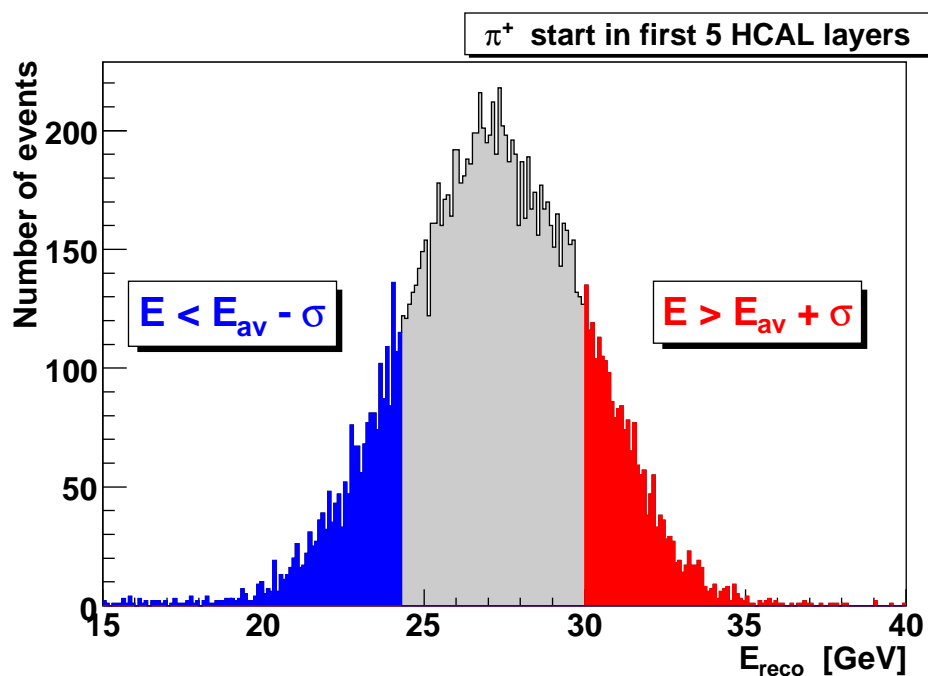
For the CALICE AHCAL  $e/\pi \approx 1.09$  for 10–80 GeV

## Software compensation: approach

based on hit spectrum analysis  $\Rightarrow$  only events with shower start in first 5 HCAL layers will be analyzed

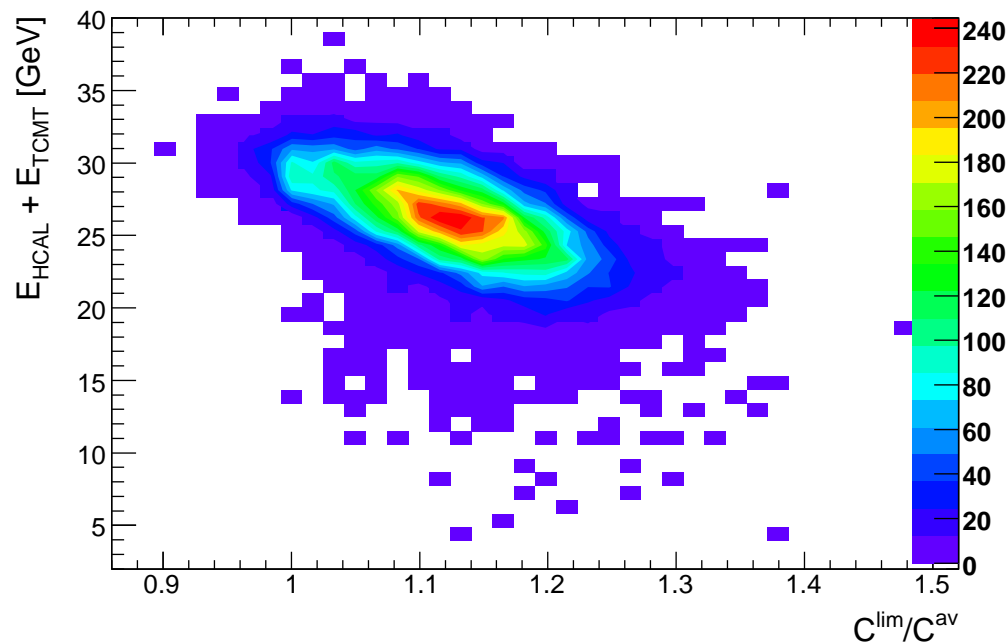
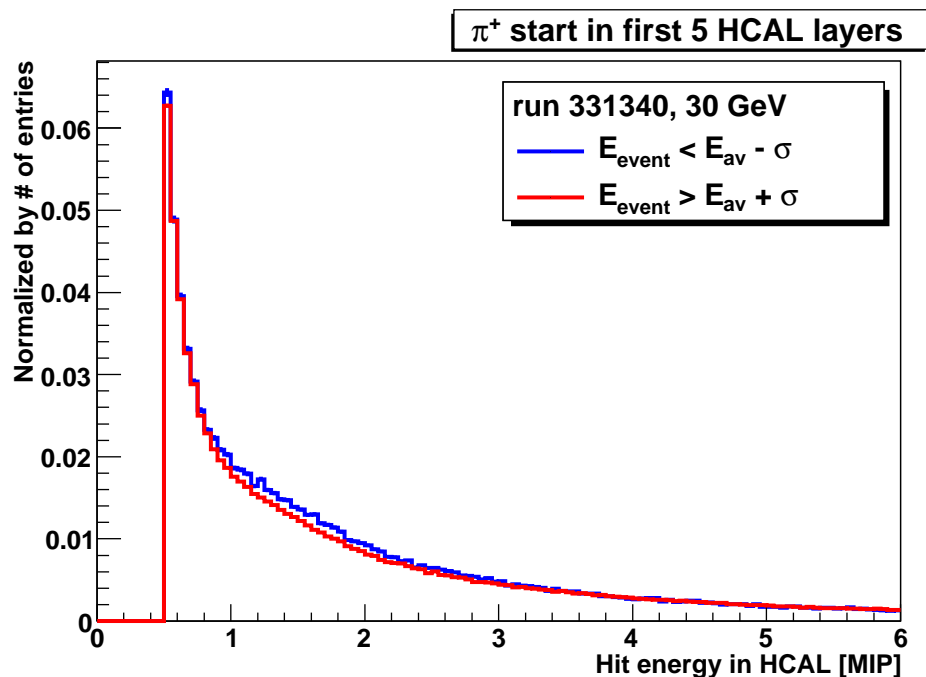
Energy distribution for 30-GeV  $\pi^+$

Hit spectra for energy distribution tails



The technique proposed in CAN-015 and also based on hit spectrum analysis includes individual hit weighting in 8 regions of hit spectrum and 6 parameters to correct energy dependence.

# Software compensation: integral spectrum characteristics



The following integral values can characterize a hit spectrum  $h_i(e)$  of the  $i$ -th event:

$$C_i^{\text{lim}} = \int_0^{e_{\text{limit}}} h_i(e) de \quad \text{and} \quad C_i^{\text{av}} = \int_0^{e_i^{\text{av}}} h_i(e) de$$

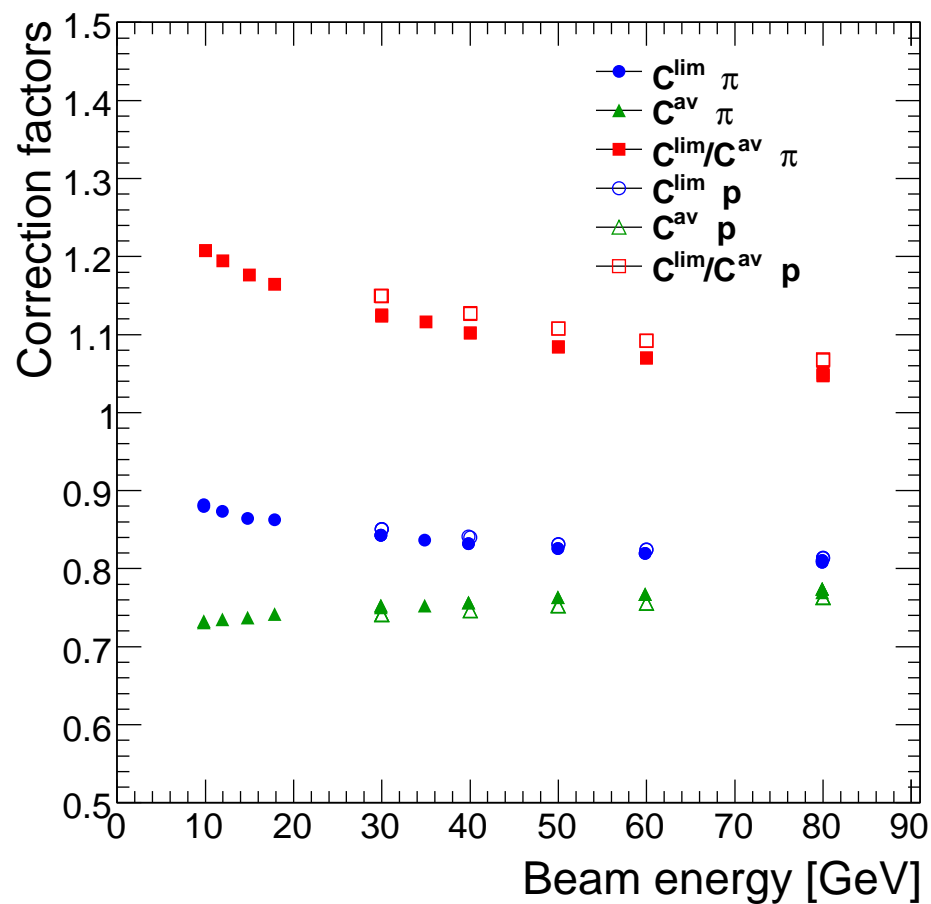
where  $e_{\text{limit}} = 5.5 \text{ MIPs}$  and  $e_i^{\text{av}} = \int_0^{e_{\text{max}}} e h_i(e) de$

The ratio  $\frac{C_i^{\text{lim}}}{C_i^{\text{av}}}$  is inversely correlated with the energy deposited in the event.

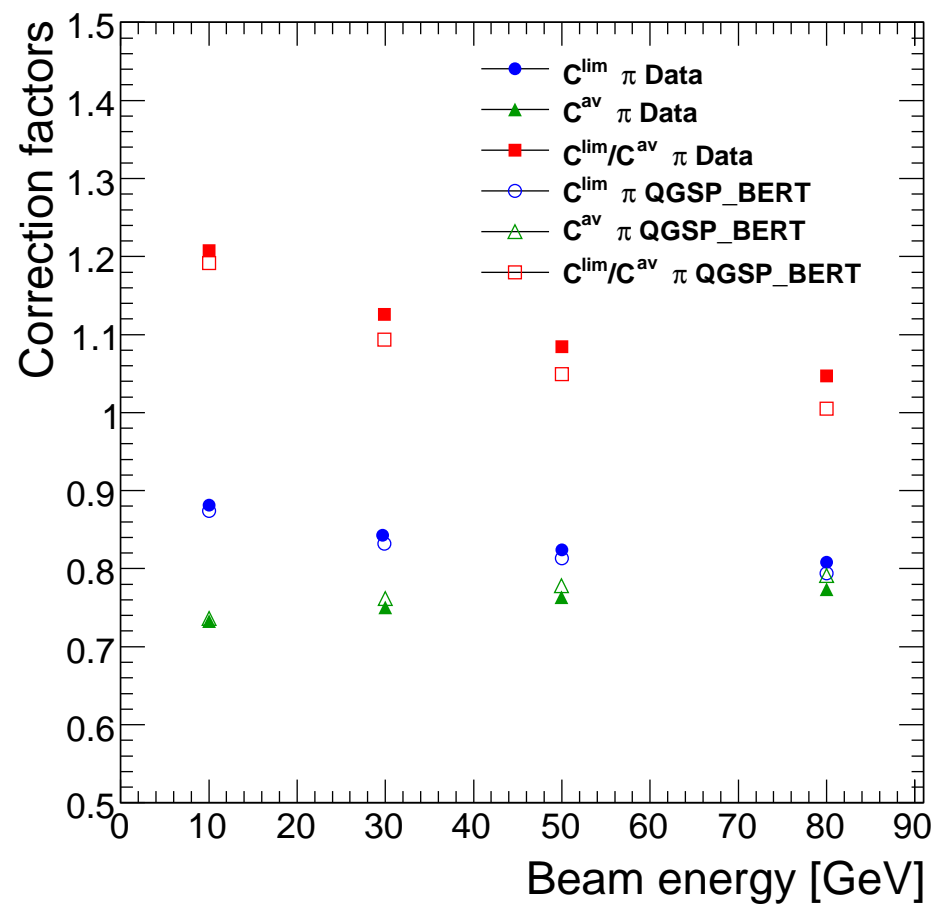
# Software compensation: correction factor behavior

The mean value of the correction factor  $\frac{C^{lim}}{C^{av}}$  exhibits a relatively weak but visible energy dependence

Data  $\pi^+$  and Data  $p$



Data  $\pi^+$  and QGSP\_BERT  $\pi^+$



This dependence can be fitted by linear or parabolic function.

## Software compensation: correction procedure for i-th event

Calculate uncorrected deposited energies  $E_i^{HCAL}$ ,  $E_i^{ECAL}$ , and  $E_i^{TCMT}$



Calculate correction factors  $C_i^{lim}$  and  $C_i^{av}$



Correct HCAL energy:  $E_i^{cor} = E_i^{HCAL} \cdot \frac{C_i^{lim}}{C_i^{av}}$



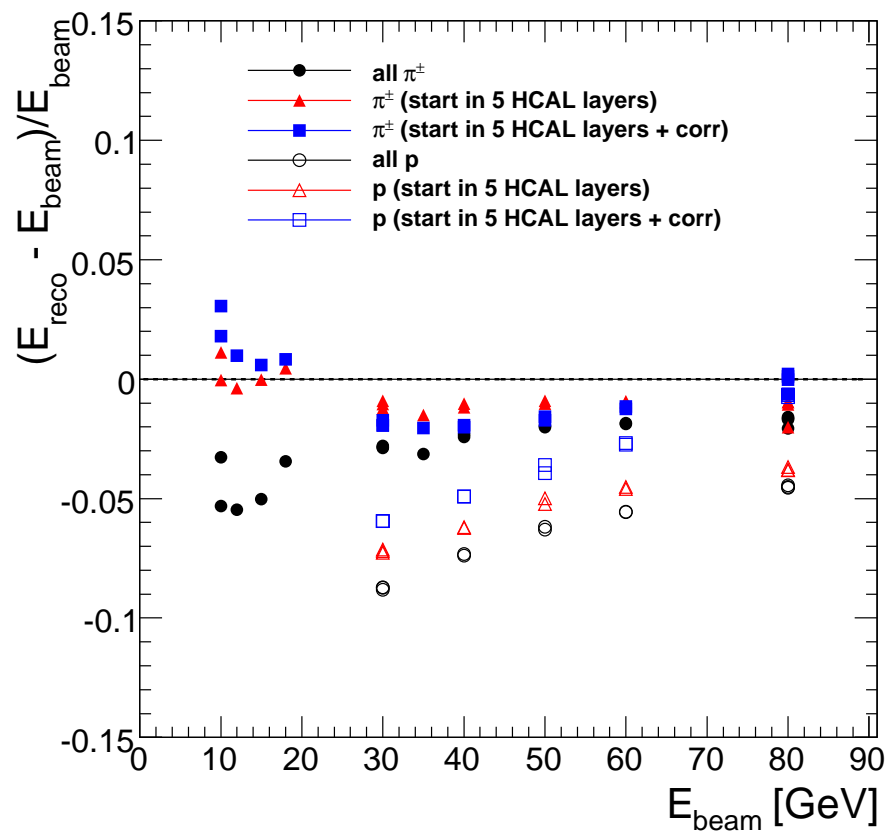
Calculate total energy:  $E_i^{dep} = E_i^{ECAL} + E_i^{cor} + E_i^{TCMT}$



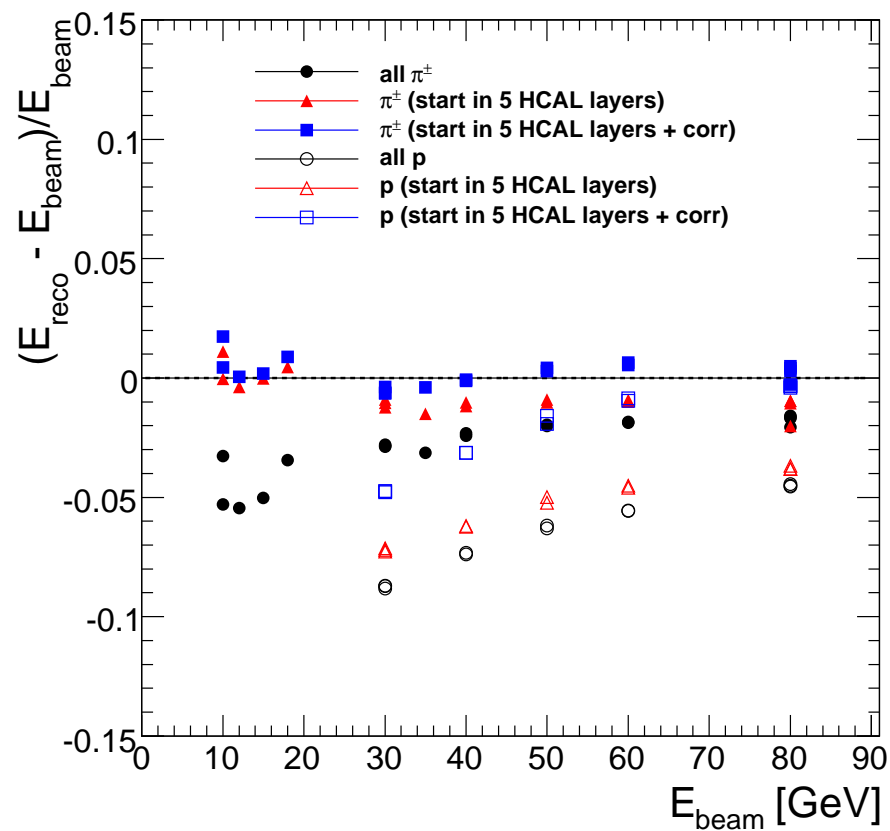
Correct energy dependence  $E_i^{reco} = E_i^{dep} (a_1 + a_2 E_i^{dep})$ ,  $a_1 = 0.91$ ,  $a_2 = 0.002 \text{ GeV}^{-1}$  or  
 $E_i^{reco} = E_i^{dep} (a_1 + a_2 E_i^{dep} + a_3 (E_i^{dep})^2)$ ,  $a_1 = 0.876$ ,  $a_2 = 0.0043 \text{ GeV}^{-1}$  and  $a_3 = -2.4 \cdot 10^{-5} \text{ GeV}^{-2}$

**Only 3 (or 4) parameters and no adjustment of ranges in hit spectra**

## Software compensation: linearity after correction



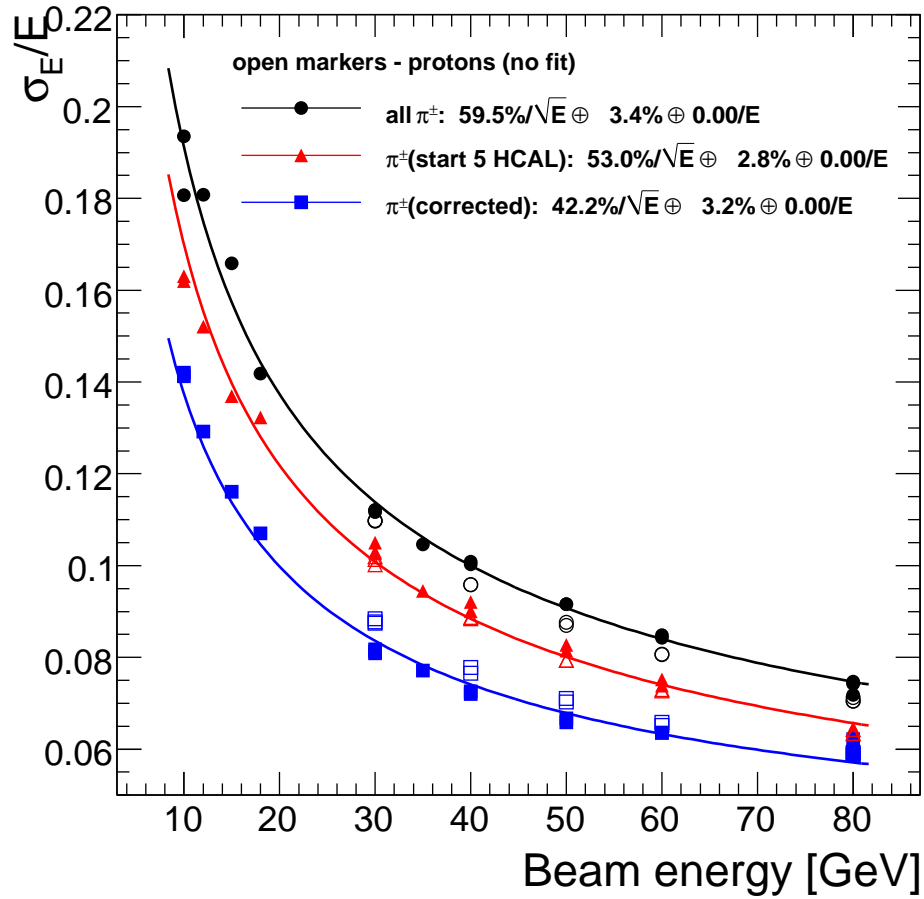
Linear



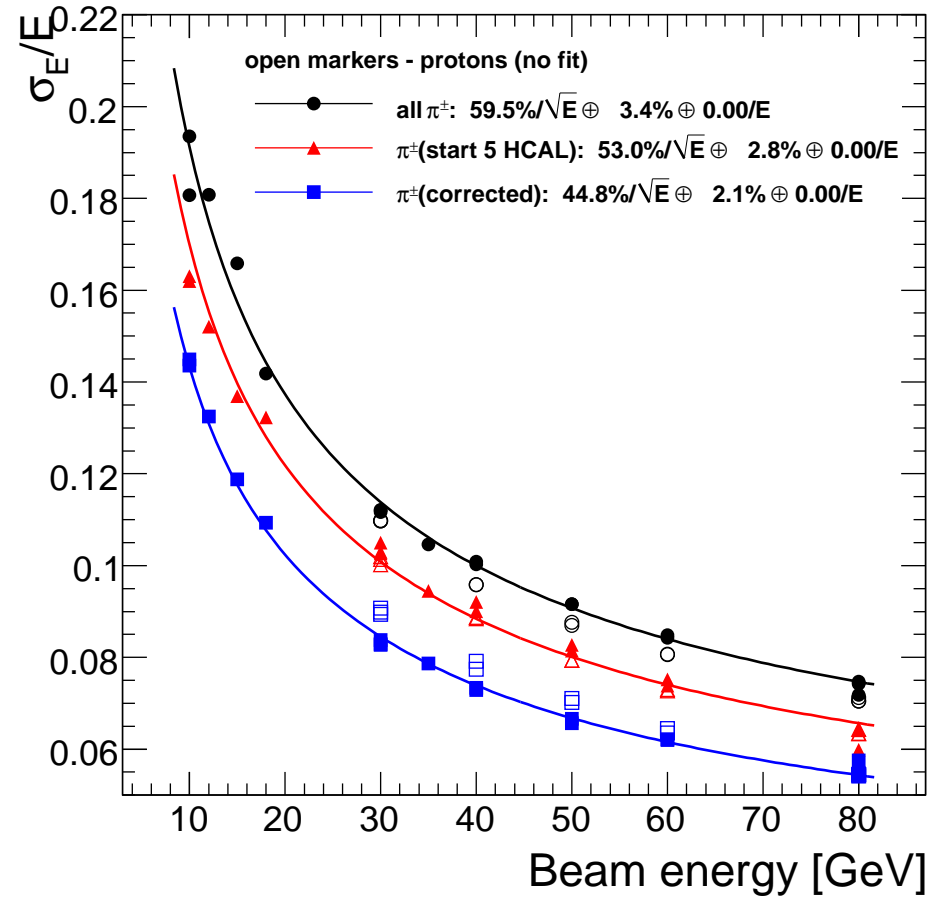
Parabolic



# Software compensation: resolution after correction

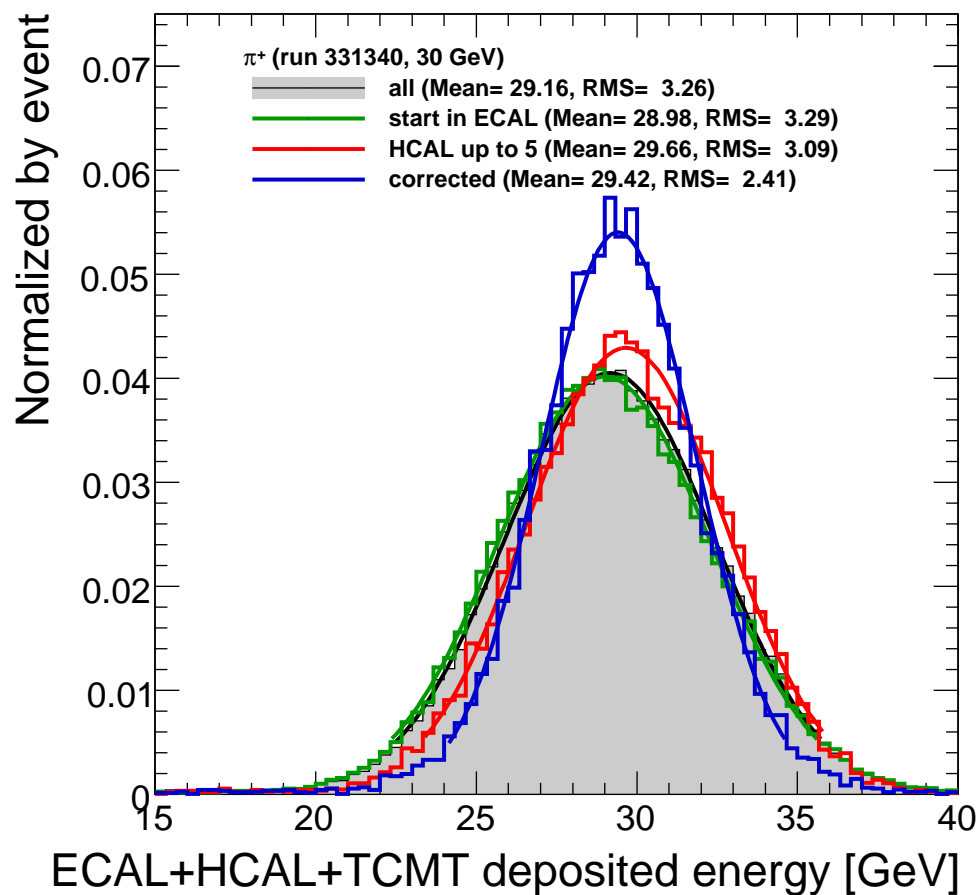


Linear

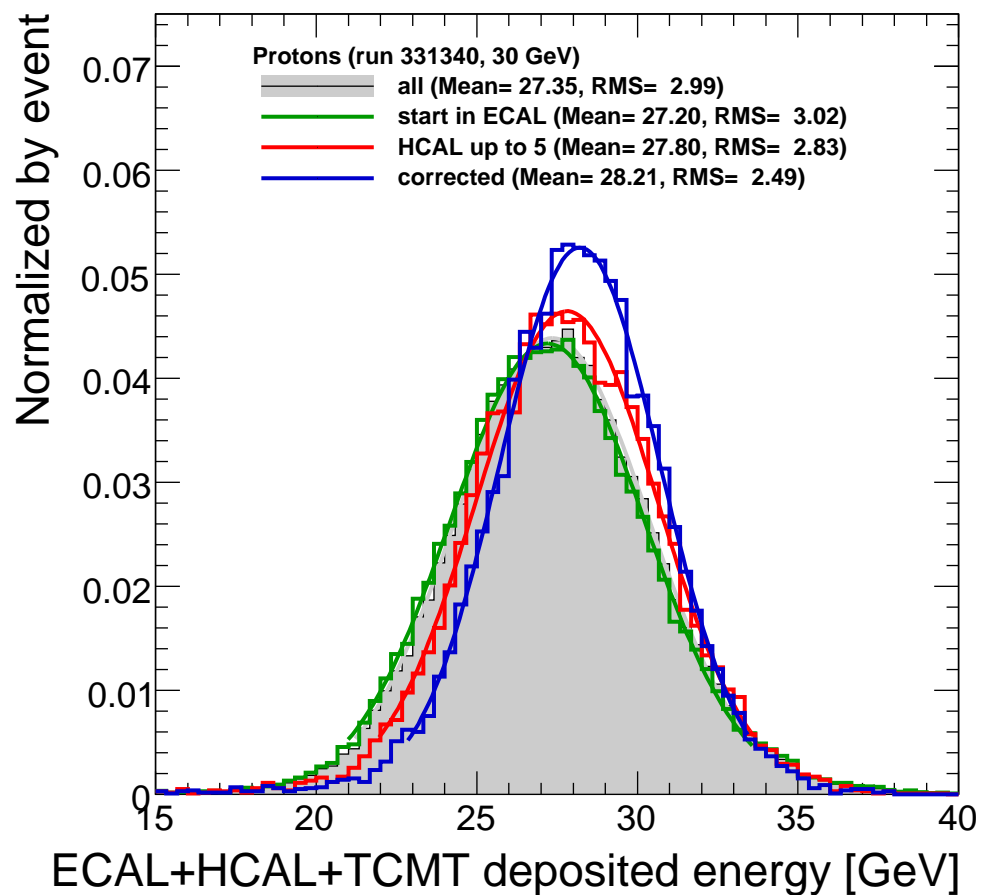


Parabolic

## Software compensation: energy distributions after correction

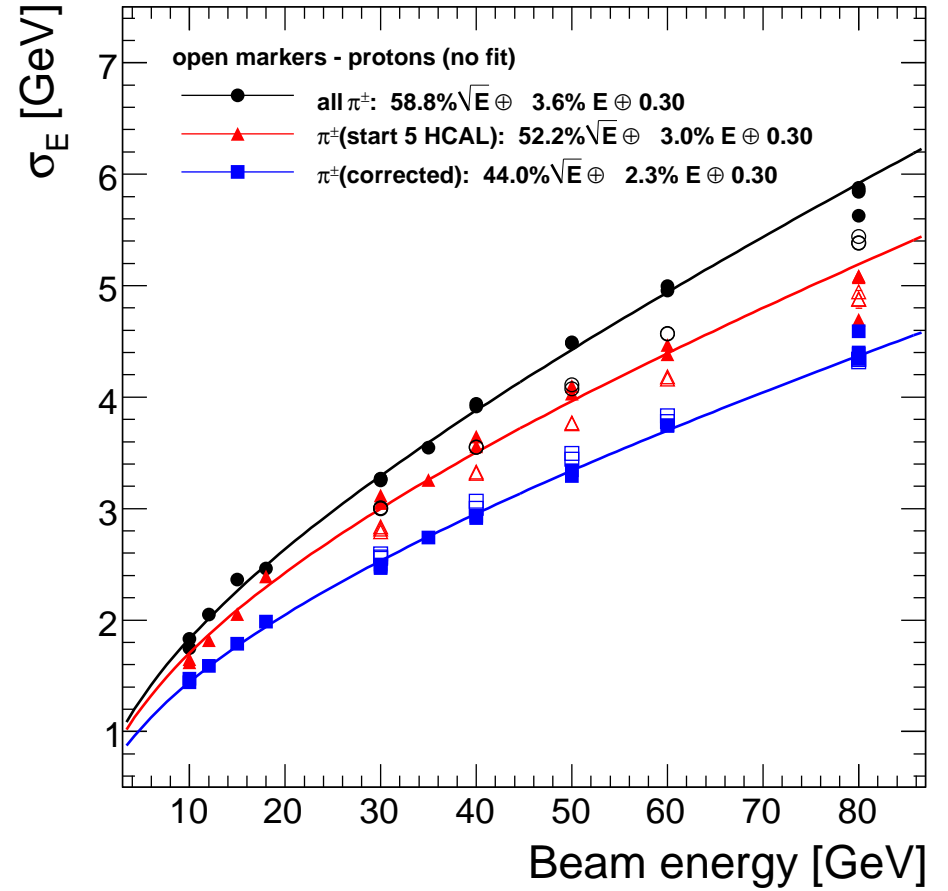
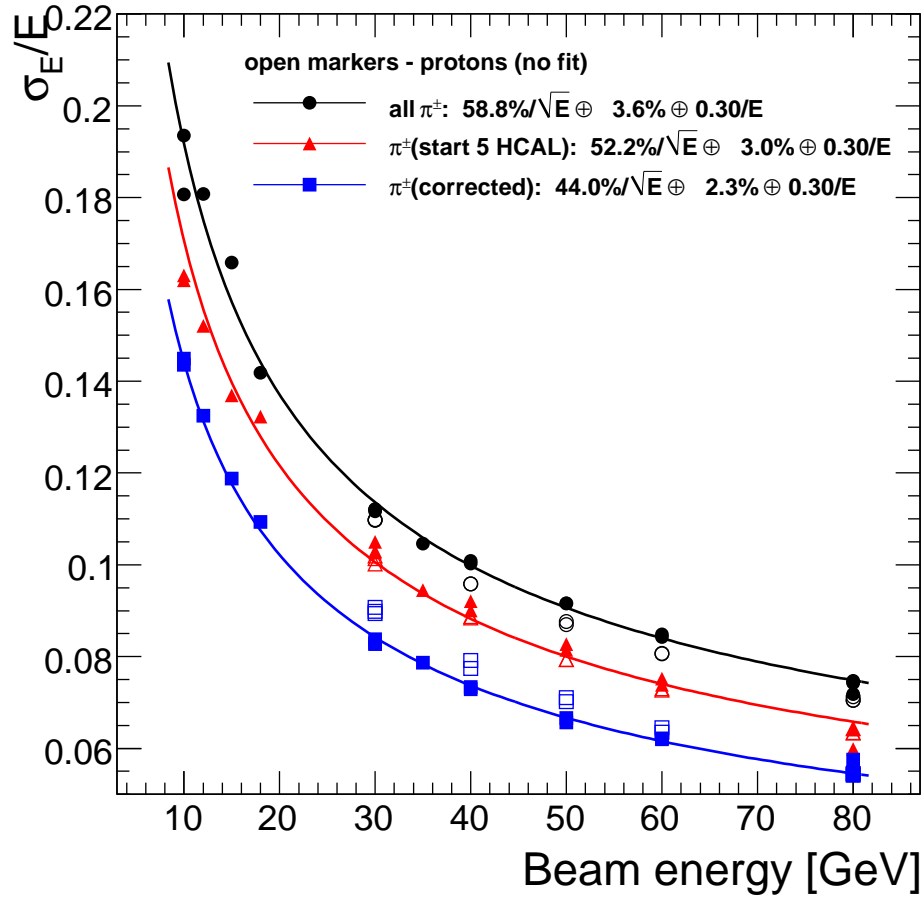


$\pi^+$



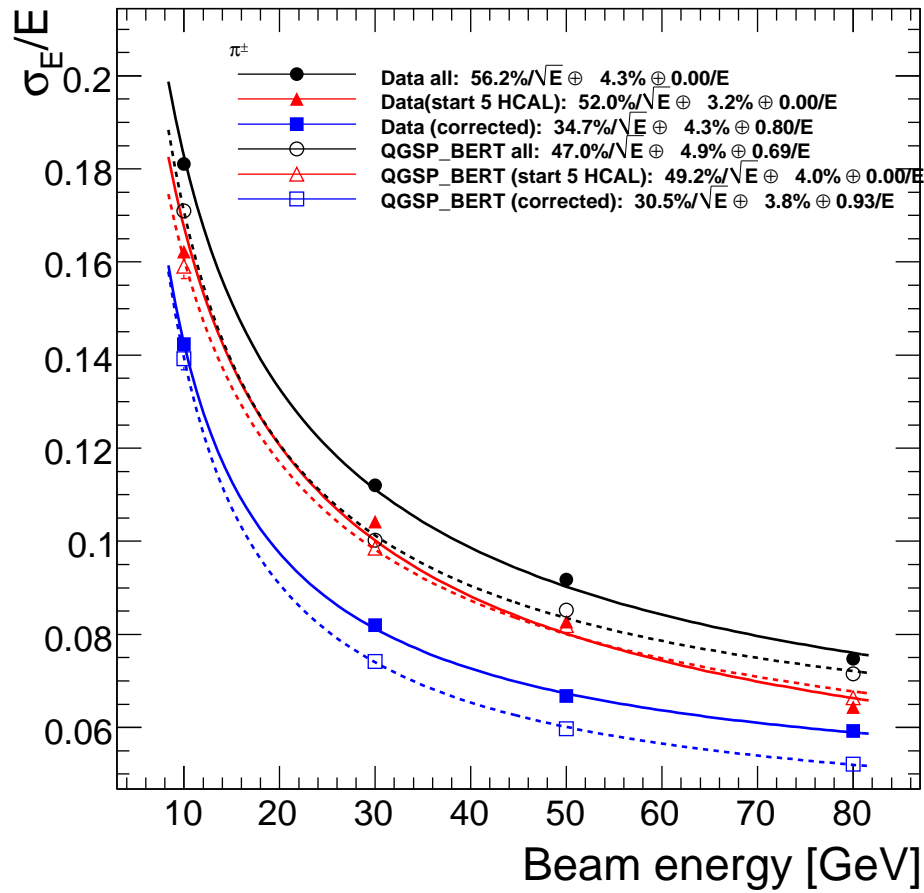
proton

# Software compensation: fit with fixed $\sigma_{noise}$

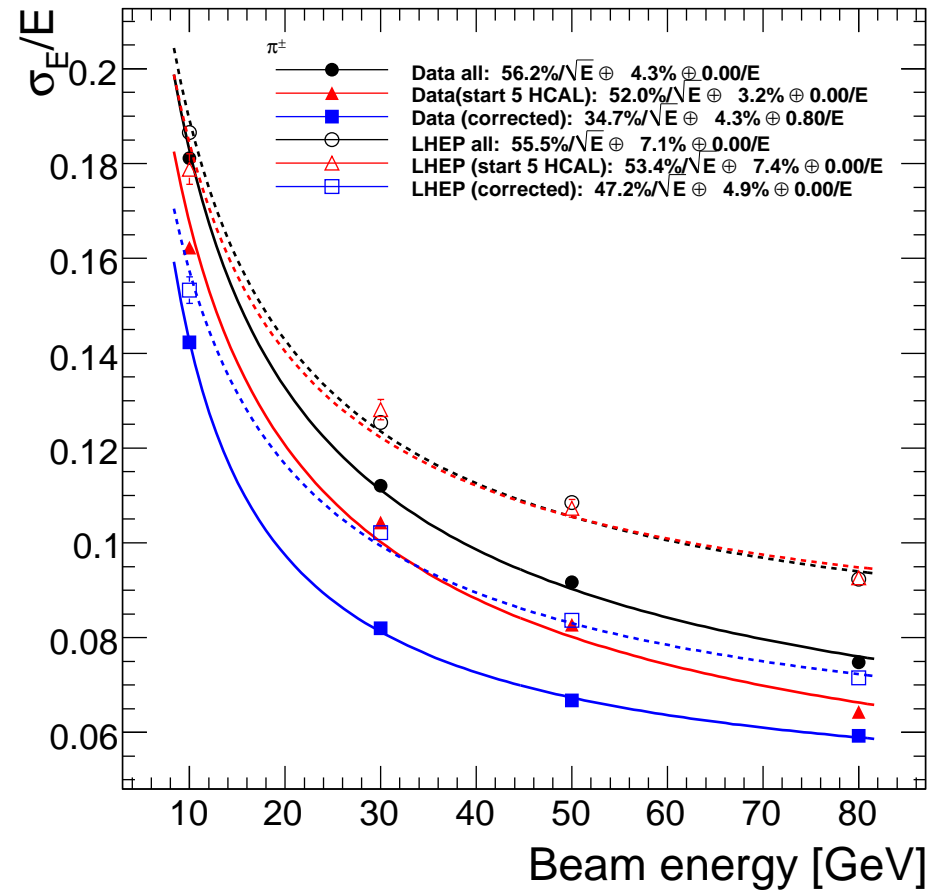


$$\frac{\sigma_E}{E} = \frac{(44.0 \pm 0.2)\%}{\sqrt{E/GeV}} \oplus (2.3 \pm 0.1)\% \oplus \frac{0.3}{E/GeV}$$

# Software compensation: MC resolution after correction

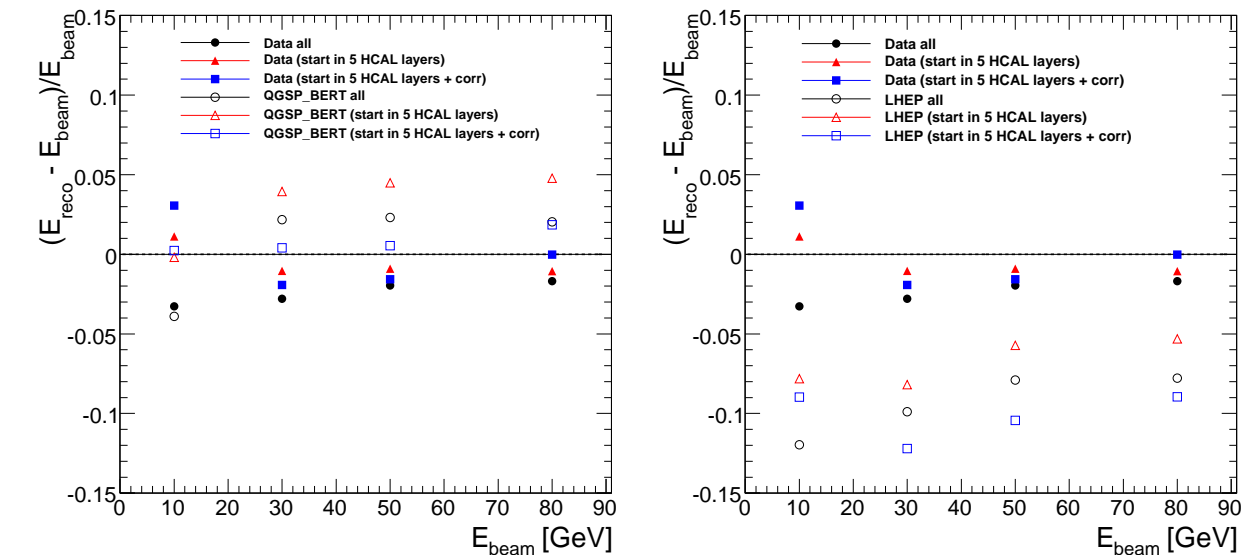


QGSP\_BERT

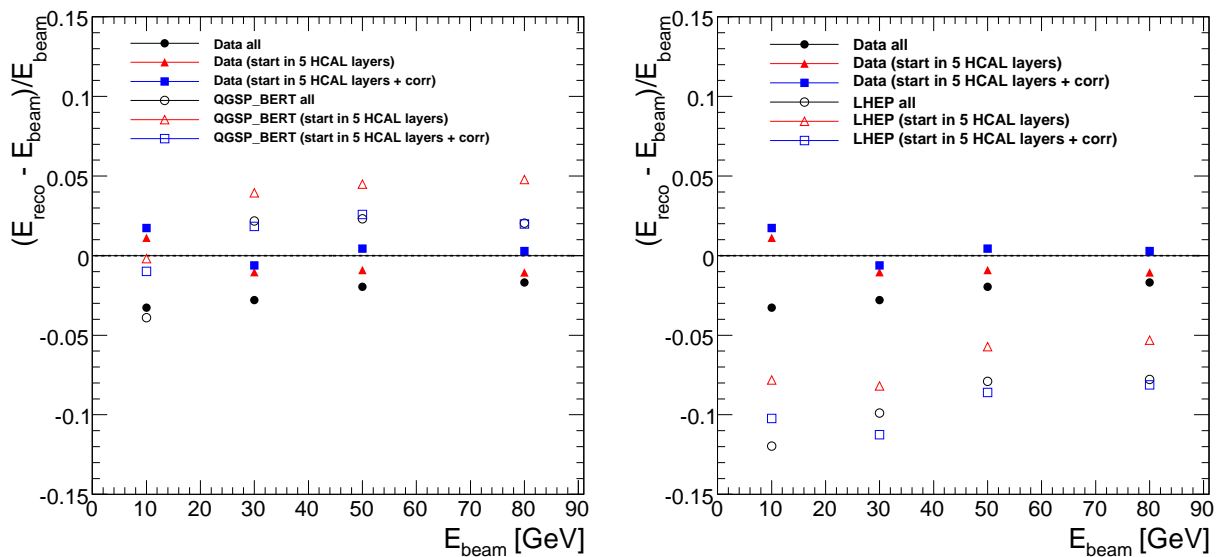


LHEP

# Software compensation: MC linearity after correction



Linear



QGSP\_BERT

LHEP

Parabolic

## Software compensation: summary

Proposed method of software compensation enables to improve resolution by  $\sim 10\%$

### Advantages

3 (or 4) parameters only

weak energy dependence

similar behavior of MC resolution

### Problems

full hit spectrum necessary

energy range extension

distortion of MC linearity

### TO DO:

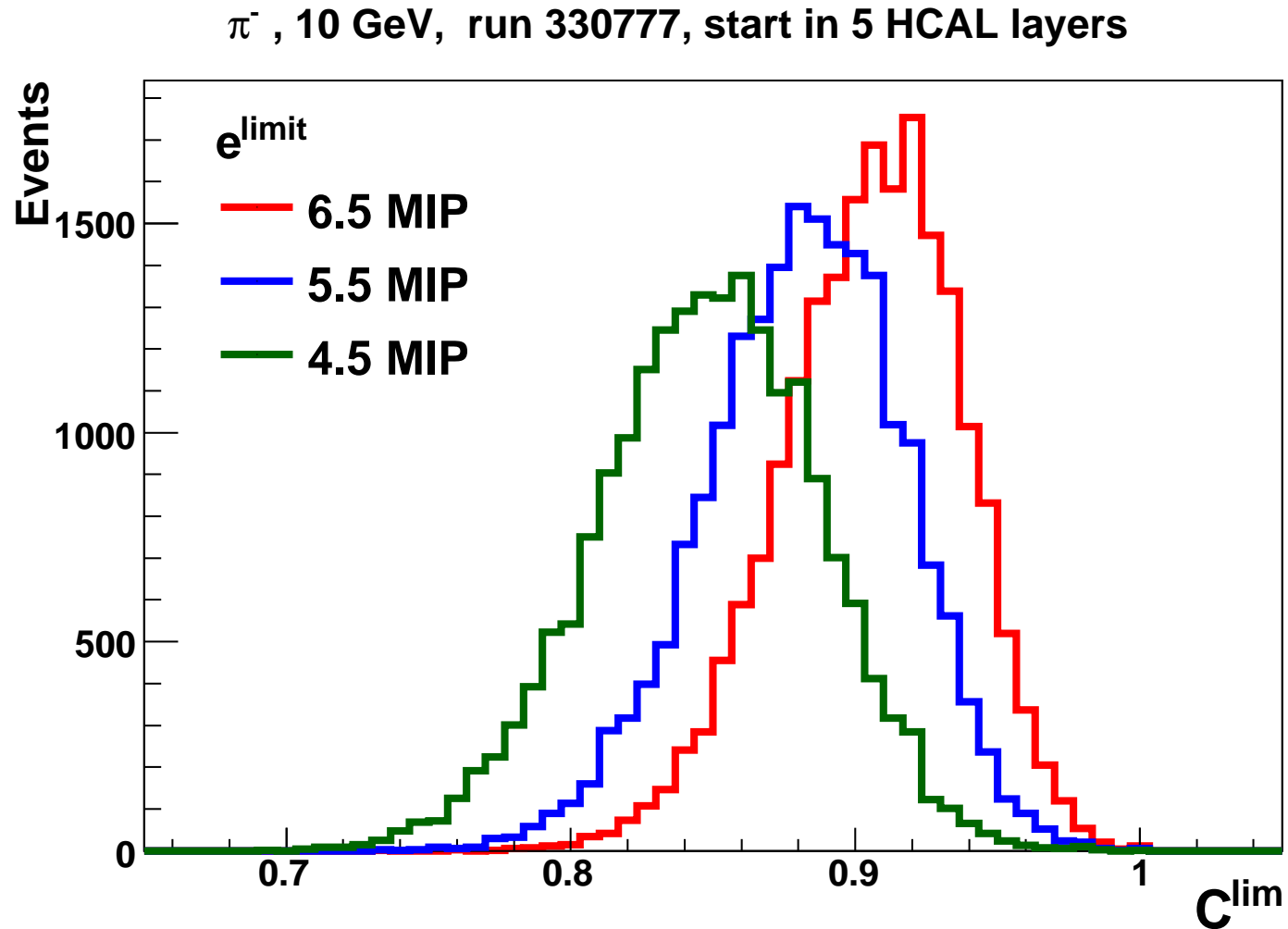
Hadron selection procedure for lower energies

More energy points and simulated MC physics lists

Energy range expansion

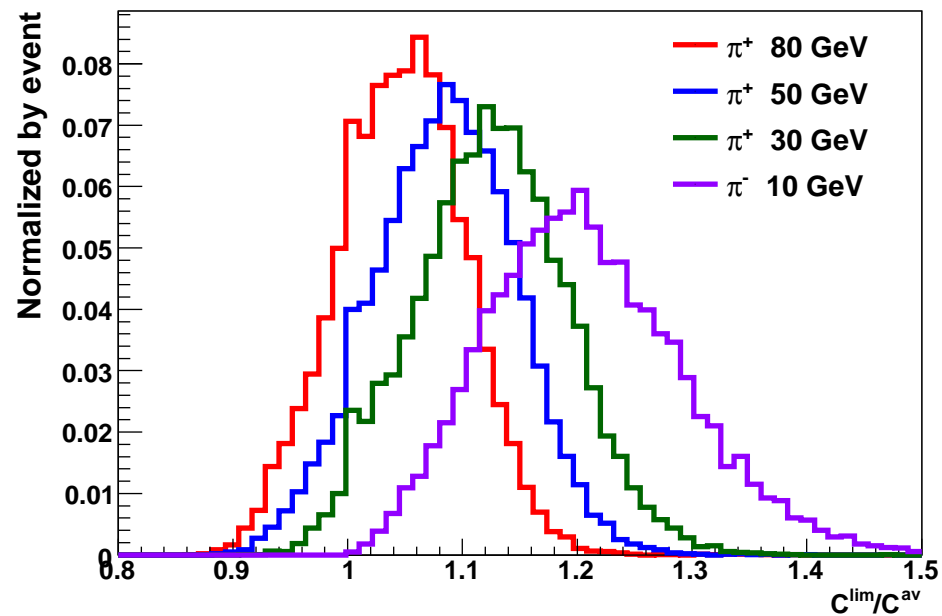
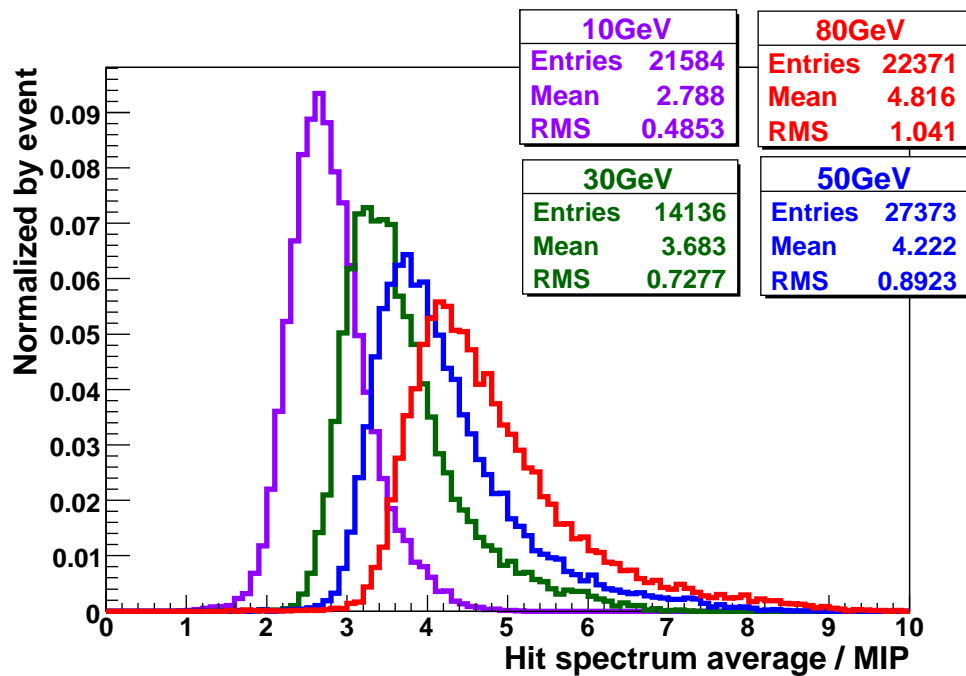
# Backup slides

## Software compensation: correction factor behavior





# Software compensation: correction factor behavior



## Software compensation: correction factor behavior

