Minimal Dark Matter at Colliders

Satoshi Shirai (Kavli IPMU)

based on

arXiv:1504.03402 with

Keisuke Harigaya, Koji Ichikawa, Anirban Kund and Shigeki Matsumoto

and arXiv:1711.05449 with

Shigeki Matsumoto and Michihisa Takeuchi

Plan

1. Minimal Dark Matter

2. Indirect Search @Lepton Collider

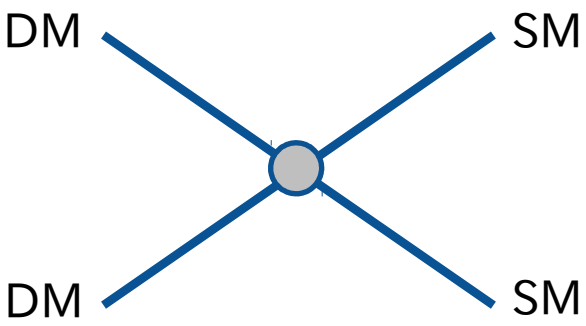
   Electroweak precision @LEP

   di-fermion and boson process @ILC250

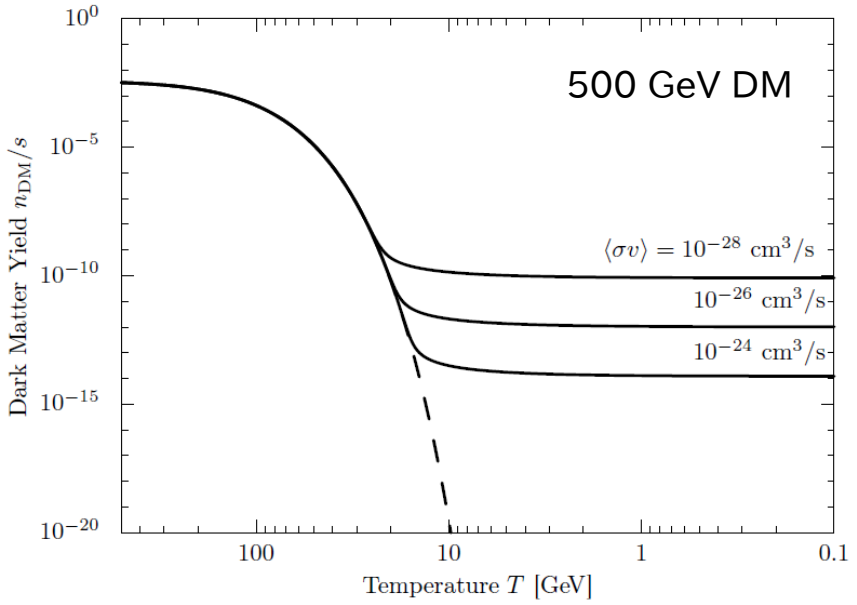
3. Indirect Search @LHC

WIMP Dark Matter

Weakly Interacting Massive Particle



DM abundance



Time

Concrete WIMP Model

WIMP: Weakly Interacting Massive Particle

SM gauge interaction

Minimal Dark Matter

Minimal Dark Matter (MDM)

- Well known interaction: concrete prediction
- Collider, direct detection, indirect detection
- Many new physics models include MDM
- Stability of DM
- …

Gauge interacting DM

DM's nature depends on SU(2)xU(1) charge

SU(2)

U(1)

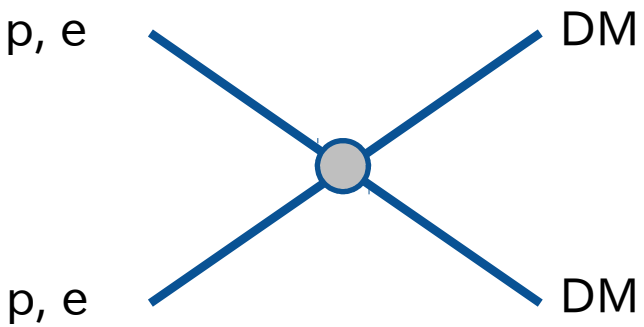
Y \ n	1	2	3	4	5
0		C	wino	C	5plet
1/2	C	higgsino	C		C
1	C	C		C	
3/2	C	C	C		C
2	C	C	C	C	
…	C	C	C	C	C

Example: Higgsino

Higgsino is

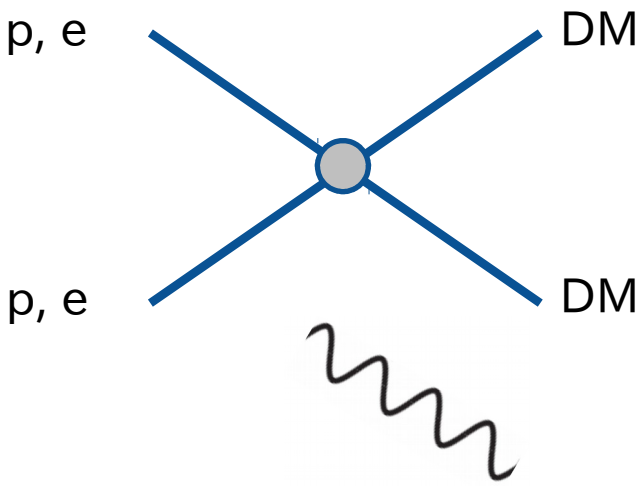
- (pseudo)Dirac fermion
- Hypercharge |Y|=1/2
- SU(2)doublet $\begin{pmatrix} \widetilde{H}_u^+ \\ \widetilde{H}_u^0 \end{pmatrix}, \begin{pmatrix} \widetilde{H}_d^0 \\ \widetilde{H}_d^- \end{pmatrix}$
- <1 TeV $\Omega h^2 \simeq 0.1 \left(\frac{m_{\widetilde{H}}}{1.1 \text{ TeV}} \right)^2$

Collider Signals of DM



DM is invisible

Collider Signals of DM

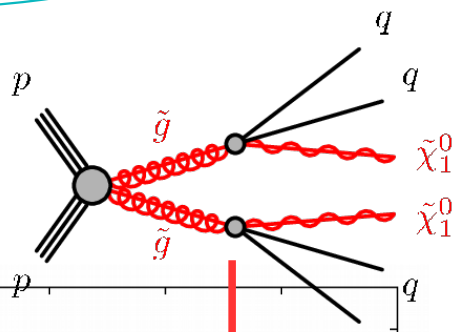


DM is invisible

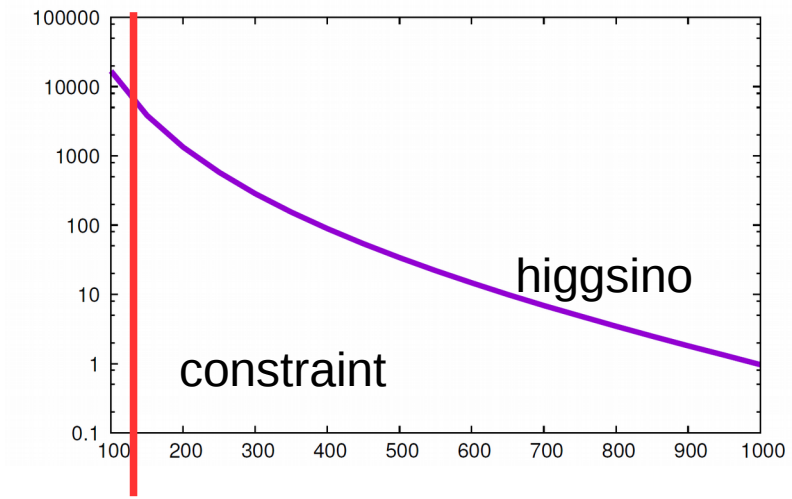
Additional objects are needed to see DM.
Missing energy (MET) search

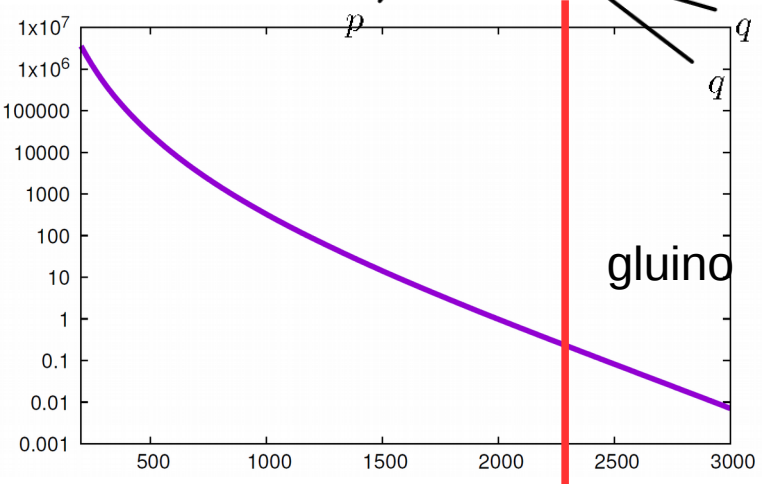
gluon photon, …

MDM is hard target



cross section [fb]





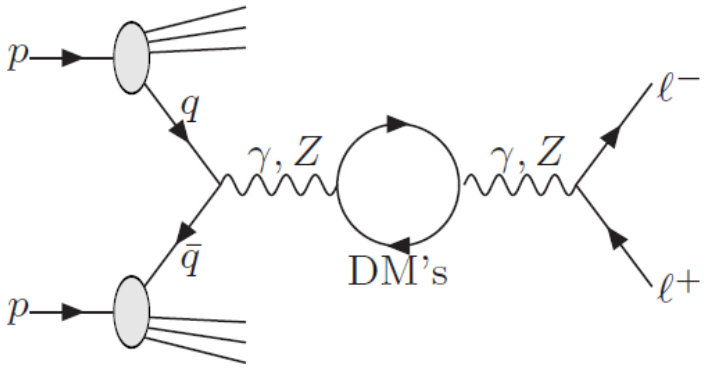
mass [GeV]

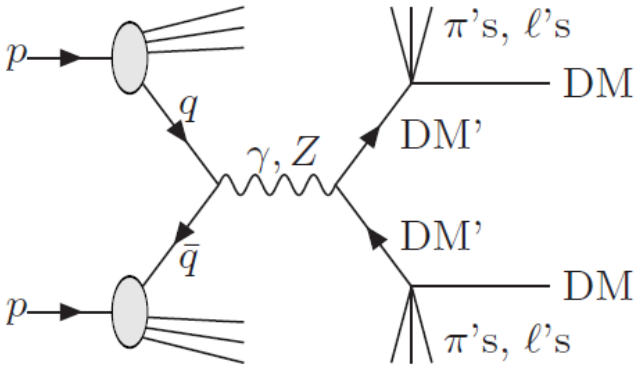
O(10^5) particle production

O(10) particle production

Efficient BG reduction and/or new observable are needed

MDM Search





Indirect Search from SM precision

Exotic tracks:

- Disappearing track
- Displaced soft track

Indirect Search @ Lepton Collider

Effective Operator

Case of beam energy < 2 x mass.
Integrating out of MDM leads effective operators

$$-\frac{1}{\Lambda_{2W}^2}(D_\mu W^a_{\mu\nu})^2 - \frac{1}{\Lambda_{2B}^2}(\partial_\mu B_{\mu\nu})^2 + \frac{g}{\Lambda_{3W}} W^3 \cdots$$

$$\frac{1}{\Lambda_{2W}^2} = \frac{g^2}{16\pi^2}\frac{1}{15m^2}\frac{n(n-1)(n+1)}{6}$$

$$\frac{1}{\Lambda_{2B}^2} = \frac{g'^2}{16\pi^2}\frac{2nY^2}{15m^2}$$

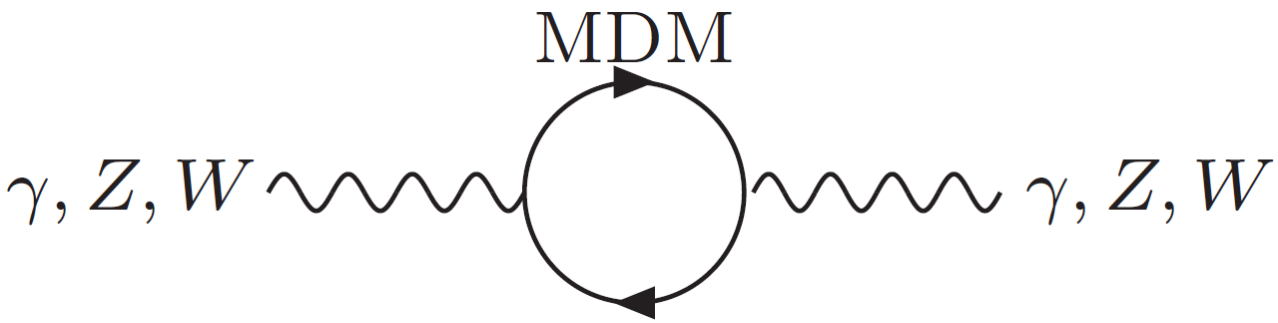
$$\frac{1}{\Lambda_{3W}^2} = \frac{-1}{12\Lambda_{2W}^2}$$

(Dirac Fermion)

Correction to gauge boson propagator and self-coupling

On-Z Observables

Heavy DM cannot be produced at collider, but affects SM processes

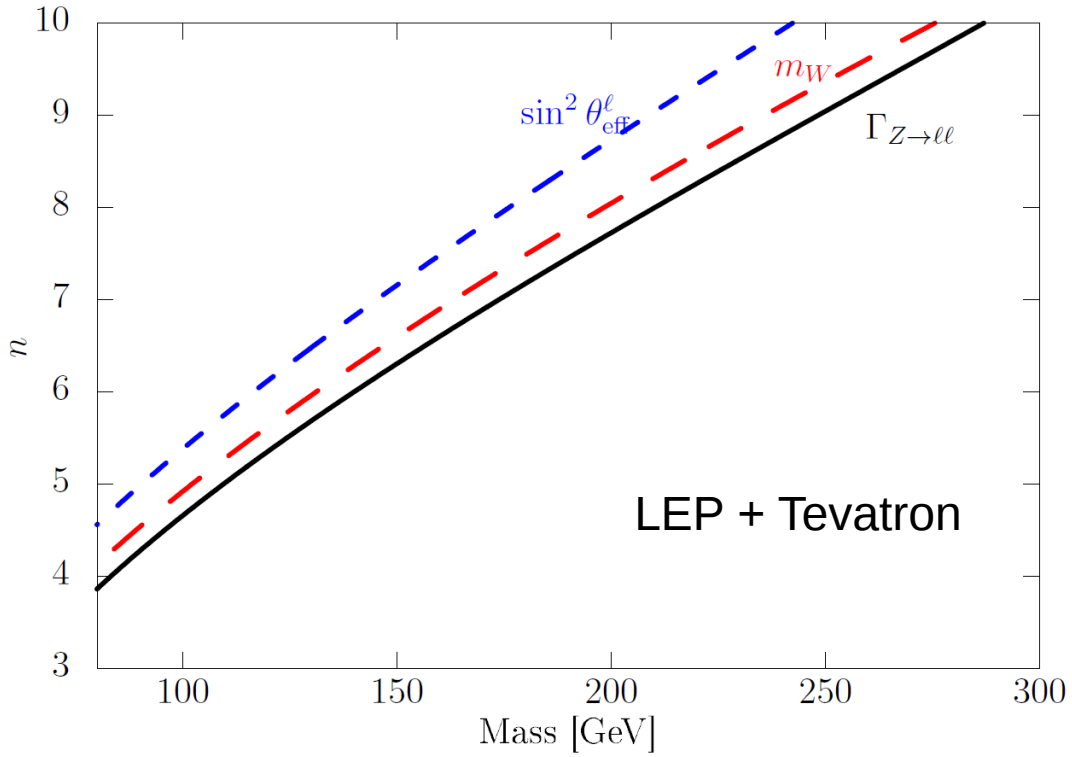


Case 1: Oblique correction to electoweak precision observables(EWPO)
weak boson mass and so on

Current Constraints from EWPO

of SU(2) representation

Majorana Fermion



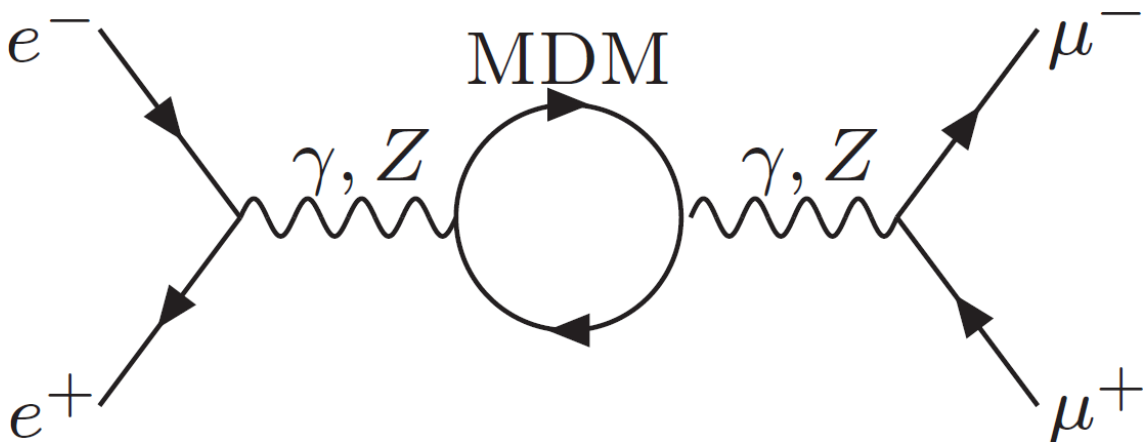
Corrections are roughly proportional to n^3 / m^2

[Harigaya+, 1504.03402]

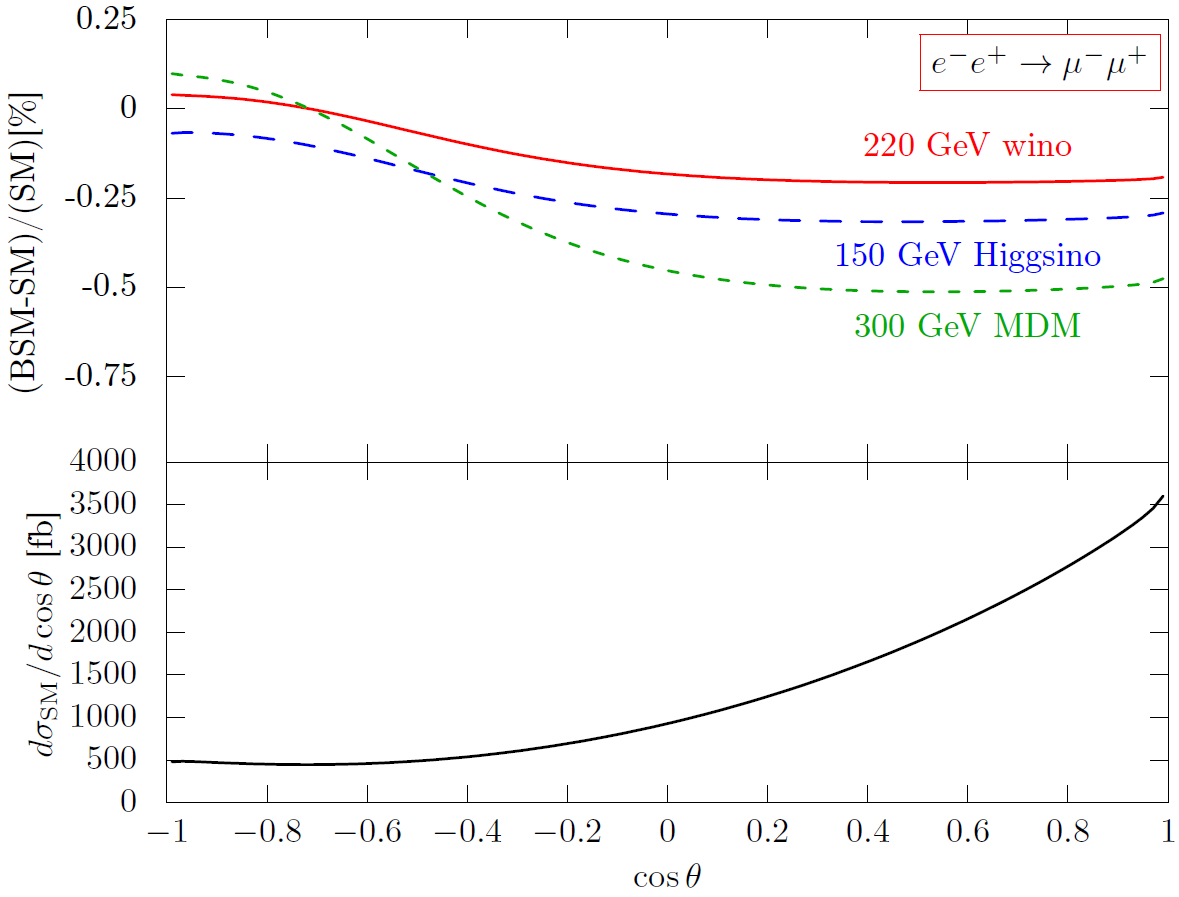
ILC250 for MDM

See Suehara's talk

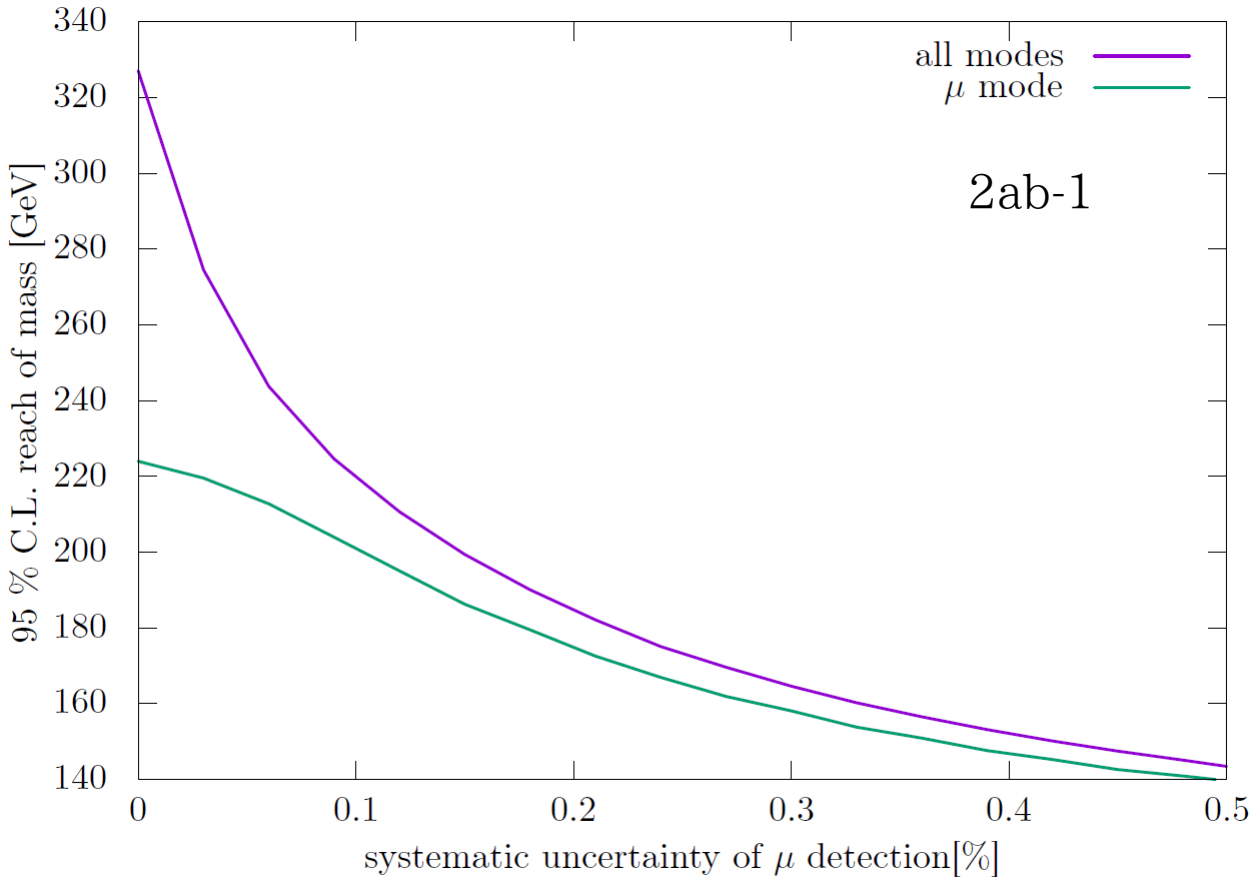
Case 2: Correction to di-fermion process at ILC.



ILC250 for Heavy MDM

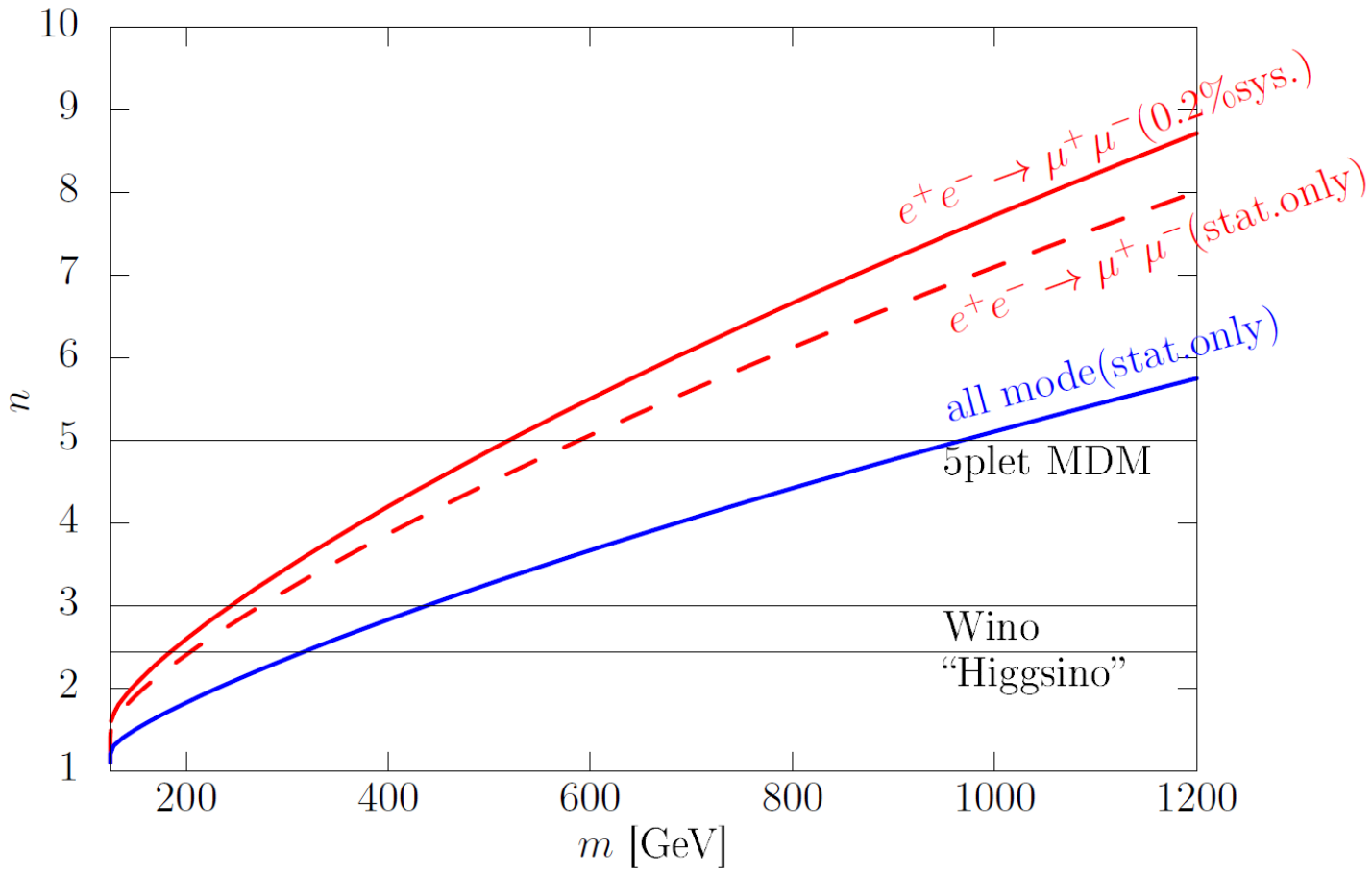


Reach of ILC250 for Higgsino



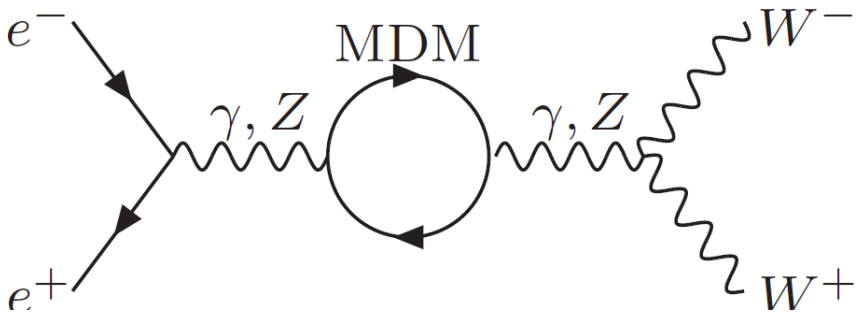
mu, e, b, c sys. Ratios: 2 : 1.5 : 5 : 10 assumed

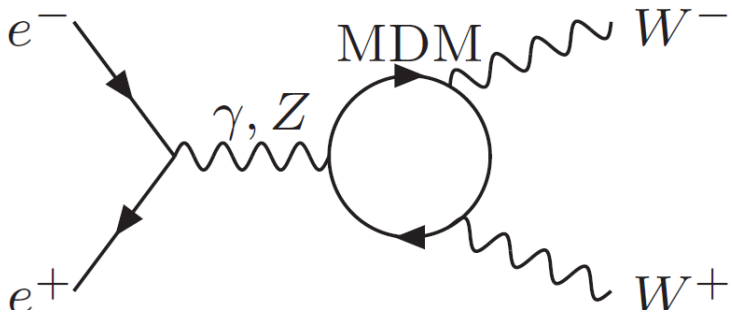
Reach of ILC250



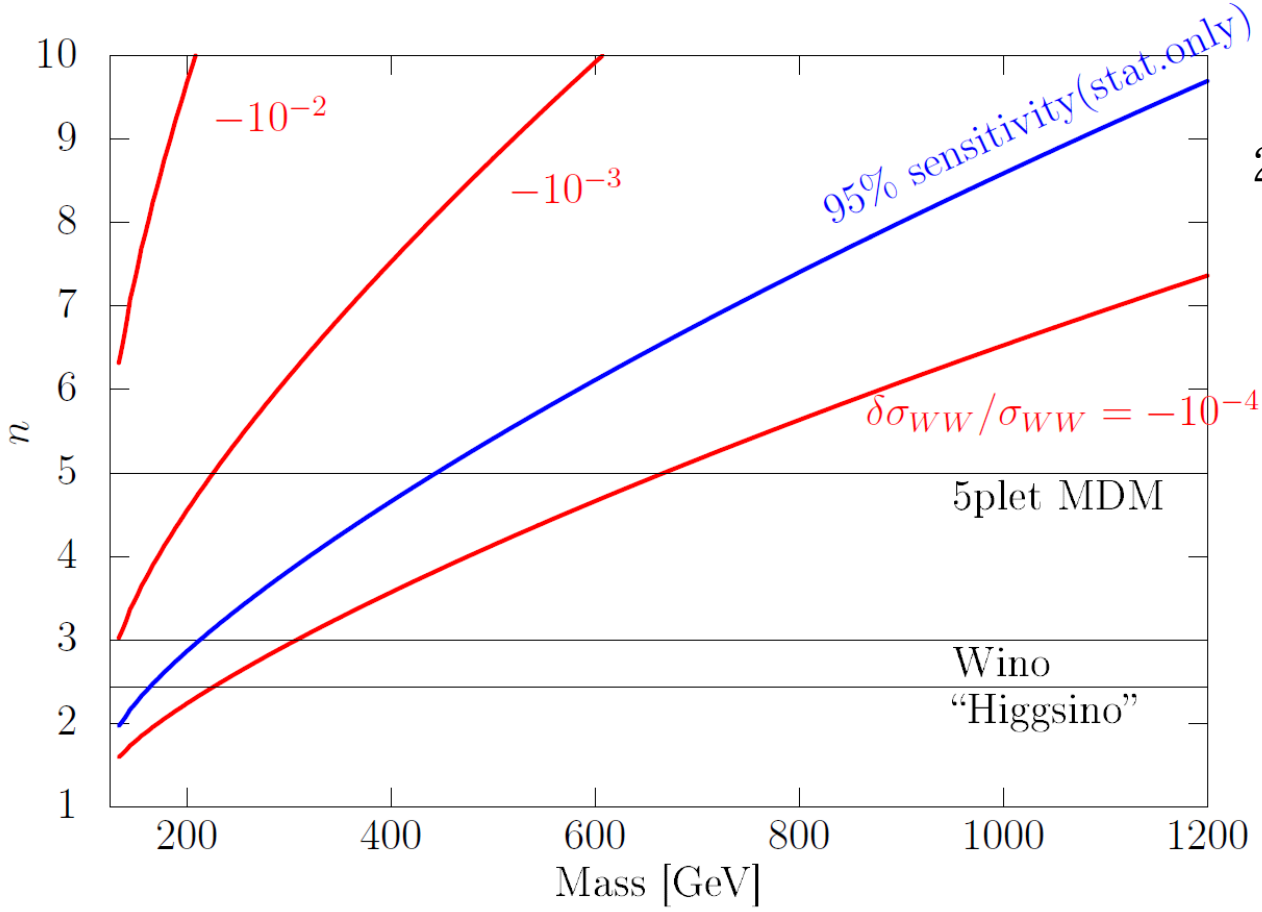
ILC250 for MDM

Case 3: Correction to diboson process at ILC.





ILC250 for MDM

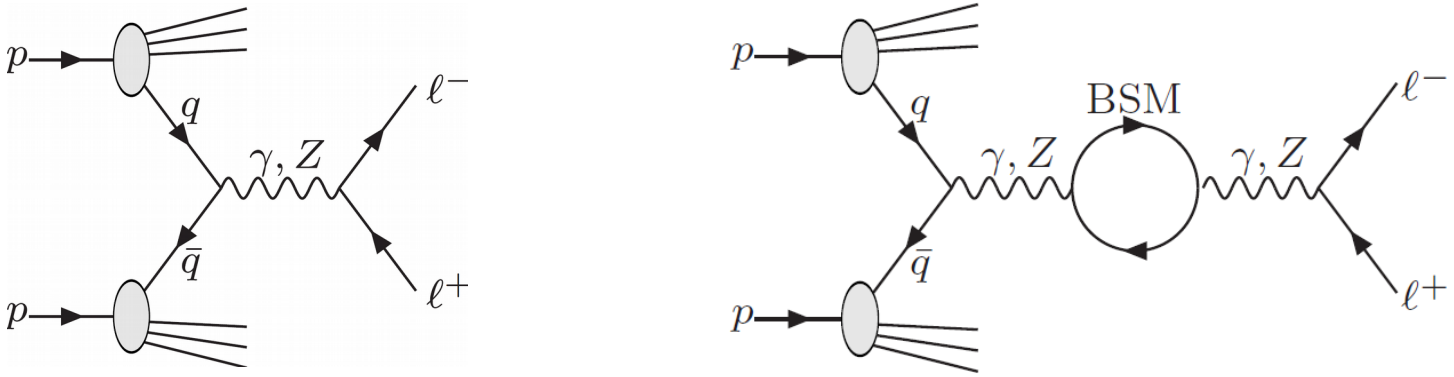


2ab-1

Indirect Search @ LHC

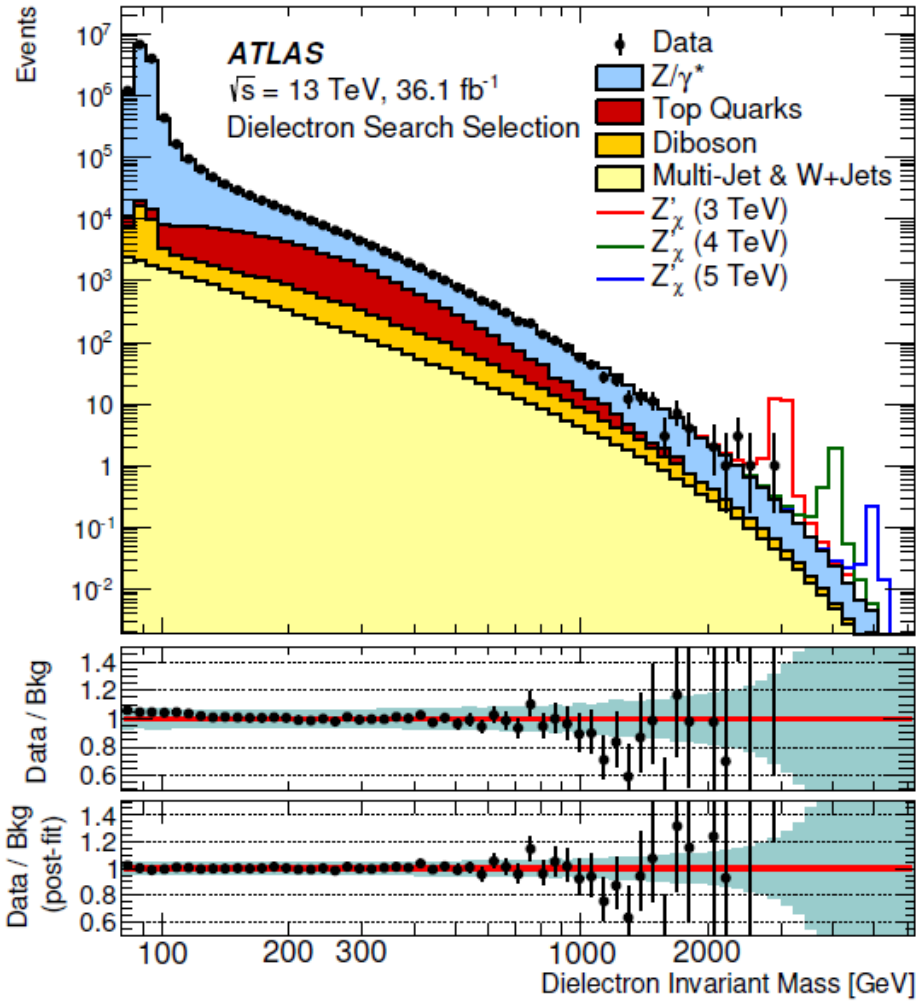
[Matsumoto, SS, Takeuchi, 1711.05449]

Indirect Probe at LHC



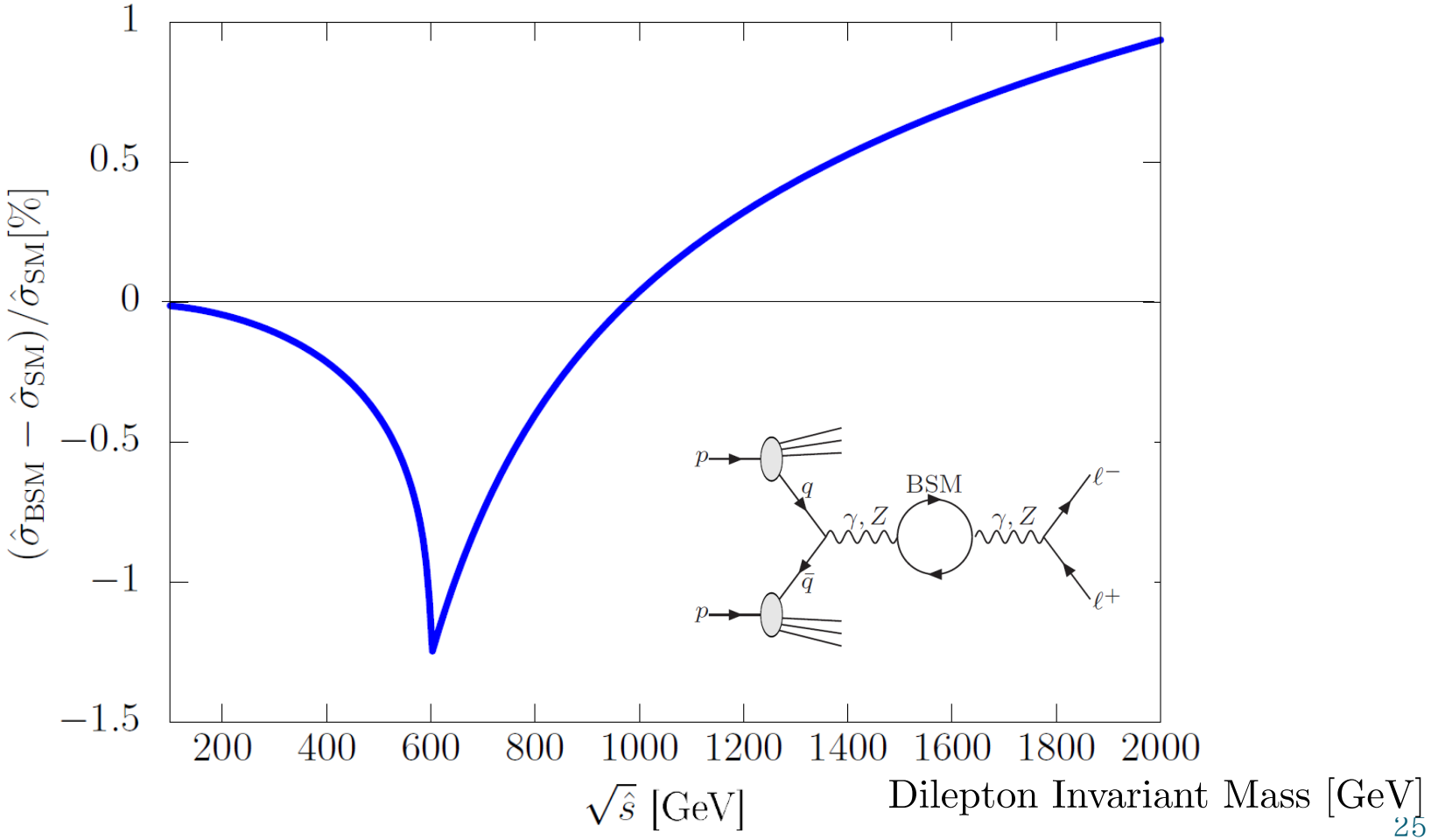
Interference between SM and BSM gives correction

Observed Data



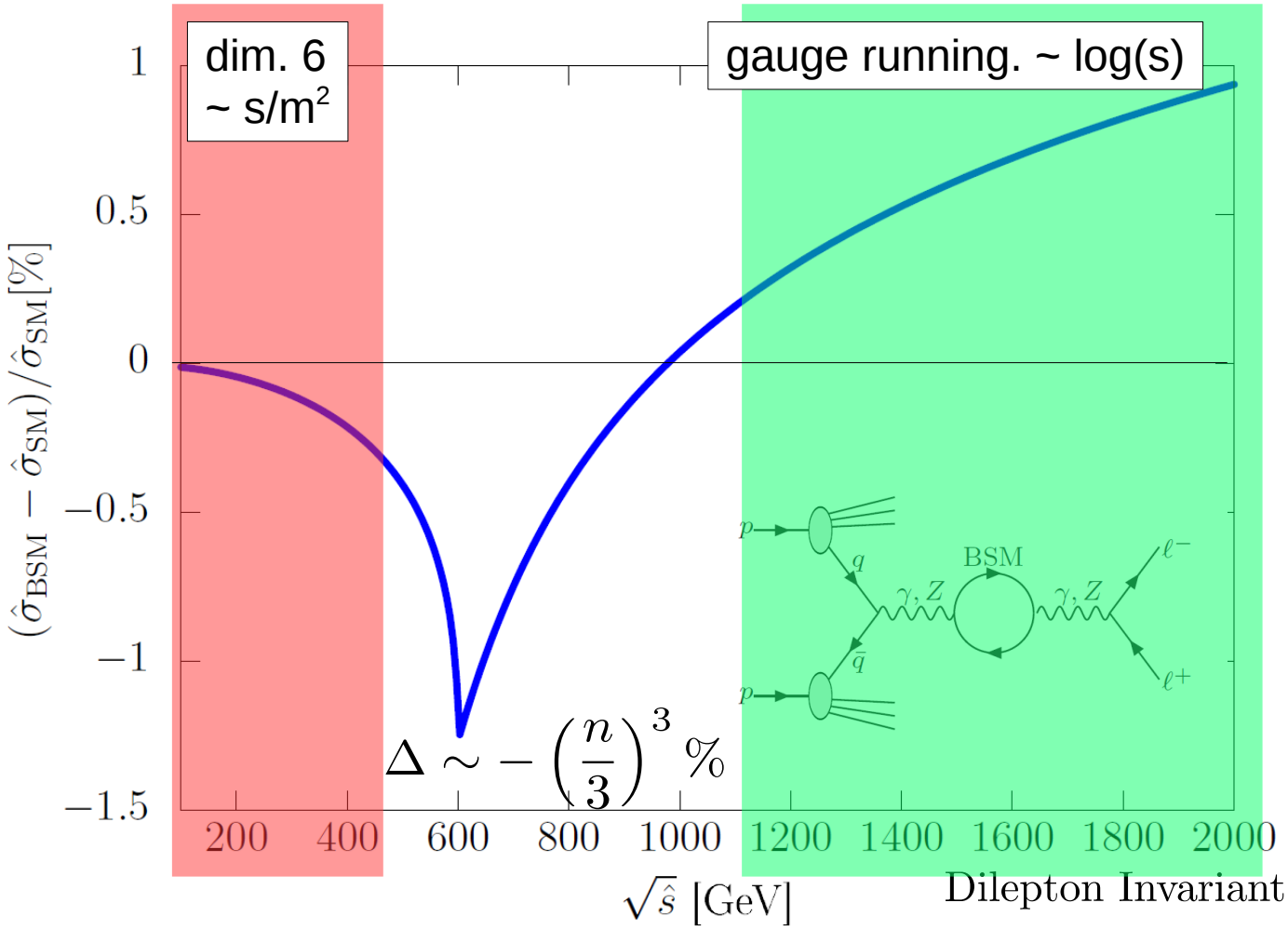
Correction from MDM

Relative difference from 300 GeV Wino

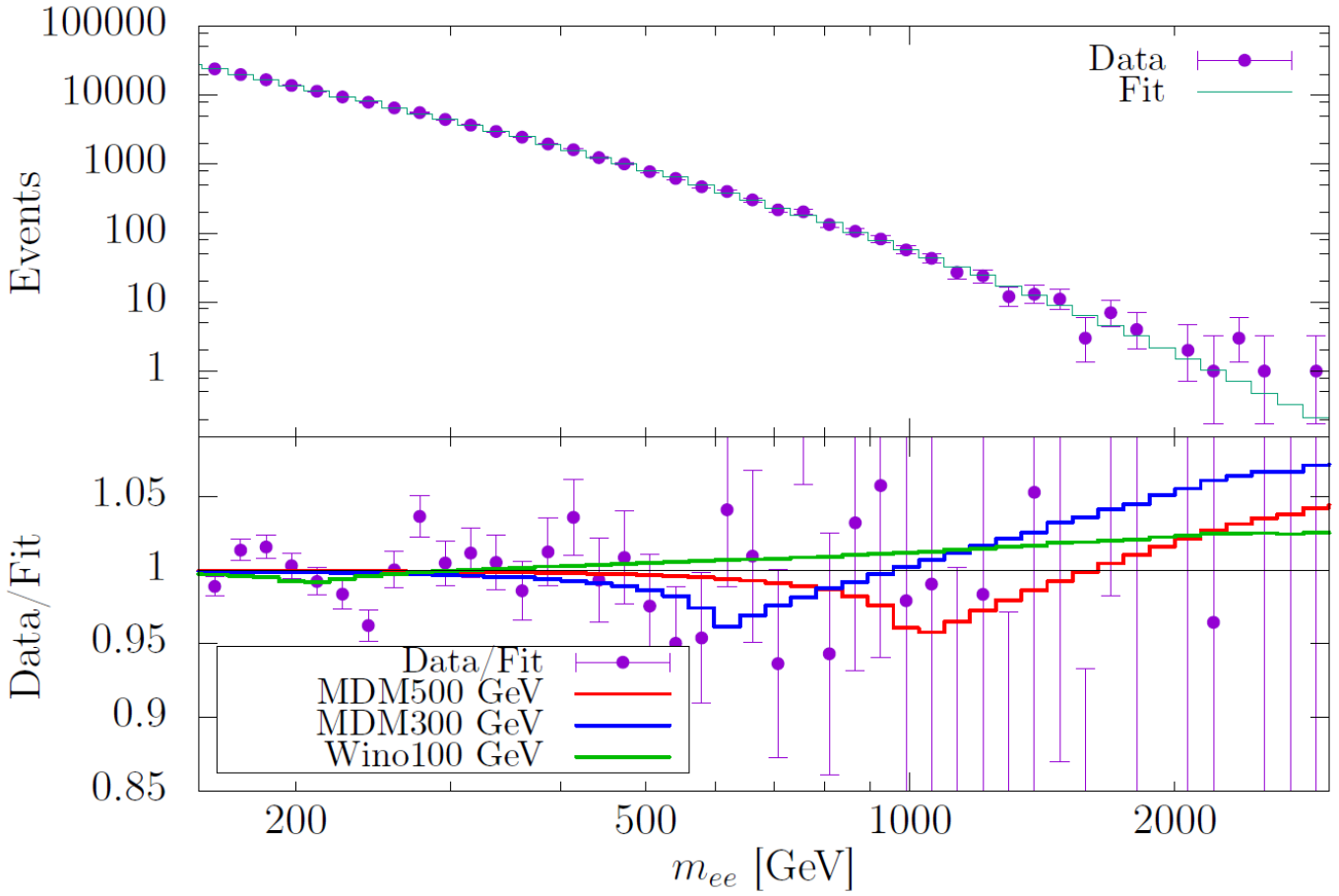


Correction from MDM

Relative correction from 300 GeV Wino

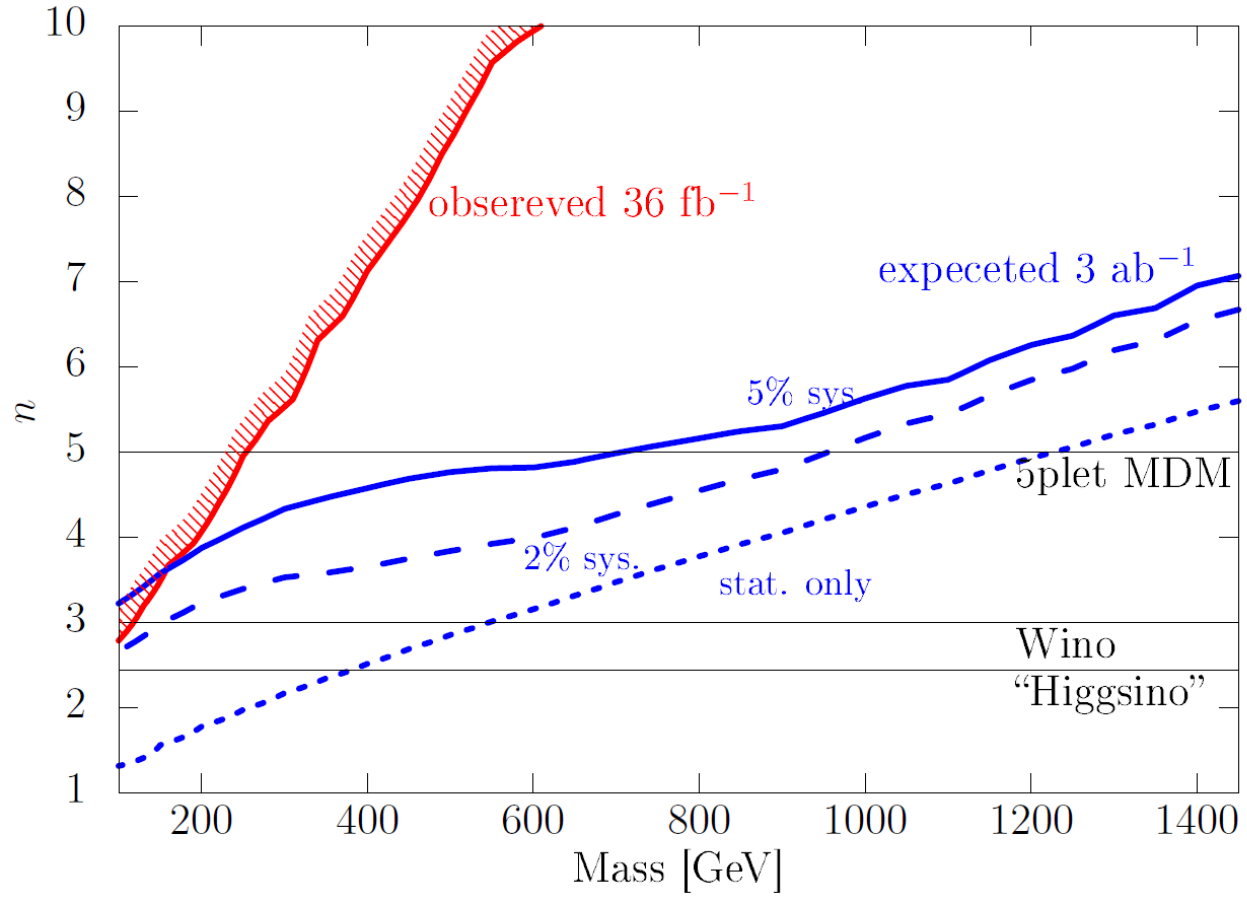


Indirect Probe at LHC



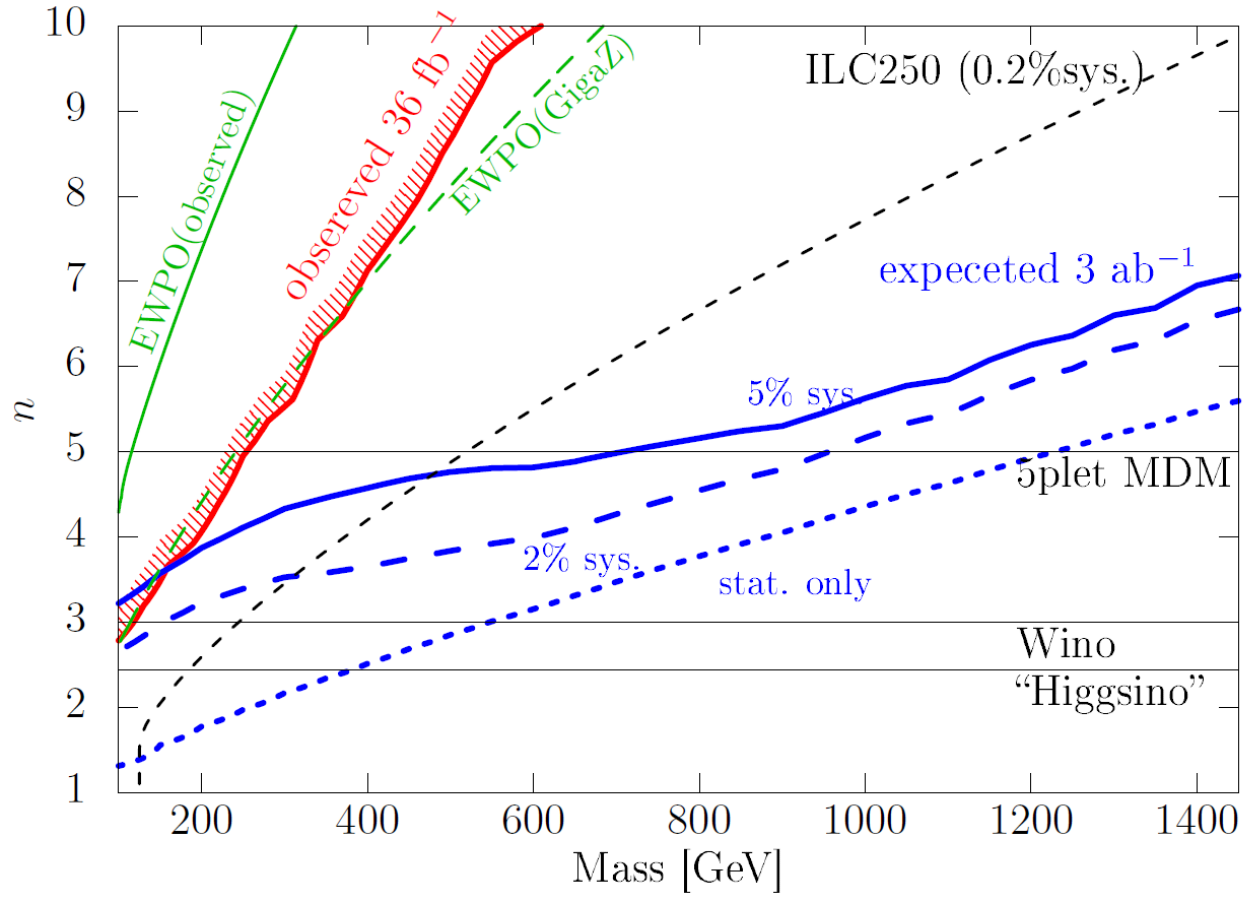
Indirect Probe at LHC

of SU(2) representation



Indirect Probe Summary

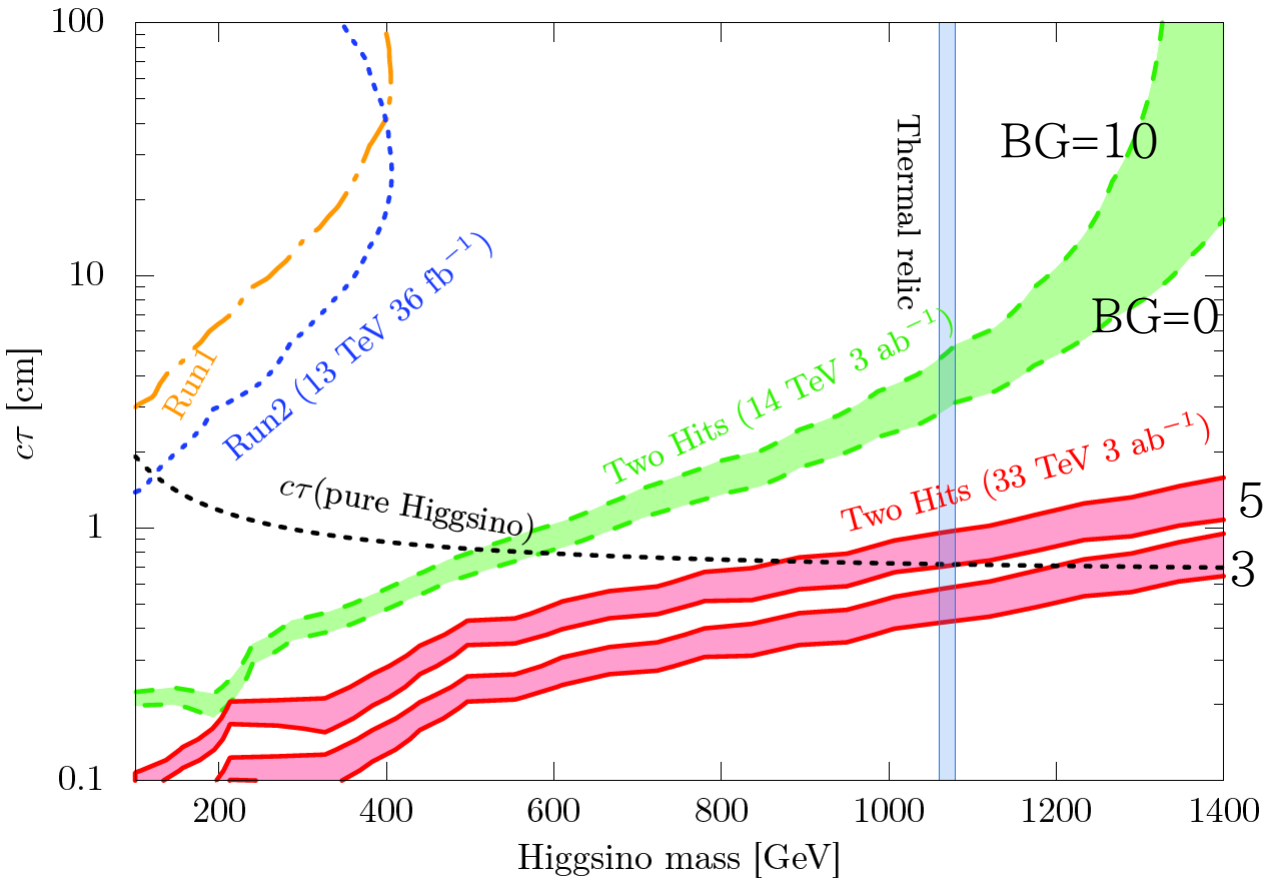
of SU(2) representation



Summary

- MDM can be indirectly tested at LHC and ILC
- As powerful as mono-jet + MET search
- This method uses only EW gauge interaction
  - Any EW interacting particle

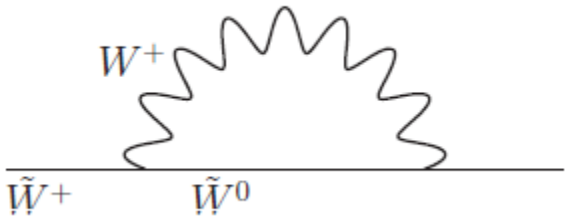
Prospects for Higgsino



MET>400 GeV

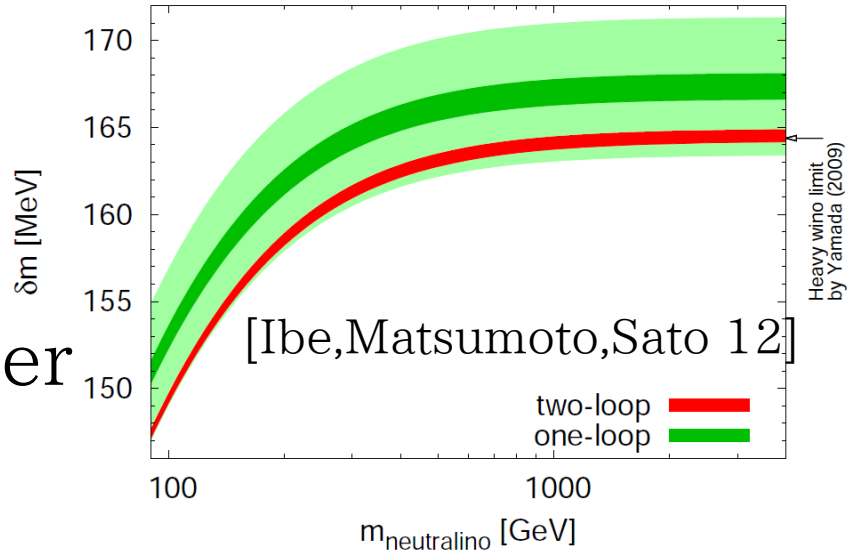
MET>600 GeV

Wino Spectrum



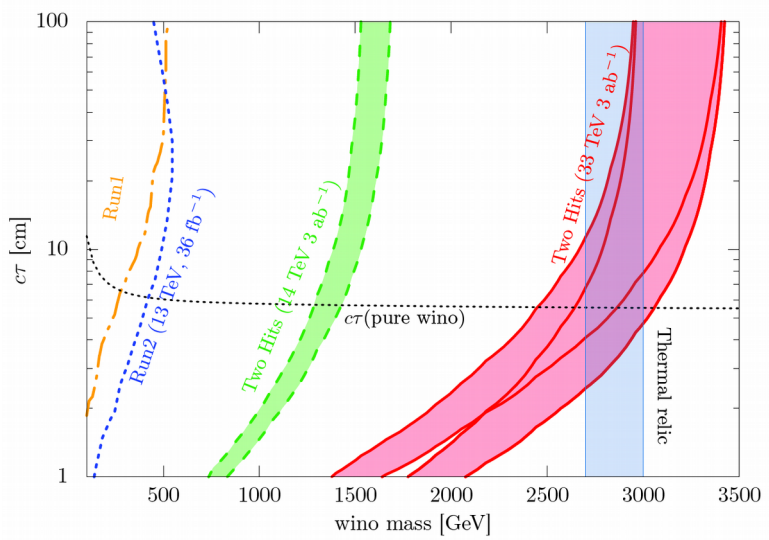
Radiative correction

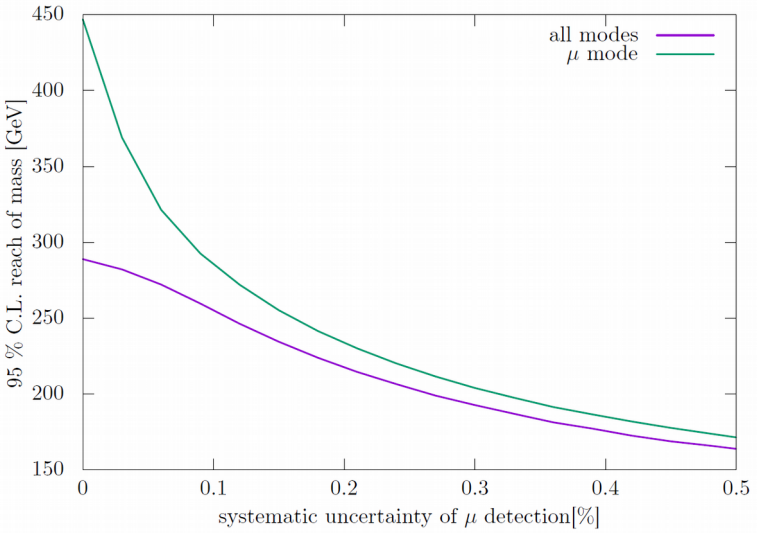
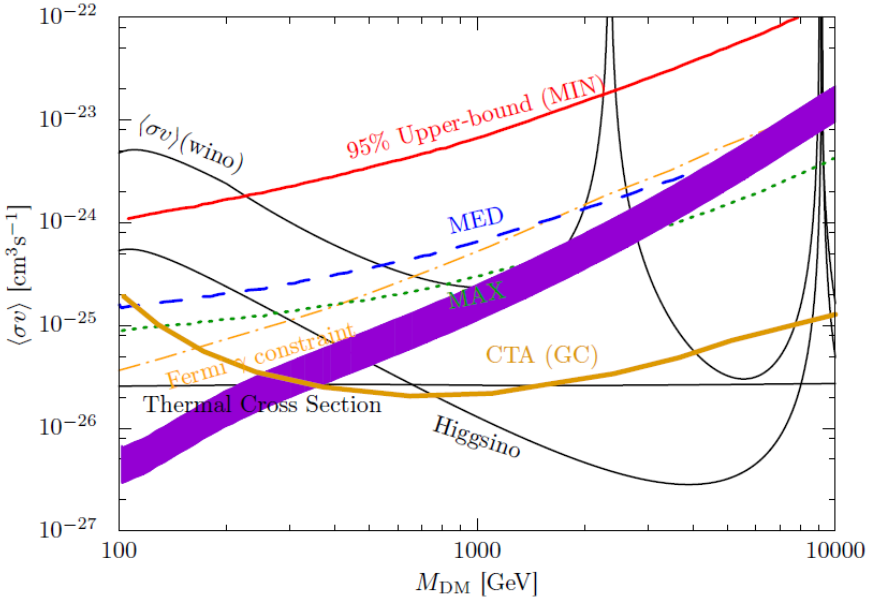
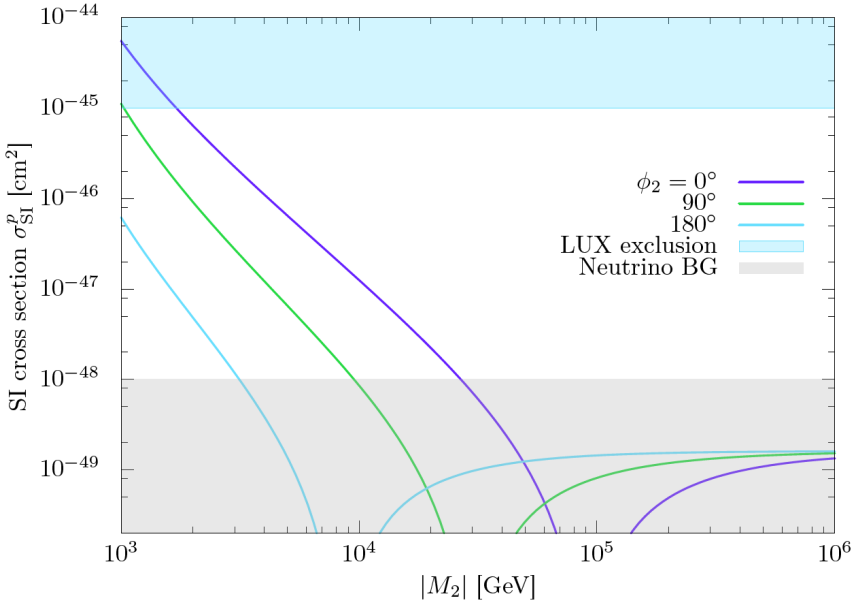
$$\begin{pmatrix} \widetilde{W}^+ \\ \widetilde{W}^0 \\ \widetilde{W}^- \end{pmatrix}$$ Charged slightly heavier



$$c\tau(\widetilde{W}^{\pm} \to \widetilde{W}^0 \pi^{\pm}) \simeq 7 \text{ cm} \left(\frac{\Delta m}{165 \text{ MeV}} \right)^{-3}$$

Wino Case


100
10
1
$c\tau$ [cm]
500
1000
1500
2000
2500
3000
3500
wino mass [GeV]
Run1
Run2 (13 TeV, 36 fb^{-1})
Two Hits (14 TeV 3 ab^{-1})
$c\tau$(pure wino)
Two Hits (33 TeV 3 ab^{-1})
Thermal relic


450
400
350
300
250
200
150
95 % C.L. reach of mass [GeV]
0
0.1
0.2
0.3
0.4
0.5
systematic uncertainty of μ detection[%]
all modes
μ mode

Higgsino DM is tough