



VERITAS

Imaging Calorimetry at Very High Energies



E. Hays

University of Chicago / Argonne National Lab



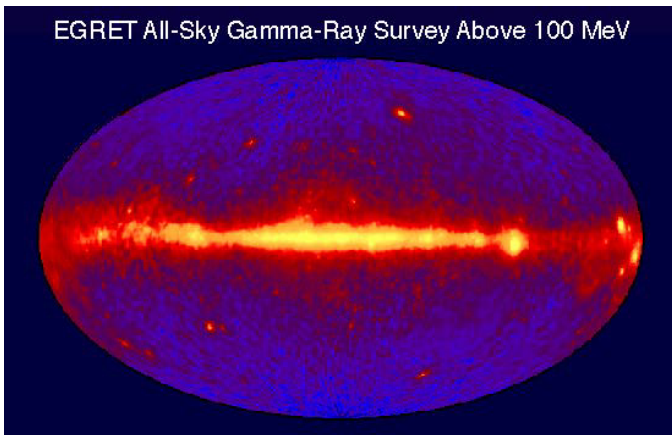
The Very Energetic Radiation Imaging Telescope Array System

- ▶ Description and status of VERITAS
- ▶ Air showers and the stereoscopic imaging technique
- ▶ Background rejection of hadronic showers
- ▶ Energy resolution



Astrophysics above 50 GeV

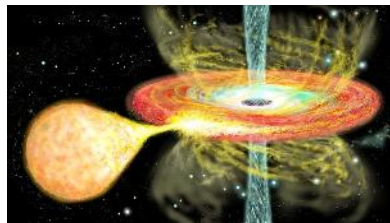
EGRET All-Sky Gamma-Ray Survey Above 100 MeV



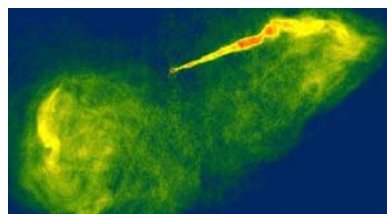
Probing galactic and extragalactic acceleration processes, particle populations, and interactions

????
Dark Matter?
GRBs?
Galactic Clusters?

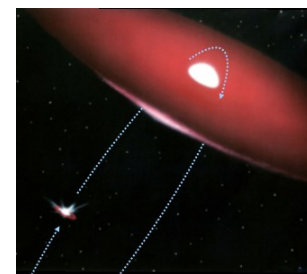
Microquasars



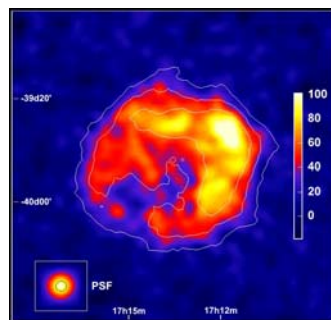
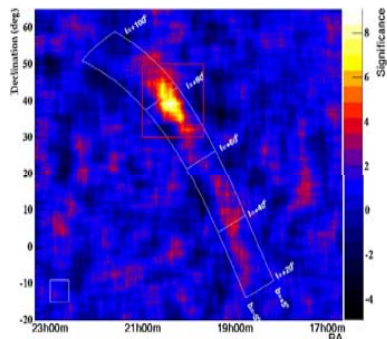
Blazars



Compact Binary Systems



Diffuse Galactic Emission

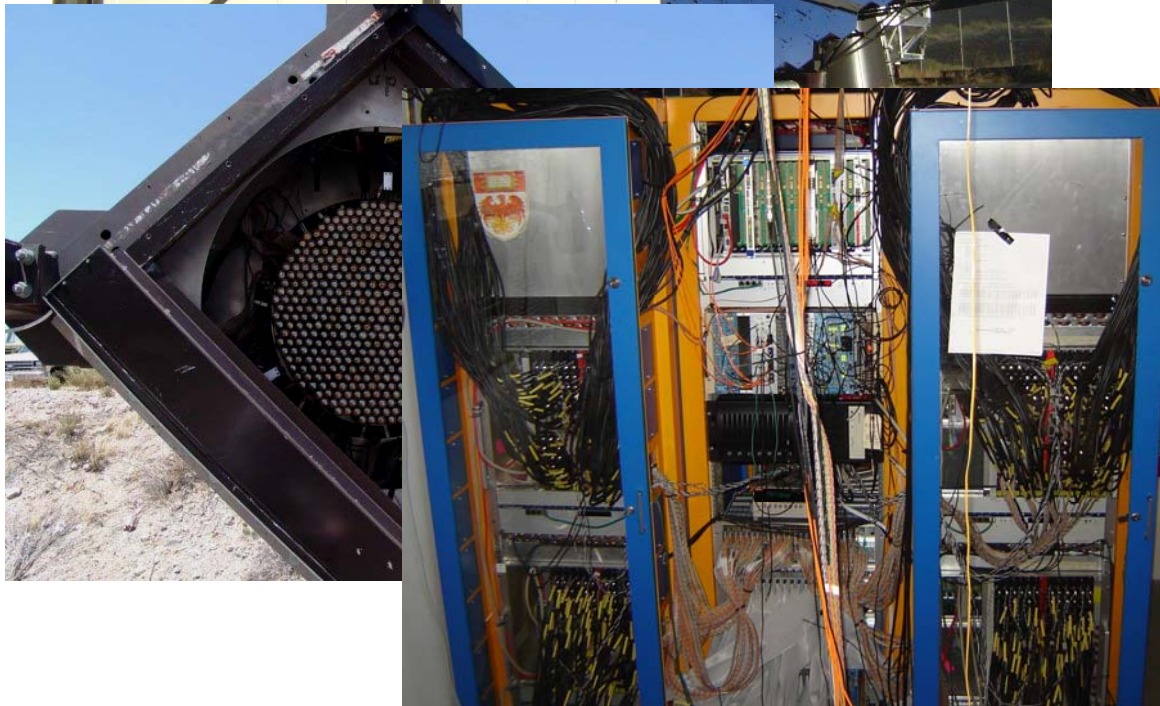


Pulsar Wind Nebulae

Supernova Remnants



VERITAS



- ▶ System of 4 imaging air Cerenkov telescopes to detect γ -rays above >50 GeV
 - 12-m reflectors, 110 m^2 area
 - 499 PMT cameras, 3.5° fov, 0.15° pixel spacing
 - 500 MHz FADC
 - 3 level trigger
 - **channel**: constant fraction discriminators, ~ 5 PE threshold
 - **telescope**: pattern of neighboring channels
 - **array**: coincident telescope triggers



VERITAS Status



Future T3
T2
Future T4
T1

- Constructing array at the basecamp of Mt Hopkins in AZ
- Running 2 telescopes now using array level trigger
- Will complete the array in Fall 2006
- Will run the 4 telescope array at Mt Hopkins for 2 years

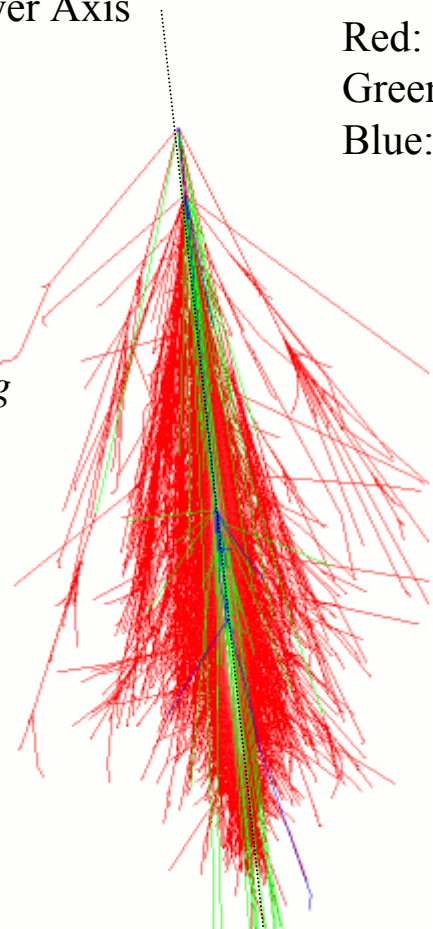


Imaging Air Showers

Primary Direction
Shower Axis

$\sim 10 \text{ g/cm}^2$ at
30 km a.s.l.

*density
increasing
 $\sim 1/e^{(\text{height})}$*



Simulated air shower using Corsika

Red: Electromagnetic

Green: Muons

Blue: Hadrons

Interaction length
TeV proton - air
 $\sim 80 \text{ g/cm}^2$

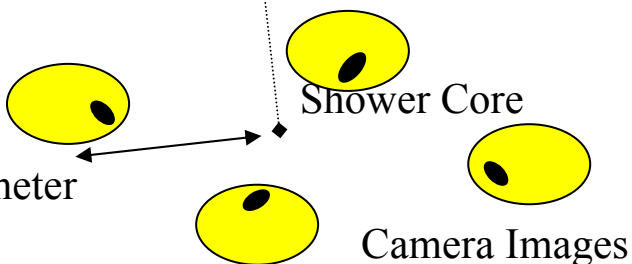
Electromagnetic
radiation length
 $\sim 34 \text{ g/cm}^2$

Shower Parameters

- Primary Type
- Primary Direction
- Primary Energy
- First Interaction Depth
- Shower Core Location
 - intersection of shower axis with detector plane

$\sim 800 \text{ g/cm}^2$
at 2 km a.s.l.

Impact Parameter



Cerenkov telescopes record 2D
projections in angular space of the track
of the shower through the atmosphere.

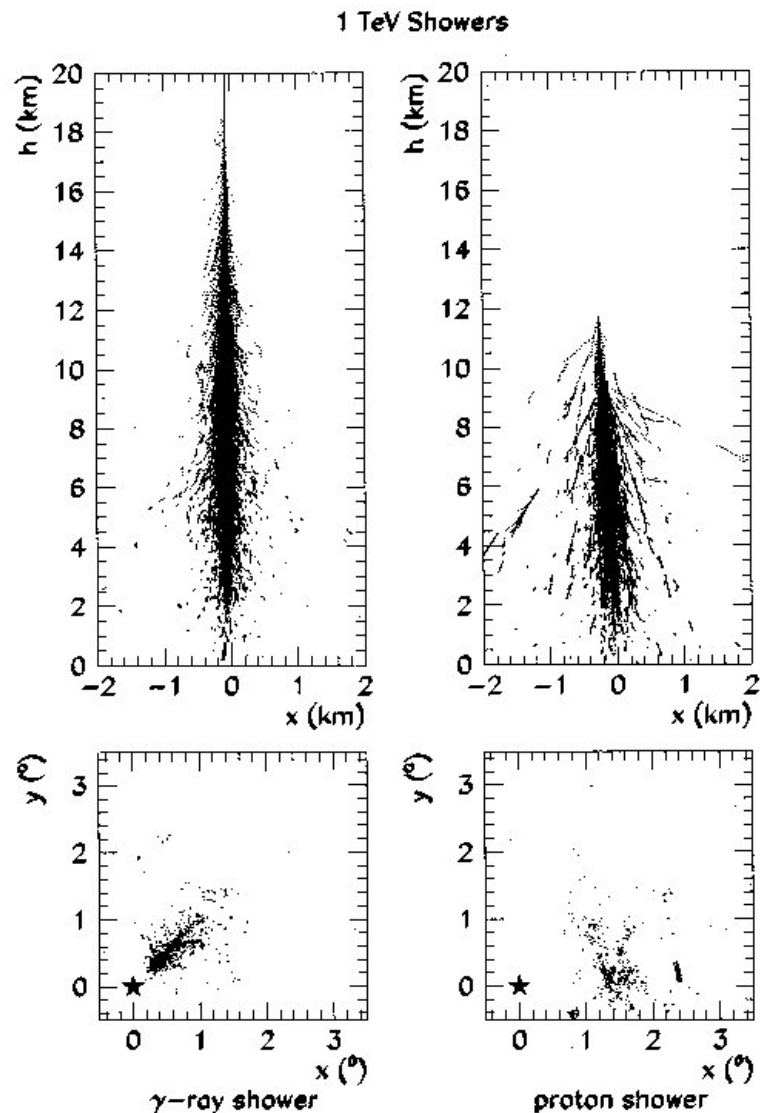
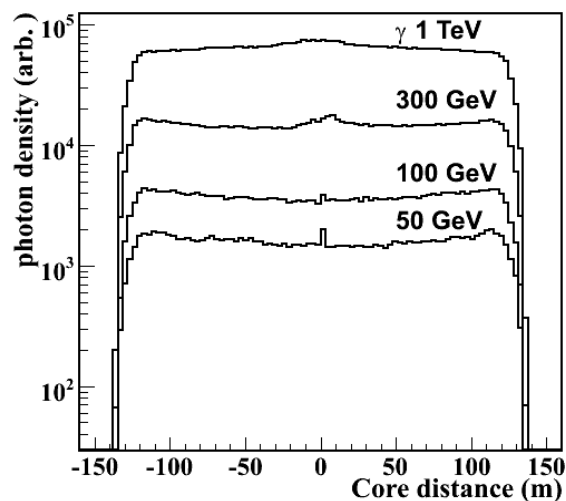
Image shape, orientation, and intensity
depend on the shower parameters.



Shower Development

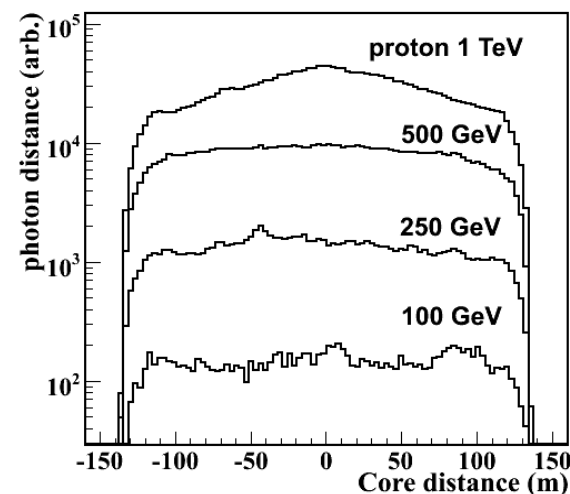
Gamma Ray

- › uniform shower development
- › higher Cerenkov light intensity
- › relatively symmetric
- › compact images without large fluctuations



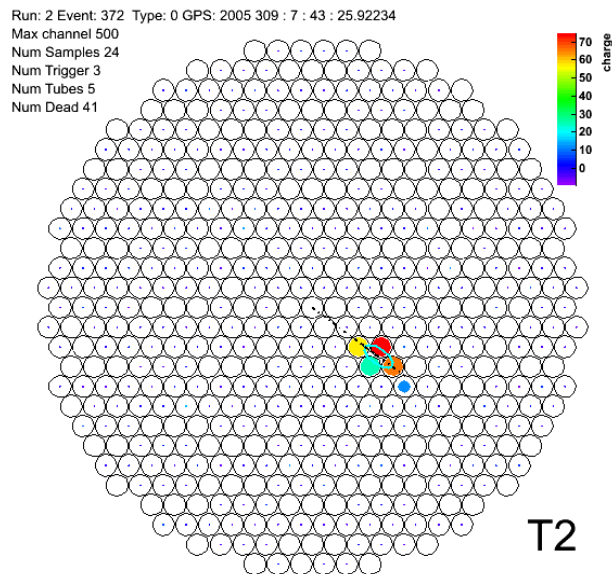
Proton

- › large fluctuations in shower development
- › lower Cerenkov intensity
- › less symmetric than electromagnetic shower
- › Laterally broader with fluctuations

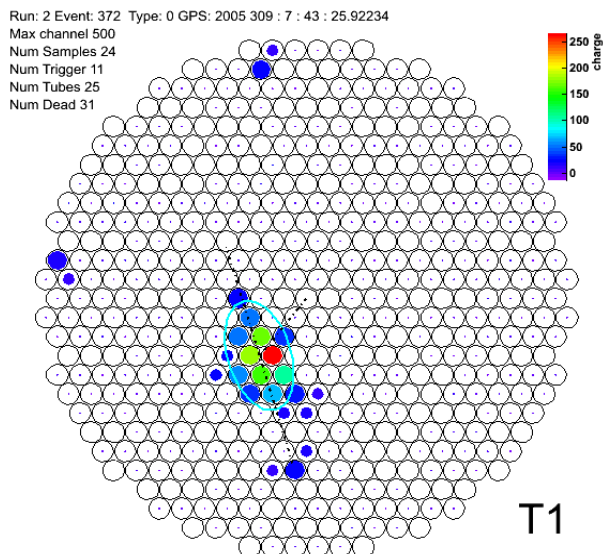




Imaging Air Showers: Mono



GEO: c_x=0.43, c_y=-0.32, dist=0.53, length=0.1077, width=0.0481, α =0.10, size=232.19



GEO: c_x=-0.31, c_y=-0.38, dist=0.49, length=0.3774, width=0.2067, α =55.50, size=1412.46

Single telescope allows

Good constraint of primary direction, but determined statistically

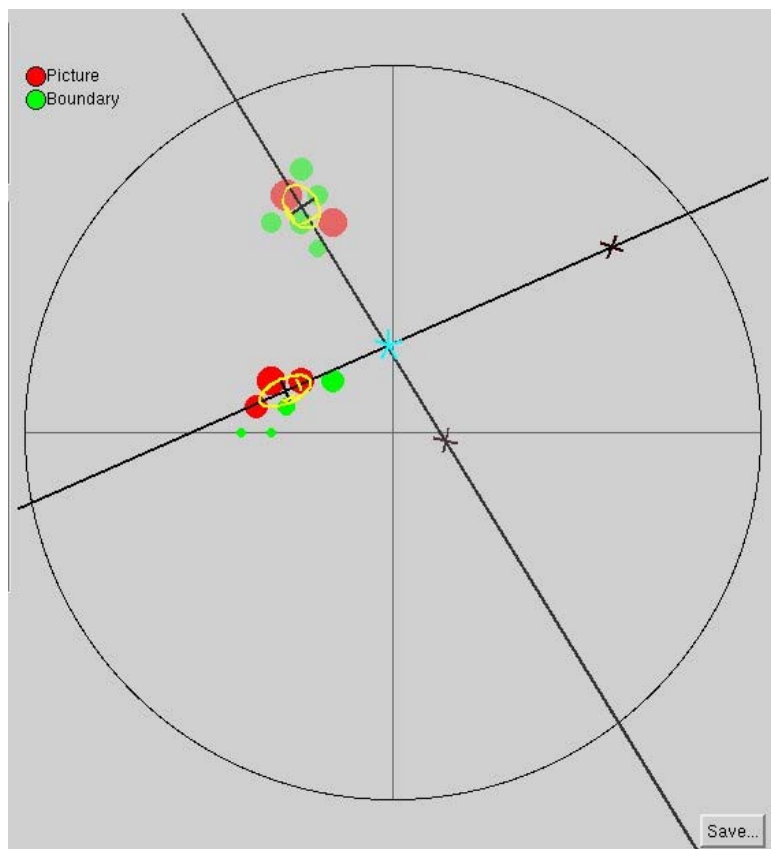
Best fits for elliptical showers, $R \sim 50-100$ m, higher energy

Some knowledge of core location

Energy resolution $\sim 30\%$



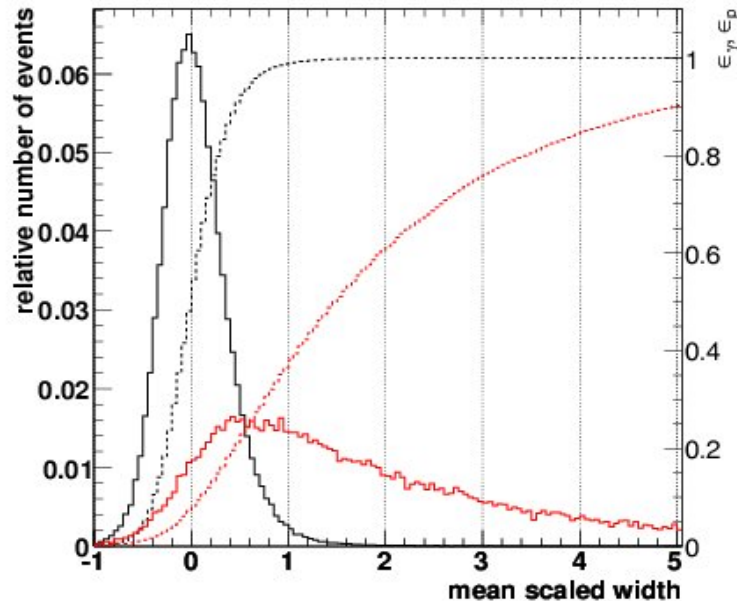
Imaging Air Showers: Stereo



- ▶ Array trigger eliminates local muon background
- ▶ Core Resolution $\sim 10\text{m}$
 - improves data quality cuts
 - hadron rejection
 - energy resolution 15-20%
- ▶ Multiple images not 100% correlated
 - array provides several images favorable for determining shower parameters
 - Fits improve substantially for 3 telescopes in event
- ▶ Primary direction determined per event
 - Reconstruction resolution $\sim 0.1^\circ$
- ▶ Primary energy, type, interaction height and shower max statistical



Background Rejection



Orientation still a powerful tool for point source analysis.

What about extended or diffuse sources?

► Rejection in trigger and image cleaning

- Hadronic showers of same energy produce less Cerenkov light than electromagnetic showers (loss to μ and ν)

► State of the art rejection still based on shape

- Knowledge of the shower core improves things substantially

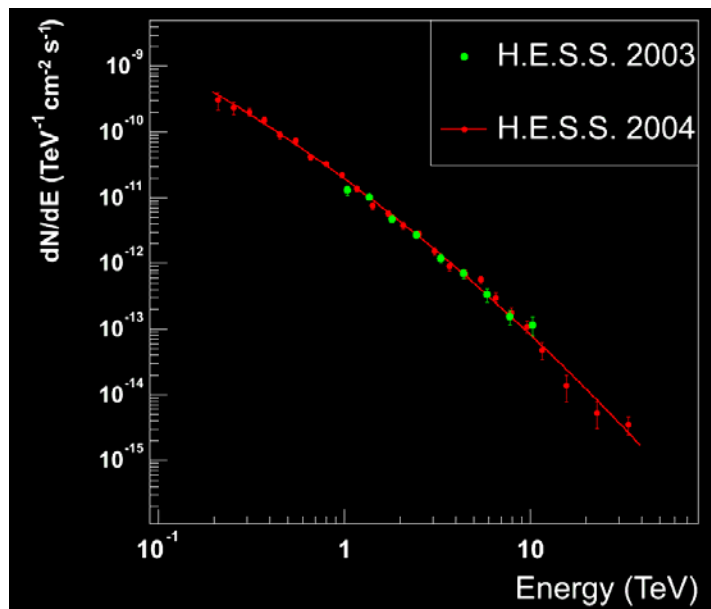
► Additional rejection potential

- Shower symmetries (widths, lateral density profile...)
- Timing



Energy Parameterization

$$N_c \propto \ln(E)$$



Differential flux for Supernova remnant RX J1713-394 from H.E.S.S.

But...this is only approximate. The atmosphere as a calorimeter is inhomogeneous and the Cerenkov angle varies with height.

And there are fluctuations in interaction height for gamma rays of a given energy.

Shower core and number of telescopes fitting the shower have a big impact.



Conclusions

- ▶ VERITAS already capable of stereo observations and full array completed soon
- ▶ Stereoscopic imaging allows excellent determination of
 - primary direction
 - shower core
- ▶ Hadron rejection significantly improved
 - ongoing work for improved methods without use of direction
- ▶ Additional methods still to be exploited