# Determination of anomalous VVH couplings at the ILC

- 1). An overview of the anomalous VVH study ZZH/ $\gamma$ ZH and WWH induced with dim-6 operators
- 2). An application of a Matrix Element method toward further improvement of the sensitivity

Tomohisa Ogawa Junping Tian Keisuke Fujii

# 1). An overview of the anomalous VVH study ZZH/γZH and WWH induced with dim-6 operators

2). An application of a Matrix Element method toward further improvement of the sensitivity

Effective Higgs Self Coupling Systematic Error Uncertainties for Higgs Self Coupling Systematic Error Uncertainties and other, BSM, couplings in  $O(e^{e^{-t}})$  (and other, BSM, couplings) in  $O(e^{-t})$  HHZ) General definition  $\mathcal{E}_{eff}^{\mathbf{ZZHH}}$  (and other,  $\frac{\mathbf{g}}{\Lambda^{1}}$ )  $\frac{\mathbf{g}}{\Lambda^{2}}$  in  $(\mathbf{g}, \mathbf{e}, \mathbf$ Nucl. Phys. B268 (1986) 621-653. dim-5 ( $L^{\dagger}\Phi\Phi^{\dagger}L$ ) gives possible to describe dynamics below  $\Lambda$ , majonara neutrino mass can reflect symmetries of an underlying theory.  $\sigma(e^+e^- \to HHZ)$  can be described by by introducing general operators based on the gauge symmetry. We assume that  $\sigma(e^+e^- \to HHZ)$  can be described by an effective field theory (EFT) containing We assume that  $\sigma(e^+e^- \to HHZ)$  can be described by an effective field theory (EFT) containing a general Se (2) x e (r) gauge invariant Lagrangian with dimension to operators in addition to the SM. a general Se (2) x e (r) gauge invariant Lagrangian with dimension to operators in addition to the SM. Warsaw bases

General structures

10 CP-conserving dim-6 operators relevant to this an before spring the "Warsaw" basis, with the pure Higgs operators in the "SILH" basis, these are the

10 GB-conserving dimensorperators relativistically Higgs shift Combination  $w/\Phi$   $\Delta \mathcal{L} = \frac{c_H}{2v^2} \partial^{\mu}(\Phi^{\dagger}\Phi) \partial_{\mu}(\Phi^{\dagger}\Phi) + \frac{c_T}{2v^2} (\Phi^{\dagger} \stackrel{\rightarrow}{D}^{\mu}\Phi) (\Phi^{\dagger} \stackrel{\rightarrow}{D}_{\mu} \Phi) - \frac{c_6 \lambda}{2v^2} (\Phi^{\dagger}\Phi)^3$ T. Barklow et al... In addition there are 4 GP-violating  $\sigma^2 \tilde{c}_{\mu\nu\nu} = \sigma^2 \tilde{c}_{\mu\nu\nu}$ 

T. Barklow et al., Phys. Rev. D 97 (2018) 053003. There are 4  $\frac{1}{m_W^2}$   $\Phi^{\dagger}\Phi W_{\mu\nu}^a W^{a\mu\nu} + \frac{1}{m_W^2}$   $\Phi^{\dagger}\Phi W_{\mu\nu}^a \widetilde{W}^{a\mu\nu} + \frac{1}{m_W^2}$   $\Phi^{\dagger}\Phi W_{\mu\nu}^a \widetilde{W}^{a\nu\nu} + \frac{1}{m_$ 

Combination w/ V, Φ

## **Effective Field Theory**

$$\mathcal{L}_{eff} = \mathcal{L}_{SM}^{(4)} + \sum_{i} \frac{c_i^{(5)}}{\Lambda^1} \mathcal{O}_i^{(5)} + \sum_{i} \frac{c_i^{(6)}}{\Lambda^2} \mathcal{O}_i^{(6)} + \cdots$$

W. Buchmuller, D. Wyler, Nucl. Phys. B268 (1986) 621-653.

possible to describe dynamics below  $\Lambda$ ,

dim-5 ( $L^{\dagger}\Phi\Phi^{\dagger}L$ ) gives majonara neutrino mass

can reflect symmetries of an underlying theory.

by introducing general operators based on the gauge symmetry.

The number of relevant dim-6 operators @ ILC = 17

Warsaw bases

 $+\eta_W \frac{2m_W^2}{v_0}W_\mu^+W^{-\mu}h + \eta_{2W}\frac{m_W^2}{v_0^2}W_\mu^+W^{-\mu}h^2$ 

Grządkowski et al. arXiv:arXiv: 1008.4884,

General structures

After symmetry breaking the EWSB very successful to the symmetry breaking and the symmetry brea

$$\Delta \mathcal{L}_{h} = -\eta_{h} \lambda_{0} v_{0} h^{3} + \frac{\theta_{h}}{v_{0}} h \partial_{\mu} h \partial^{\mu} h + \eta_{Z} \frac{m_{Z}^{2}}{v_{0}} Z_{\mu} Z^{\mu} h + \frac{1}{2} \eta_{2Z} \frac{m_{Z}^{2}}{v_{0}^{2}} Z_{\mu} Z^{\mu} h^{2}$$

T. Barklow et al.,

Phys. Rev. D 97 (2018) 053003.

→ Complete formula is given

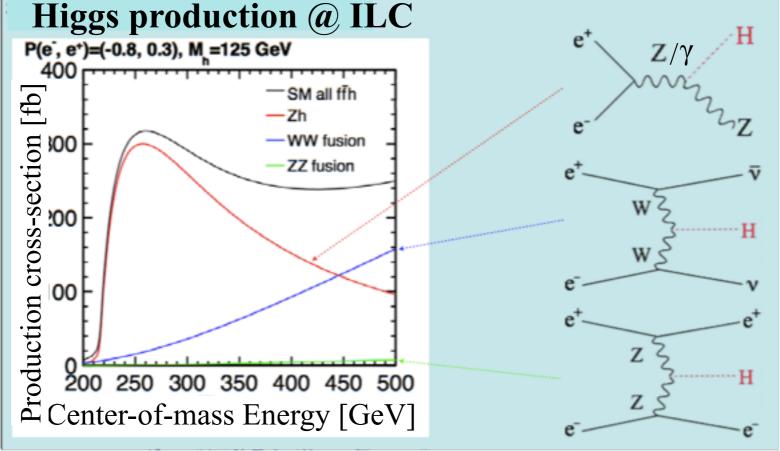
$$+\frac{1}{2}\left(\zeta_{Z}\frac{h}{v_{0}}+\frac{1}{2}\zeta_{2Z}\frac{h^{2}}{v_{0}^{2}}\right)\hat{Z}_{\mu\nu}\hat{Z}^{\mu\nu}+\left(\zeta_{W}\frac{h}{v_{0}}+\frac{1}{2}\zeta_{2W}\frac{h^{2}}{v_{0}^{2}}\right)\hat{W}_{\mu\nu}^{+}\hat{W}^{-\mu\nu}\\+\frac{1}{2}\left(\zeta_{A}\frac{h}{v_{0}}+\frac{1}{2}\zeta_{2A}\frac{h^{2}}{v_{0}^{2}}\right)\hat{A}_{\mu\nu}\hat{A}^{\mu\nu}+\left(\zeta_{AZ}\frac{h}{v_{0}}+\zeta_{2AZ}\frac{h^{2}}{v_{0}^{2}}\right)\hat{A}_{\mu\nu}\hat{Z}^{\mu\nu}.$$

$$\Delta \mathcal{L}_{eehZ} = g_{LZh}^{+\frac{1}{2}} \underbrace{\left(\tilde{\zeta}_{Z} \frac{h}{v_{\mu}} + \frac{1}{2} \tilde{\zeta}_{L}^{\frac{1}{2}} \frac{h^{2}}{v_{\nu}^{2}}\right)^{\frac{1}{2}} \underbrace{\frac{h^{2}}{v_{\nu}^{2}} \hat{Z}_{L}^{\mu\nu}}^{\frac{1}{2}} \underbrace{\left(\tilde{\zeta}_{R} \frac{h}{v_{\nu}^{2}} + \frac{1}{2} \tilde{\zeta}_{L}^{2} \frac{h^{2}}{v_{\nu}^{2}}\right)^{\frac{1}{2}} \underbrace{\frac{h^{2}}{v_{\nu}^{2}} \hat{Z}_{L}^{\mu\nu}}^{\frac{1}{2}} \underbrace{\left(\tilde{\zeta}_{R} \frac{h}{v_{\nu}^{2}} + \frac{1}{2} \tilde{\zeta}_{L}^{2} \frac{h^{2}}{v_{\nu}^{2}}\right)^{\frac{1}{2}} \underbrace{\left(\tilde{\zeta}_{R} \frac{h}{v_{\nu}^{2}} + \frac{1}{2} \tilde{\zeta}_{L}^{2} \frac{h^{2}}{v_{\nu$$

## **Effective Field**

General definition

The number o



7. Buchmuller, D. Wyler, ucl. Phys. B268 (1986) 621-653.

> dim-5 ( $L^{\dagger}\Phi\Phi^{\dagger}L$ ) gives majonara neutrino mass

Grządkowski et al. arXiv:arXiv: 1008.4884,

General structures

General structures

After symmetry breaking the EWSB very surjective and the symmetry breaking the EWSB we have,  $\Delta \mathcal{L} = \Delta \mathcal{L}_h + \Delta \mathcal{L}_{eehZ} + \Delta \mathcal{L}_{TGC}$  where

After EWSB we have,  $\Delta \mathcal{L} = \Delta \mathcal{L}_h + \Delta \mathcal{L}_{eehZ} + \Delta \mathcal{L}_{TGC}$  where

$$\Delta \mathcal{L}_{h} = -\eta_{h} \lambda_{0} v_{0} h^{3} + \frac{\theta_{h}}{v_{0}} h \partial_{\mu} h \partial^{\mu} h + \eta_{Z} \frac{m_{Z}^{2}}{v_{0}} Z_{\mu} Z^{\mu} h + \frac{1}{2} \eta_{2Z} \frac{m_{Z}^{2}}{v_{0}^{2}} Z_{\mu} Z^{\mu} h^{2}$$

T. Barklow et al., Phys. Rev. D 97 (2018) 053003.

→ Complete formula is given

$$+\frac{1}{2}\left(\zeta_{Z}\frac{h}{v_{0}}+\frac{1}{2}\zeta_{2Z}\frac{h^{2}}{v_{0}^{2}}\right)\hat{Z}_{\mu\nu}\hat{Z}^{\mu\nu}+\left(\zeta_{W}\frac{h}{v_{0}}+\frac{1}{2}\zeta_{2W}\frac{h^{2}}{v_{0}^{2}}\right)\hat{W}_{\mu\nu}^{+}\hat{W}^{-\mu\nu}\\+\frac{1}{2}\left(\zeta_{A}\frac{h}{v_{0}}+\frac{1}{2}\zeta_{2A}\frac{h^{2}}{v_{0}^{2}}\right)\hat{A}_{\mu\nu}\hat{A}^{\mu\nu}+\left(\zeta_{AZ}\frac{h}{v_{0}}+\zeta_{2AZ}\frac{h^{2}}{v_{0}^{2}}\right)\hat{A}_{\mu\nu}\hat{Z}^{\mu\nu}.$$

$$\Delta \mathcal{L}_{eehZ} = g_{LZh}^{+\frac{1}{2}} \underbrace{\left(\tilde{\zeta}_{Z} \frac{h}{\nu_{0}} + \frac{1}{2} \tilde{\zeta}_{Z} \frac{h^{2}}{v_{0}^{2}}\right)^{\frac{1}{2}} \underbrace{\frac{h^{2}}{2}}_{2} \hat{Z}_{\mu\nu}^{2} + \left(\tilde{\zeta}_{R} \frac{h}{\nu_{0}} + \frac{1}{2} \tilde{\zeta}_{Z} \frac{h^{2}}{v_{0}^{2}}\right)^{\frac{1}{2}} \underbrace{\frac{h^{2}}{2}}_{2} \hat{Z}_{\mu\nu}^{2} \hat{Z}_{\mu\nu}^{2} + \left(\tilde{\zeta}_{R} \frac{h}{\nu_{0}} + \frac{1}{2} \tilde{\zeta}_{Z} \frac{h^{2}}{2} \frac{h^{2}}{v_{0}^{2}}\right)^{\frac{1}{2}} \underbrace{\frac{h^{2}}{2}}_{2} \hat{Z}_{\mu\nu}^{2} \hat{Z}_{\mu\nu}^{2}} \underbrace{\frac{h^{2}}{2} \underbrace{\frac{h^{2}}{2}}_{2} \hat{Z}_{\mu\nu}^{2} \underbrace{\frac{h^{2}}{2}}_{2} \hat{Z}_{\mu\nu}^{2} \underbrace{\frac{h^{2}}{2}}_{2} \hat{Z}_{\mu\nu}^{2} \underbrace{\frac{h^{2}}{2}}_{2} \hat{Z}_{\mu\nu}^{2} \underbrace{\frac{h^{2}}{2}}_{2} \underbrace{\frac{h^{2}}{2}$$

 $+\eta_W \frac{2m_W^2}{v_2} W_{\mu}^+ W^{-\mu} h + \eta_{2W} \frac{m_W^2}{v_2^2} W_{\mu}^+ W^{-\mu} h^2$ 

## anomalous ZZH: 3 parameters fit

Notation on ZZH ⇒ az, bz, btz parameters assuming beam Pol. left/right

$$\mathcal{L}_{ZZH} = M_Z^2 \left(\frac{1}{\mathbf{v}} + \frac{a_Z}{\Lambda}\right) Z_{\mu} Z^{\mu} H + \frac{b_Z}{2\Lambda} \hat{Z}_{\mu\nu} \hat{Z}^{\mu\nu} H + \frac{\tilde{b}_Z}{2\Lambda} \hat{Z}_{\mu\nu} \hat{Z}^{\mu\nu} H + \frac{\tilde{b}_Z}{2\Lambda} \hat{Z}_{\mu\nu} \hat{Z}^{\mu\nu} H$$
(A=1TeV)

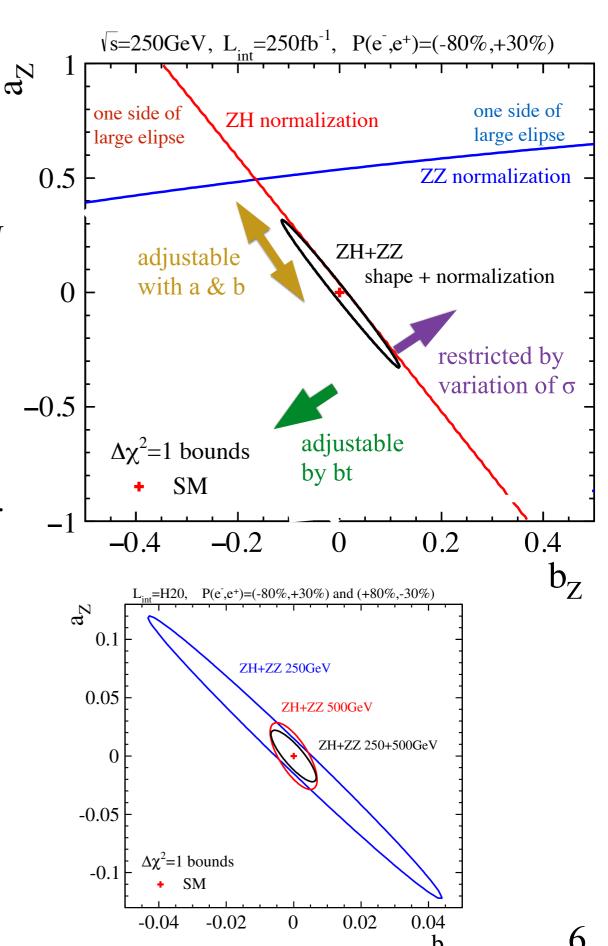
All SM bkgs are considered Detector response is considered.

The sensitivity can not be given with norm. only. The shape information is critical for the determination.

EPS17 talk
https://indico.cern.ch/event/466934/contributions/2588482/
Annual ILC physics and detector meeting
https://agenda.linearcollider.org/event/7837/contributions/
40946/attachments/32854/49991/annualMeeting18.pdf

Energy is also can improve the sensitivity

H20 operation (250GeV 2ab<sup>-1</sup>) including 500GeV H20 operation https://arxiv.org/abs/1506.07830



## anomalous ZZH/γZH: 3 parameters fit

H



$$\mathcal{L}_{ZZH} = M_Z^2 \left(\frac{1}{v} + \frac{a_Z}{\Lambda}\right) Z_\mu Z_\mu^\mu H$$

$$= 2 \text{ Rare processes at CLIC involvi} \text{ Profit on the problem of the entire of t$$

#### Five parameters fit

 $1\sigma$  bounds including 500GeV operation

ZZH / γZH structures can be measured ~2% or much better

$$ZH + ZZ$$
 at  $250 + 500$  GeV with  $H20$ 

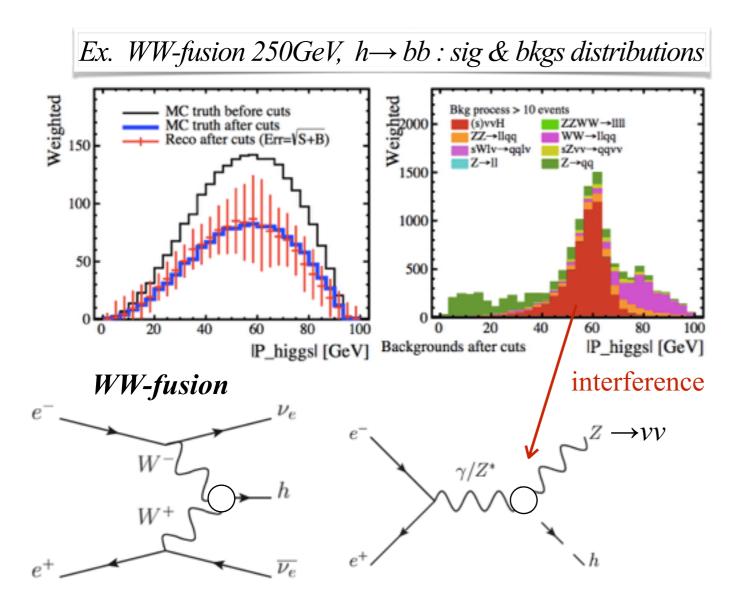
$$\begin{cases}
a_Z = \pm 0.0223 \\
\zeta_{ZZ} = \pm 0.0067 \\
\zeta_{AZ} = \pm 0.0024 , \rho = \begin{pmatrix}
1 & -.837 & -.134 & -.009 & -.010 \\
- & 1 & .040 & .008 & .013 \\
- & - & 1 & .006 & -.0012 \\
- & - & - & 1 & .600 \\
- & - & - & 1
\end{pmatrix}$$

$$\tilde{\zeta}_{AZ} = \pm 0.0006$$

## anomalous WWH: 3 parameters fit

Notation on ZZH ⇒ aw, bw, btw parameters

$$\mathcal{L}_{WWH} = 2M_W^2 \left(\frac{1}{\mathbf{v}} + \frac{a_W}{\Lambda}\right) W_\mu^+ W^{-\mu} H + \frac{b_W}{\Lambda} \hat{W}_{\mu\nu}^+ \hat{W}^{-\mu\nu} H + \frac{\tilde{b}_W}{\Lambda} \hat{W}_{\mu\nu}^+ \hat{W}^{-\mu\nu} \hat{W}^{-\mu\nu} H + \frac{\tilde{b}_W}{\Lambda} \hat{W}_{\mu\nu}^+ \hat{W}^{-\mu\nu} \hat{W}^{-\mu\nu} \hat{W}^{-\mu\nu} \hat{W}^{-\mu\nu} \hat{W}^{-\mu\nu} \hat{W}^{-\mu\nu} \hat$$



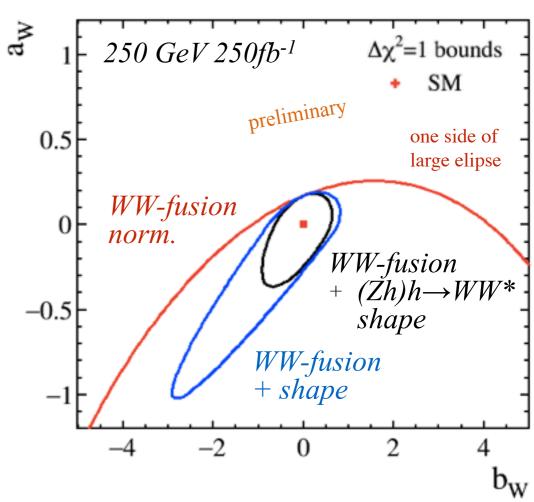
ZH w/ anomalous Same final state → contaminate WWH due to variation shape & norm.

#### LCWS17

https://agenda.linearcollider.org/event/7645/contributions/40062/attachments/32273/49230/LCWS17\_Ogawa\_v171025.pdf

#### Annual ILC physics and detector meeting

https://agenda.linearcollider.org/event/7837/contributions/40946/attachments/32854/49991/annualMeeting18.pdf



## anomalous WWH: 3 parameters fit

#### LCWS17

0.5

-0.5

https://agenda.linearcollider.org/event/7645/contributions/40062/

https://agenda.linearcollider.org/event/7837/contributions/40946,

preliminary

WW-fusion

+ shape

=1 bounds

one side of

large elipse

WW-fusion +  $(Zh)h \rightarrow WW*$ 

shape

SM

vs=250+500Gevaltachuments/32273/49230/LEWS 7 vs 250+500Gevaltachuments/32273/49230/LEWS 7 vs 250+500Gevaltachuments/32273/4920/LEWS 7 vs 250+500Gevaltachuments/3220/LEWS 7 vs 250+500Gevaltachuments/3220/LEWS 7 vs 250+500Gevaltachuments/3220/LEWS 7 vs 250+5000Gevaltachuments/3220/LEWS 7 vs 250+5000Gevaltachuments/3220/LEWS 7 vs 250+5000Gevaltachuments/3220/LEWS 7 vs 250+5000Gevaltachuments/

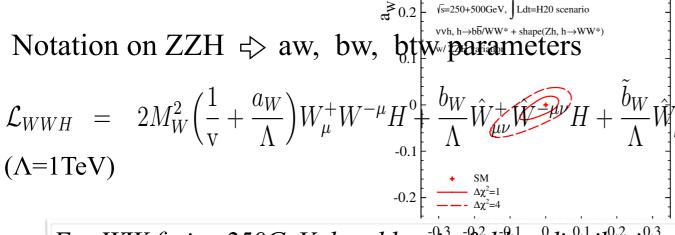
vvh, h o b \( \bar{b} \) / WW\* + shape (Zh, h \rightarrow W\*) physics and detector meeting

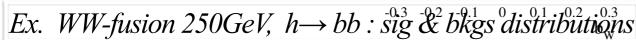
attachments/32854/49991/annualMeeting/8.pdf

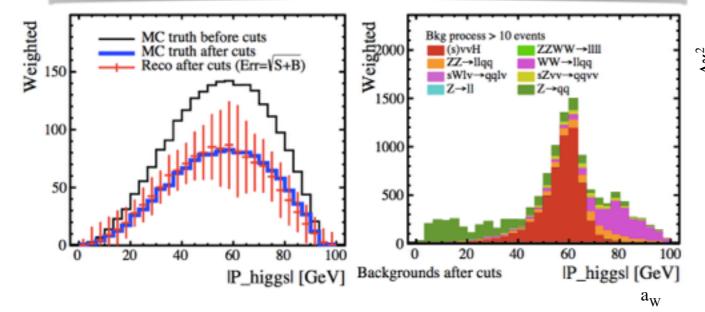
250 GeV 250fb<sup>-1</sup>

WW-fusion

norm.







$$\sqrt{s} = 250 + 500 \text{ GeV with } L_{\text{int}} =$$

1σ bounds including 500GeV operation

Six parameters fit

SM-like structure can be measured ~2% New structures a few % Need to improve for bt

 $\Delta\Phi(decay\ planes\ H\rightarrow WW^*)$ need to be reconstructed for bt

w/ ZZH contributions w/ the shape  $\nu \bar{\nu} h + \text{w}/\text{the shape } Zh, h \to WW^*$ 

### The sensitivities to anomalous VVH

· ILA fite Fire the contract of the contract o

$$\Delta \mathcal{L}_{h} = -\eta_{h} \lambda_{0} v_{0} h^{3} + \frac{\theta_{h}}{v_{0}} h \partial_{\mu} h \partial^{\mu} h + \eta_{Z} \frac{m_{Z}^{2}}{v_{0}} Z_{\mu} Z^{\mu} h + \frac{1}{2} 0.5\% (az \sim 2\%)$$

$$+ \eta_{W} \frac{2m_{W}^{2}}{v_{0}} W_{\mu}^{+} W^{-\mu} h + \sim 0.5\% (aw \sim 2\%)$$

common notation difference

$$\zeta_{ZZ} = \frac{v}{\Lambda} b_Z,$$

must convert them with factor of 4.07

$$<0.3\%(bz\sim1\%) + \frac{1}{2}\left(\zeta_{Z}\frac{h}{v_{0}} + \frac{1}{2}\zeta_{2Z}\frac{h^{2}}{v_{0}^{2}}\right)\hat{Z}_{\mu\nu}\hat{Z}^{\mu\nu} + \left(\zeta_{W}\frac{h}{v_{0}} + \frac{1}{2}\zeta_{2W}\frac{h^{2}}{v_{0}^{2}}\right)\hat{W}_{\mu\nu}^{+}\hat{W}^{-\mu\nu} \sim 1\sim2\% \text{ (bw=3\sim7\%)}$$

$$+\frac{1}{2}\left(\zeta_{A}\frac{h}{v_{0}} + \frac{1}{2}\zeta_{2A}\frac{h^{2}}{v_{0}^{2}}\right)\hat{A}_{\mu\nu}\hat{A}^{\mu\nu} + \left(\zeta_{AZ}\frac{h}{v_{0}} + \zeta_{2AZ}\frac{h^{2}}{v_{0}^{2}}\right)\hat{A}_{\mu\nu}\hat{Z}^{\mu\nu} < 0.3\%$$

heavy flavor ID, jet charge ID can improve more for especially WWH

$$< 0.3\% \underbrace{(btz\sim1\%)}_{eehZ} = \underbrace{0.3\%}_{LZhZ} \underbrace{(\tilde{\zeta}_{Z}\frac{h}{v_{0}} + \frac{1}{2}\tilde{\zeta}_{2}\frac{h^{2}}{v_{0}^{2}})^{1}}_{2} \underbrace{Z_{\nu_{0}}^{h^{2}}}_{2} \underbrace{Z_{$$

$$V = A, Z$$

The values given above are direct measurem  $e \bar{n} t^{A,Z}$  without any assumption.

When performing the global fitting by using the other channels the results could be improved more.

In the SM at tree level  $g_A = e$   $g_Z = gc_0$   $g_{1V} = \kappa_V = \eta_H = \eta_Z = \eta_{2Z} = \eta_W = \eta_{2W} = 1$ , and all others =0 In the SM at tree level  $g_A = e$   $g_Z = gc_0$   $g_{1V} = \kappa_V = \eta_H = \eta_Z = \eta_{2Z} = \eta_W = \eta_{2W} = 1$ , and all others =0

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### The sensitivities to anomalous VVH

· IL A find the white where where where

common notation difference

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must convert them with factor of 4.07

$$\mathcal{L}_{ZZH} = M_{\perp}^{2} \Delta \mathcal{L}_{h} = -\eta_{h} \lambda_{0} v_{0} h^{3} + \frac{\theta_{h}}{v_{0}} h \partial_{\mu} h \partial^{\mu} h + \eta_{Z} \frac{m_{Z}^{2}}{v_{0}} Z_{\mu} Z^{\mu} h + \frac{1}{2} 0.5 \% (az \sim 2\%)$$

$$+ \eta_{W} \frac{2m_{W}^{2}}{v_{0}} W_{\mu}^{+} W^{-\mu} h + \gamma_{Z} 0.5 \% (aw \sim 2\%)$$

$$<0.3\%(bz\sim1\%) + \frac{1}{2}\left(\zeta_{Z}\frac{h}{v_{0}} + \frac{1}{2}\zeta_{2Z}\frac{h^{2}}{v_{0}^{2}}\right)\hat{Z}_{\mu\nu}\hat{Z}^{\mu\nu} + \left(\zeta_{W}\frac{h}{v_{0}} + \frac{1}{2}\zeta_{2W}\frac{h^{2}}{v_{0}^{2}}\right)\hat{W}_{\mu\nu}^{+}\hat{W}^{-\mu\nu} \sim 1\sim2\% \text{ (bw=3\sim7\%)}$$

$$+ \frac{1}{2}\left(\zeta_{A}\frac{h}{v_{0}} + \frac{1}{2}\zeta_{2A}\frac{h^{2}}{v_{0}^{2}}\right)\hat{A}_{\mu\nu}\hat{A}^{\mu\nu} + \left(\zeta_{AZ}\frac{h}{v_{0}} + \zeta_{2AZ}\frac{h^{2}}{v_{0}^{2}}\right)\hat{A}_{\mu\nu}\hat{Z}^{\mu\nu} < 0.3\%$$

$$< 0.3\% \underbrace{(btz\sim1\%_{eehZ}^{-1})}_{eehZ} = \underbrace{0.3\%_{eehZ}^{-1}}_{LZh} \underbrace{(\tilde{\zeta}_{Z}\frac{h}{\mu\nu_{0L}})^{-1}_{Z}\tilde{\zeta}_{Z}^{-1}\frac{h^{2}}{v_{0}^{2}}}_{2}\underbrace{(\tilde{\zeta}_{Z}\frac{h}{\nu_{0}})^{-1}_{2}\tilde{\zeta}_{Z}^{-1}}_{2}\underbrace{(\tilde{\zeta}_{Z}\frac{h}{\mu\nu_{0L}})^{-1}_{2}\tilde{\zeta}_{Z}^{-1}}_{2}\underbrace{(\tilde{\zeta}_{Z}\frac{h}{\nu_{0L}})^{-1}_{2}\tilde{\zeta}_{Z}^{-1}}_{2}\underbrace{(\tilde{\zeta}_{Z}\frac{h}{\mu\nu_{0L}})^{-1}_{2}\tilde{\zeta}_{Z}^{-1}}_{2}\underbrace{(\tilde{\zeta}_{Z}\frac{h}{\mu\nu_{0L}})^{-1}_{2}\tilde{\zeta}_{Z}^{-1}}_{2}\underbrace{(\tilde{\zeta}_{Z}\frac{h}{\nu_{0L}})^{-1}_{2}\tilde{\zeta}_{Z}^{-1}}_{2}\underbrace{(\tilde{\zeta}_{Z}\frac{h}{\mu\nu_{0L}})^{-1}_{2}\tilde{\zeta}_{Z}^{-1}}_{2}\underbrace{(\tilde{\zeta}_{Z}\frac{h}{\mu\nu_{0L}})^{-1}_{2}\tilde{\zeta}_{Z}^{-1}}_{2}\underbrace{(\tilde{\zeta}_{Z}\frac{h}{\mu\nu_{0L}})^{-1}_{2}\tilde{\zeta}_{Z}^{-1}}_{2}\underbrace{(\tilde{\zeta}_{Z}\frac{h}{\mu\nu_{0L}})^{-1}_{2}\tilde{\zeta}_{Z}^{-1}}_{2}\underbrace{(\tilde{\zeta}_{Z}\frac{h}{\mu\nu_{0L}})^{-1}_{2}\tilde{\zeta}_{Z}^{-1}}_{2}\underbrace{(\tilde{\zeta}_{Z}\frac{h}{\mu\nu_{0L}})^{-1}_{2}\tilde{\zeta}_{Z}^{-1}}_{2}\underbrace{(\tilde{\zeta}_{Z}\frac{h}{\mu\nu_{0L}})^{-1}_{2}\tilde{\zeta}_{Z}^{-1}}_{2}\underbrace{(\tilde{\zeta}_{Z}\frac{h}{\mu\nu_{0L}})^{-1}_{2}\tilde{\zeta}_{Z}^{-1}}_{2}\underbrace{(\tilde{\zeta}_{Z}\frac{h}{\mu\nu_{0L}})^{-1}_{2}\tilde{\zeta}_{Z}^{-1}}_{2}\underbrace{(\tilde{\zeta}_{Z}\frac{h}{\mu\nu_{0L}})^{-1}_{2}\tilde{\zeta}_{Z}^{-1}}_{2}\underbrace{(\tilde{\zeta}_{Z}\frac{h}{\mu\nu_{0L}})^{-1}_{2}\tilde{\zeta}_{Z}^{-1}}_{2}\underbrace{(\tilde{\zeta}_{Z}\frac{h}{\mu\nu_{0L}})^{-1}_{2}\tilde{\zeta}_{Z}^{-1}}_{2}\underbrace{(\tilde{\zeta}_{Z}\frac{h}{\mu\nu_{0L}})^{-1}_{2}\tilde{\zeta}_{Z}^{-1}}_{2}\underbrace{(\tilde{\zeta}_{Z}\frac{h}{\mu\nu_{0L}})^{-1}_{2}\tilde{\zeta}_{Z}^{-1}}_{2}\underbrace{(\tilde{\zeta}_{Z}\frac{h}{\mu\nu_{0L}})^{-1}_{2}\tilde{\zeta}_{Z}^{-1}}_{2}\underbrace{(\tilde{\zeta}_{Z}\frac{h}{\mu\nu_{0L}})^{-1}_{2}\tilde{\zeta}_{Z}^{-1}}_{2}\underbrace{(\tilde{\zeta}_{Z}\frac{h}{\mu\nu_{0L}})^{-1}_{2}\tilde{\zeta}_{Z}^{-1}}_{2}\underbrace{(\tilde{\zeta}_{Z}\frac{h}{\mu\nu_{0L}})^{-1}_{2}\tilde{\zeta}_{Z}^{-1}}_{2}\underbrace{(\tilde{\zeta}_{Z}\frac{h}{\mu\nu_{0L}})^{-1}_{2}\tilde{\zeta}_{Z}^{-1}}_{2}\underbrace{(\tilde{\zeta}_{Z}\frac{h}{\mu\nu_{0L}})^{-1}_{2}\tilde{\zeta}_{Z}^{-1}}_{2}\underbrace{(\tilde{\zeta}_{Z}\frac{h}{\mu\nu_{0L}})^{-1}_{2}\tilde{\zeta}_{Z}^{-1}}_{2}\underbrace{(\tilde{\zeta}_{Z}\frac{h}{\mu\nu_{0L}})^{-1}_{2}\tilde{\zeta}_{Z}^{-1}}_{2}\underbrace{(\tilde{\zeta}_{Z}\frac{h}{\mu\nu_{0L}})^{-1}_{2}\tilde{\zeta}_{Z}^{-1}}_{2}\underbrace{(\tilde{\zeta}_{Z}\frac{h}{\mu\nu_{0L}})^{-1}_{2}\tilde{\zeta}_{Z}^{-1}}_{2}\underbrace{(\tilde{\zeta}_{Z}\frac{h}{\mu\nu_{0L}})^{-1}_{2}\tilde{\zeta}_{Z}^{-1}}_{2}\underbrace{(\tilde{\zeta}_{Z}\frac{h}{\mu\nu_{0L}})^{-1}_{2}\tilde{\zeta}_{Z}^{-1}}_{2}\underbrace{(\tilde{\zeta}_{Z}\frac{h}{\mu\nu_{0L}})^{-1}_{2}\tilde{\zeta}_{Z}^{-1}}_{2}\underbrace{(\tilde{\zeta}_{Z}\frac{h}{\mu\nu_{0L}})^{-1}_{2}\tilde{\zeta}_{Z}^{-1}}_{2}\underbrace{(\tilde{\zeta}_{Z}\frac{h}{\mu\nu_{0L}})^{-1}_{2}\tilde{\zeta}_{Z}^{-1}}_{2}\underbrace{(\tilde{\zeta}_{Z}\frac{h}{\mu\nu_{0L}})^{-1}_{2}\tilde{\zeta}_{Z}^{-1}}_{2}\underbrace{(\tilde{\zeta}_{Z}\frac{h}{\mu\nu_{0L}})^{-1}_{2}\tilde{\zeta}$$

heavy flavor ID, jet charge ID can improve more for especially WWH

LHC ATLAS: EFT analysis

JHEP 03 (2018) 095

DOI: 10.100 // JHEP 40 2 (2018) 095

V = A, Z

$$\mathcal{L}_{0}^{V} = \left\{ \kappa_{\text{SM}} \left[ \frac{1}{2} g_{HZZ} Z_{\mu} Z^{\mu} + g_{HWW} W_{\mu}^{+} W^{-\mu} \right] \right.$$

$$- \frac{1}{4} \left[ \kappa_{Hgg} g_{Hgg} G_{\mu\nu}^{a} G^{a,\mu\nu} + \tan \alpha \kappa_{Agg} g_{Agg} G_{\mu\nu}^{a} \tilde{G}^{a,\mu\nu} \right] =$$

$$- \frac{1}{1} \frac{1}{4 \ln^{2}} \left[ \kappa_{HZZ} Z_{\mu\nu} Z^{\mu\nu} + \tan \alpha \kappa_{AZZ} Z_{\mu\nu} \tilde{Z}^{\mu\nu} \right]$$

$$- \frac{1}{2} \frac{1}{\Lambda} \left[ \kappa_{HWW} W_{\mu\nu}^{+} W^{-\mu\nu} + \tan \alpha \kappa_{AWW} W_{\mu\nu}^{+} W^{-\mu\nu} \right] \left\{ \kappa_{HWW} W_{\mu\nu}^{+} W^{-\mu\nu} + \tan \alpha \kappa_{AWW} W_{\mu\nu}^{+} W^{-\mu\nu} \right\} \left\{ \kappa_{HWW} W_{\mu\nu}^{+} W^{-\mu\nu} + \tan \alpha \kappa_{AWW} W_{\mu\nu}^{+} W^{-\mu\nu} \right\} \left\{ \kappa_{HWW} W_{\mu\nu}^{+} W^{-\mu\nu} + \tan \alpha \kappa_{AWW} W_{\mu\nu}^{+} W^{-\mu\nu} \right\} \left\{ \kappa_{HWW} W_{\mu\nu}^{+} W^{-\mu\nu} + \tan \alpha \kappa_{AWW} W_{\mu\nu}^{+} W^{-\mu\nu} \right\} \left\{ \kappa_{HWW} W_{\mu\nu}^{+} W^{-\mu\nu} + \tan \alpha \kappa_{AWW} W_{\mu\nu}^{+} W^{-\mu\nu} \right\} \left\{ \kappa_{HWW} W_{\mu\nu}^{+} W^{-\mu\nu} + \tan \alpha \kappa_{AWW} W_{\mu\nu}^{+} W^{-\mu\nu} \right\} \left\{ \kappa_{HWW} W_{\mu\nu}^{+} W^{-\mu\nu} + \tan \alpha \kappa_{AWW} W_{\mu\nu}^{+} W^{-\mu\nu} \right\} \left\{ \kappa_{HWW} W_{\mu\nu}^{+} W^{-\mu\nu} + \tan \alpha \kappa_{AWW} W_{\mu\nu}^{+} W^{-\mu\nu} \right\} \left\{ \kappa_{HWW} W_{\mu\nu}^{+} W^{-\mu\nu} + \tan \alpha \kappa_{AWW} W_{\mu\nu}^{+} W^{-\mu\nu} \right\} \right\} \left\{ \kappa_{HWW} W_{\mu\nu}^{+} W^{-\mu\nu} + \tan \alpha \kappa_{AWW} W_{\mu\nu}^{+} W^{-\mu\nu} \right\} \left\{ \kappa_{HWW} W_{\mu\nu}^{+} W^{-\mu\nu} + \tan \alpha \kappa_{AWW} W_{\mu\nu}^{+} W^{-\mu\nu} \right\} \left\{ \kappa_{HWW} W_{\mu\nu}^{+} W^{-\mu\nu} + \tan \alpha \kappa_{AWW} W_{\mu\nu}^{+} W^{-\mu\nu} \right\} \right\} \left\{ \kappa_{HWW} W_{\mu\nu}^{+} W^{-\mu\nu} + \tan \alpha \kappa_{AWW} W_{\mu\nu}^{+} W^{-\mu\nu} \right\} \left\{ \kappa_{HWW} W_{\mu\nu}^{+} W^{-\mu\nu} + \tan \alpha \kappa_{AWW} W_{\mu\nu}^{+} W^{-\mu\nu} \right\} \left\{ \kappa_{HWW} W_{\mu\nu}^{+} W^{-\mu\nu} + \tan \alpha \kappa_{AWW} W_{\mu\nu}^{+} W^{-\mu\nu} \right\} \left\{ \kappa_{HWW} W_{\mu\nu}^{+} W^{-\mu\nu} + \tan \alpha \kappa_{AWW} W_{\mu\nu}^{+} W^{-\mu\nu} \right\} \left\{ \kappa_{HWW} W_{\mu\nu}^{+} W^{-\mu\nu} + \tan \alpha \kappa_{AWW} W_{\mu\nu}^{+} W^{-\mu\nu} \right\} \left\{ \kappa_{HWW} W_{\mu\nu}^{+} W^{-\mu\nu} + \tan \alpha \kappa_{AWW} W_{\mu\nu}^{+} W^{-\mu\nu} \right\} \left\{ \kappa_{HWW} W_{\mu\nu}^{+} W^{-\mu\nu} + \tan \alpha \kappa_{HWW} W_{\mu\nu}^{+} W^{-\mu\nu} \right\} \left\{ \kappa_{HWW} W_{\mu\nu}^{+} W^{-\mu\nu} + \tan \alpha \kappa_{HWW} W_{\mu\nu}^{+} W^{-\mu\nu} \right\} \left\{ \kappa_{HWW} W_{\mu\nu}^{+} W^{-\mu\nu} + \tan \alpha \kappa_{HWW} W_{\mu\nu}^{+} W^{-\mu\nu} \right\} \left\{ \kappa_{HWW} W_{\mu\nu}^{+} W^{-\mu\nu} + \tan \alpha \kappa_{HWW} W_{\mu\nu}^{+} W^{-\mu\nu} \right\} \left\{ \kappa_{HWW} W_{\mu\nu}^{+} W^{-\mu\nu} + \tan \alpha \kappa_{HWW} W_{\mu\nu}^{+} W^{-\mu\nu} \right\} \left\{ \kappa_{HWW} W_{\mu\nu}^{+} W^{-\mu\nu} + \tan \alpha \kappa_{HWW} W_{\mu\nu}^{+} W^{-\mu\nu} \right\} \left\{ \kappa_{HWW} W_{\mu\nu}^{+} W^{-\mu\nu} + \tan \alpha \kappa_{HWW} W_{\mu\nu}^{+} W^{-\mu\nu} \right\} \left\{ \kappa_{HWW} W_{\mu\nu}^{+} W^$$

Expected and observed confidence intervals at 95% CL with 36.1 fb<sup>-1</sup> of data at  $\sqrt{s}$  = 13 TeV.

	V - 1.7	7					
BSM cou	pling	Fit	Expected	Observed	Best-fit	Best-fit	Deviation
$\kappa_{ m BSN}$	_	configuration	conf. inter.	conf. inter.	$\hat{\kappa}_{ ext{BSM}}$	$\hat{\kappa}_{ ext{SM}}$	from SM
$= n_{xx} = \kappa_{HVV}$	$= n_{\alpha \beta} \stackrel{(\kappa_H)}{=}$	$g_{gg} \equiv \frac{1}{\eta_{2W}} \kappa_{SM} = 1$ , a	nd all others	$=0^{[0.8, 4.5]}$	2.9	-	$2.3\sigma$
$=e^{-KH}YY$	$V = gC_{\circ}^{(\kappa_H)}$	$g_{\mathcal{S}} = 1$ , $k_{\text{SM}} = 1$	$= [\bar{\eta}_{3}, \underline{1}, 4, 0] =$	$n^{[-0.6,4.2]}$	and all	others	$=0^{1.7\sigma}$
$K_{AVV}$	$\kappa_H \sim \kappa_H$	$\kappa_{gg} = 1, \kappa_{SM} = 1$	[-3.5, 3.5]	[-5.2, 5.2]	±2.9	-	$1.4\sigma$
$\kappa_{AVV}$	$\kappa_H$	$ \chi_{gg} = 1,  \kappa_{\text{SM}} \text{ free} $	[-4.0, 4.0]	[-4.4, 4.4]	±1.5	1.2	$0.5\sigma$

( given inverse power is  $\Lambda$  )

In the SZZHI tree Tevel  $\frac{30\%}{200\%}$ ,  $\frac{200\%}{200\%}$  =  $gc_0^{HL}$  -  $gc_0^{HL}$  -

- 1). An overview of an anomalous VVH study
  ZZH/γZH and WWH induced with dim-6 operators
- 2). An application of a Matrix Element method toward further improvement of the sensitivity  $ZH \rightarrow \mu^+\mu^- H$ ,  $\sqrt{s} = 250 GeV$

## **Matrix Element Method**

An objective is clear

Try to encode all available kinematical information on an event into a single observable. LHC, Tevatron ... have used it!

Observation in an event in terms of differential  $\sigma$ 

LO

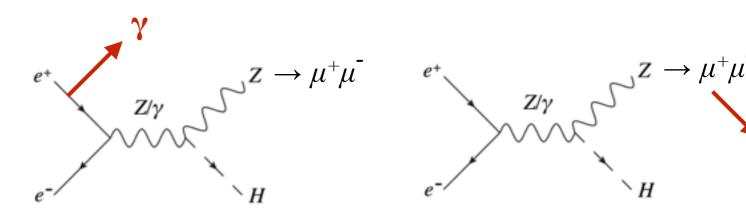
$$P(\vec{p}^{\,\mu}) = \frac{|\mathcal{M}(\vec{p}^{\,\mu})|^2}{\sigma} d\Phi$$

Probability = 
$$\begin{vmatrix} e^{+} & Z/\gamma & Z \rightarrow \mu^{+}\mu^{-} \\ e^{-} & H \end{vmatrix}^{2} d\Phi$$
Matrix Element

However, ISR, beam-strahlung, and FSR

NLO effects

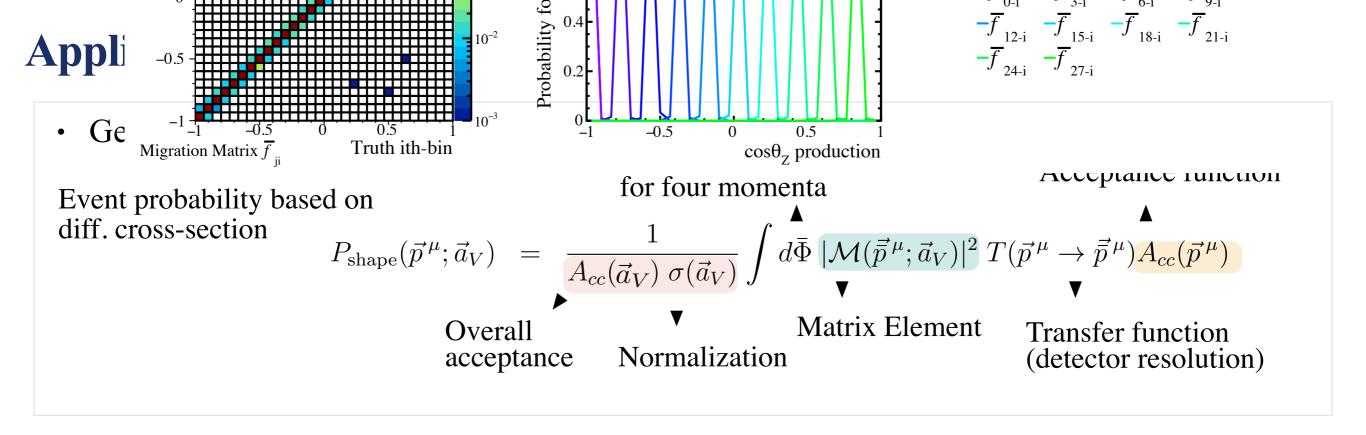
Matrix Element doesn't fit reaction anymore

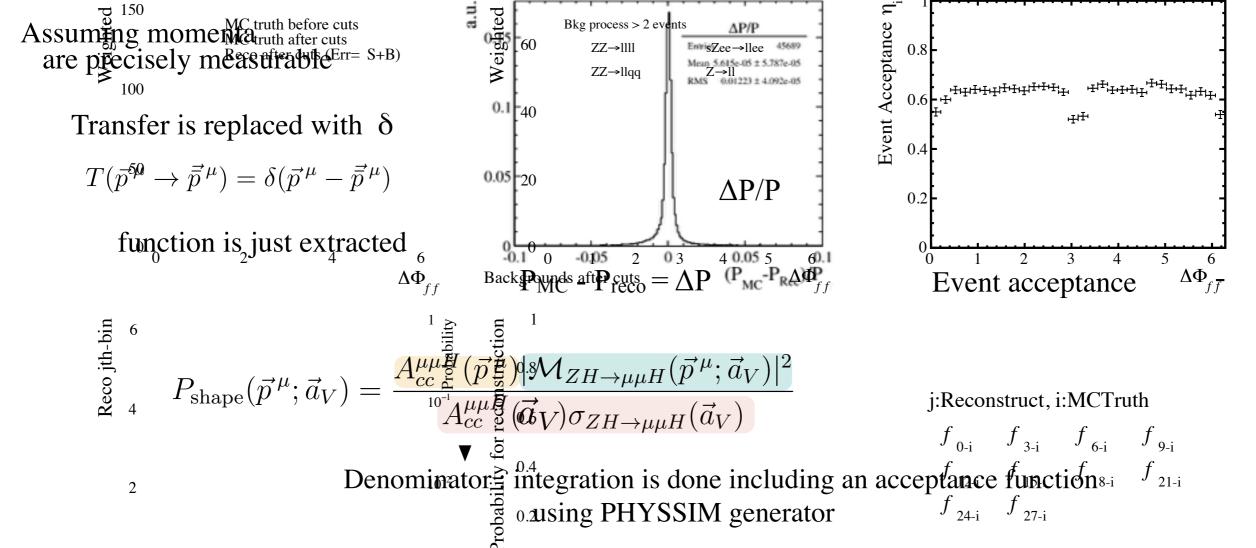


• ILCSoft framework: Marlin-PHYSSIM

The development is on going by Junping, Keisuke

Matrix Element Calculation
based on PHYSSIM, Junping Tian
<a href="https://agenda.linearcollider.org/event/6301/contributions/29469/attachments/24440/37804/MatrixElement\_AWLC14.pdf">https://agenda.linearcollider.org/event/6301/contributions/29469/attachments/24440/37804/MatrixElement\_AWLC14.pdf</a>





## **Application: trial for the signal**

Chi-squared

$$\chi^2 = -2\log\Delta\mathcal{L} = -2w(\log\mathcal{L}(\vec{a}_V) - \log\mathcal{L}_{SM})$$

W: a factor for scaling the norm. to #expected ~1623 (after bkg suppression in the shape analysis)

Likelihood function (unbinned estimation)

$$\mathcal{L}(\vec{a}_{V}) = \mathcal{L}_{\text{shape}}(\vec{a}_{V}) \cdot \mathcal{L}_{\text{norm}}(\vec{a}_{V})$$

$$= \prod_{i=1}^{\text{MCevents}} P_{\text{shape}}(\vec{p}_{i}^{\,\mu}; \vec{a}_{V}) \cdot P_{\text{norm}}(\vec{a}_{V})$$

Event probability

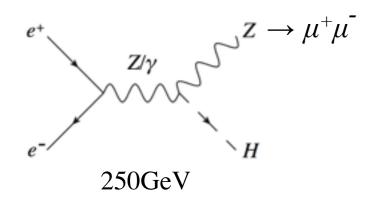
$$P_{\text{shape}}(\vec{p}^{\,\mu}; \vec{a}_{V}) = \frac{A_{cc}^{\mu\mu H}(\vec{p}^{\,\mu}) |\mathcal{M}_{ZH \to \mu\mu H}(\vec{p}^{\,\mu}; \vec{a}_{V})|^{2}}{A_{cc}^{\mu\mu H}(\vec{a}_{V}) \sigma_{ZH \to \mu\mu H}(\vec{a}_{V})} \blacktriangleright$$

momenta:  $\mu$ ,  $\mu$ , and it's recoil info.

Denominator:

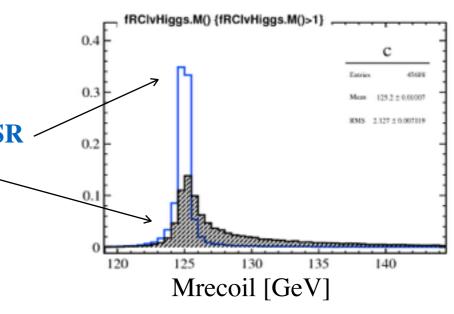
integration is done including Acc is also calculated without ISR, BSL, FSR

MarlinPhyssim: Calculator is LO



Sample:

- 1). no ISR, no BSL, and no FSR -
- 2). with ISR, BSL and FSR



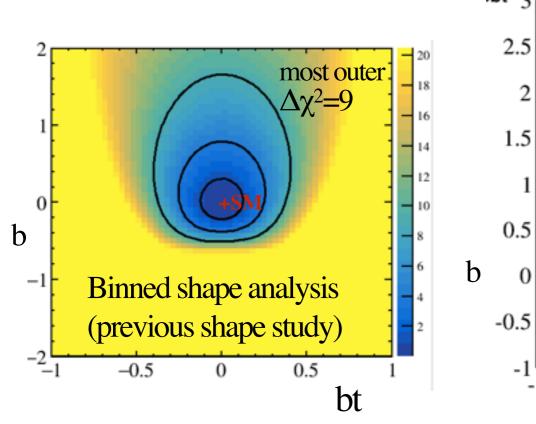
## Application: trial for the signal

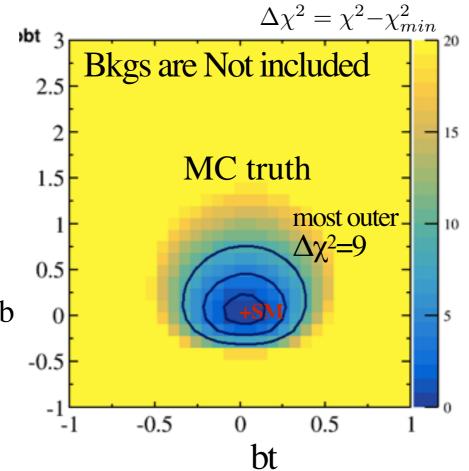
- b bs bt contours in the 2-parameter space
  - A consistent situation: LO, hopefully it's perspective improvement

• ZH $\rightarrow \mu^+\mu^-H$  (signal only) 250GeV 250fb<sup>-1</sup> [b vs bt] ME : is LO

Sample: no ISR, BSL, and FSR

Denomi. : is calculated based on LO





## **Application: trial for the signal**

- b bs bt contours in the 2-parameter space
  - A consistent situation: LO, hopefully it's perspective improvement
  - NLO effects,  $\rightarrow$  change shape, direct usage of momenta give large impact → shift minimum, falsehood sensitivity
    - Need to handle NLO effects correctly if wants to exceed 1% sensitivity

 $ZH \rightarrow \mu^+\mu^- H$  (signal only) 250GeV 250fb<sup>-1</sup> [b vs bt]

ME : is LO

Sample : no ISR, BSL, and FSR

Denomi.: is calculated based on LO

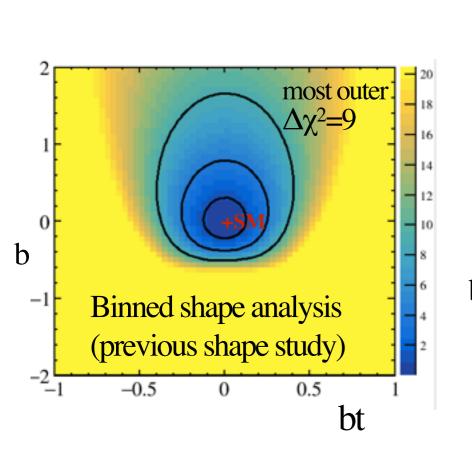
ME : is LO

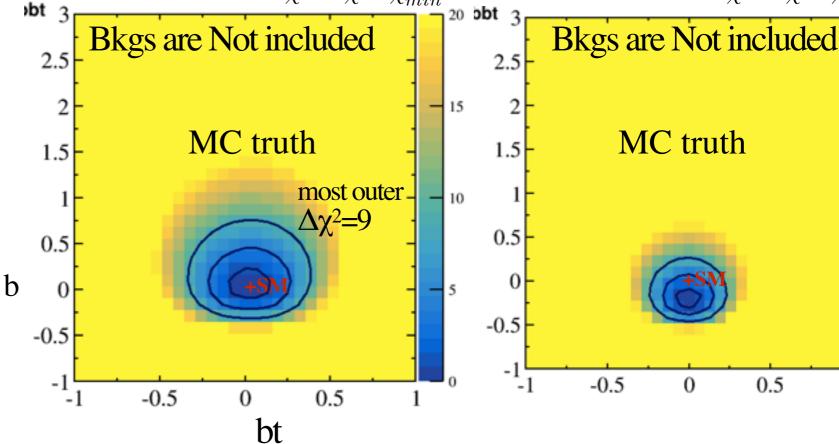
: with ISR, BSL, and FSR Sample

Denomi. : is calculated based on LO

 $\Delta \chi^2 = \chi^2 - \chi^2_{min}$ 

0.5





 $\Delta\chi^2=\chi^2{-}\chi^2_{min}$ 

## **Summary**

- 1). An overview of the anomalous VVH study ZZH/γZH and WWH induced with dim-6 operators
  - Model independently the sensitivities to the structures were evaluated. (including 500GeV operation)
    - SM-like ZZH/WWH structures ~2%
    - new ZZH/ $\gamma$ ZH structures < 1%
    - new WWH structures  $3\sim7\%$  and  $\sim17\%$
- 2). An application of a Matrix Element method toward further improvement of the sensitivity
  - Try to encode all information into a single observable

Intrinsically the improvement could be given, however, it turns out that NLO effects (ISR, BSL) affect to results largely when discussing the sensitivity ~1%

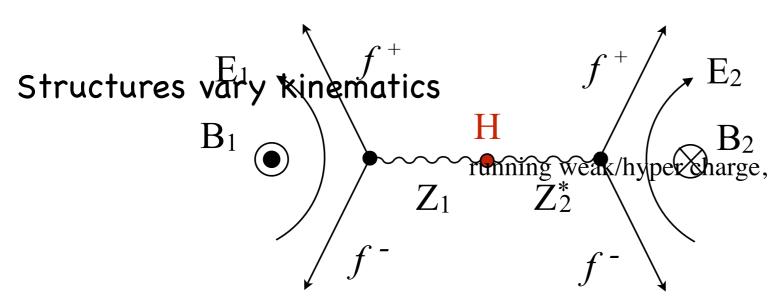
Need to handle carefully, we will start to develop it to include ISR & BSL



### Focusing on ZZH

SM-like coupling

• a different CP-even structure



• a CP-violating structure

one like an electroweak magnetic field could be generated

#### Result in EM dynamics

would give peculiar kinematical distributions

$$\hat{F}_{\mu\nu}\hat{F}^{\mu\nu}$$
  $\propto$   $\boldsymbol{B}_1\cdot\boldsymbol{B}_2-\boldsymbol{E}_1\cdot\boldsymbol{E}_2$   $\hat{F}_{\mu\nu}\tilde{F}^{\mu\nu}$   $\propto$   $\boldsymbol{E}_1\cdot\boldsymbol{B}_2$  take a parallel state

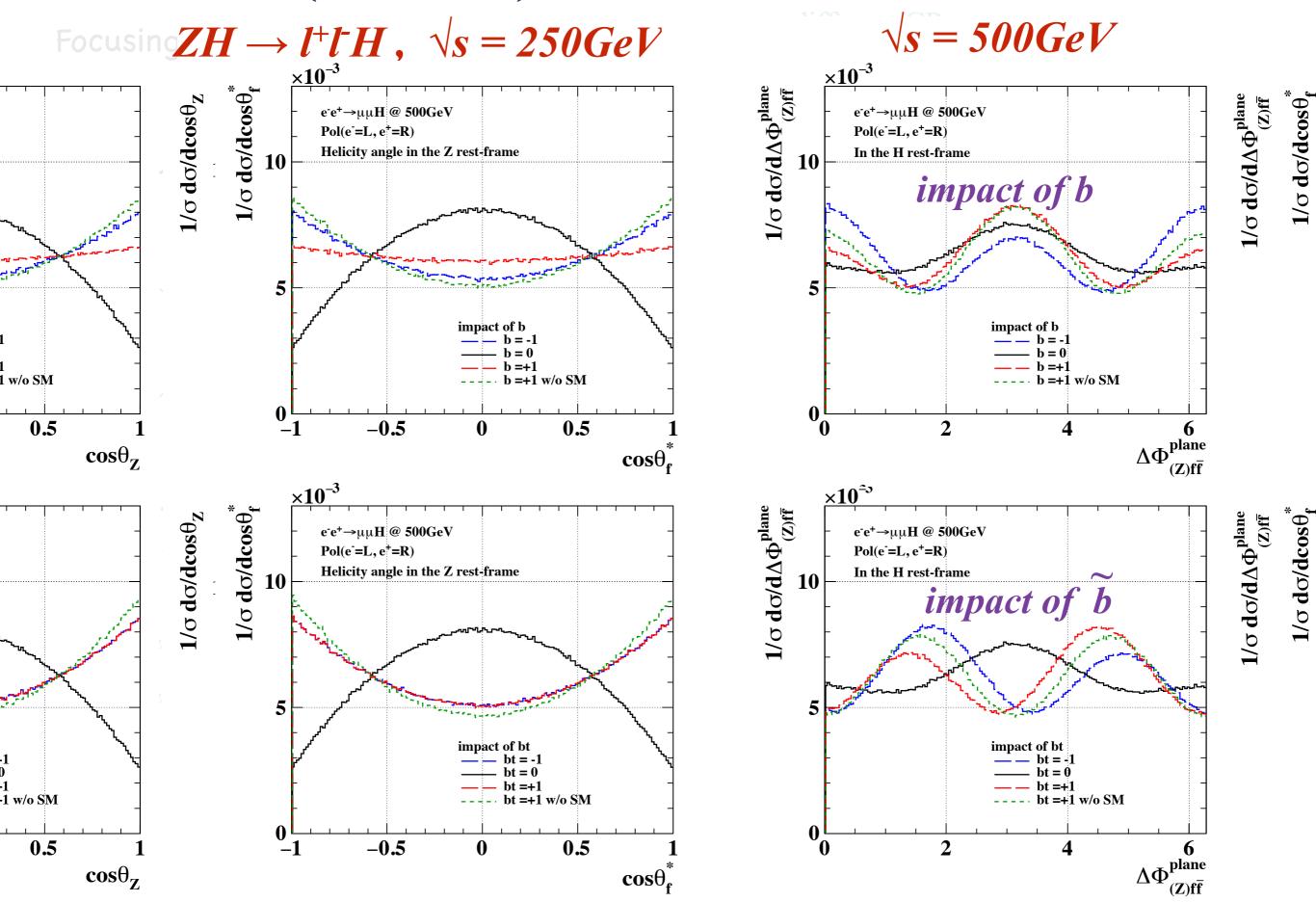
makes both planes tend to take a perpendicular state

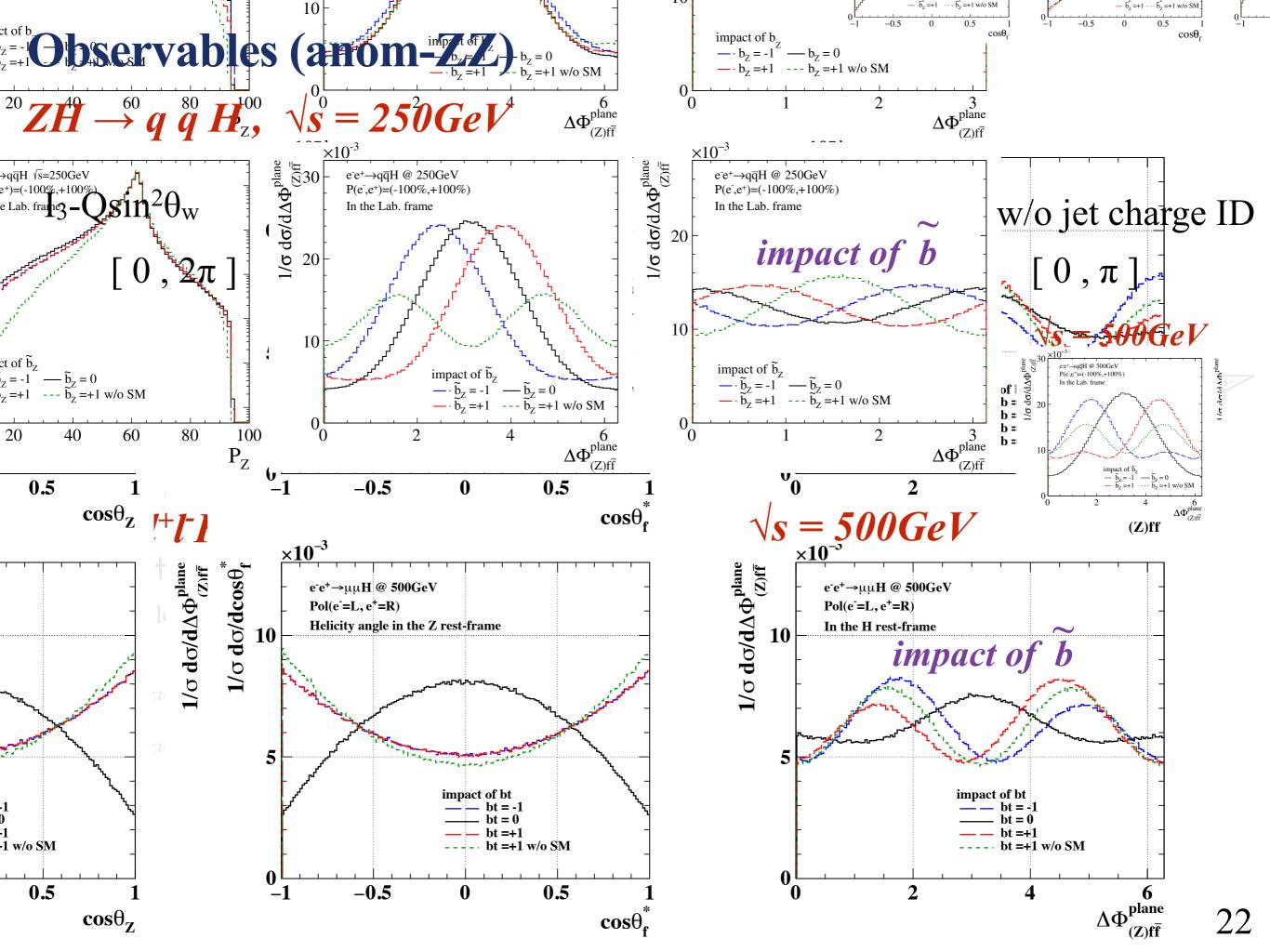
$$e^+e^- \rightarrow ZH \rightarrow l^+l^-H$$

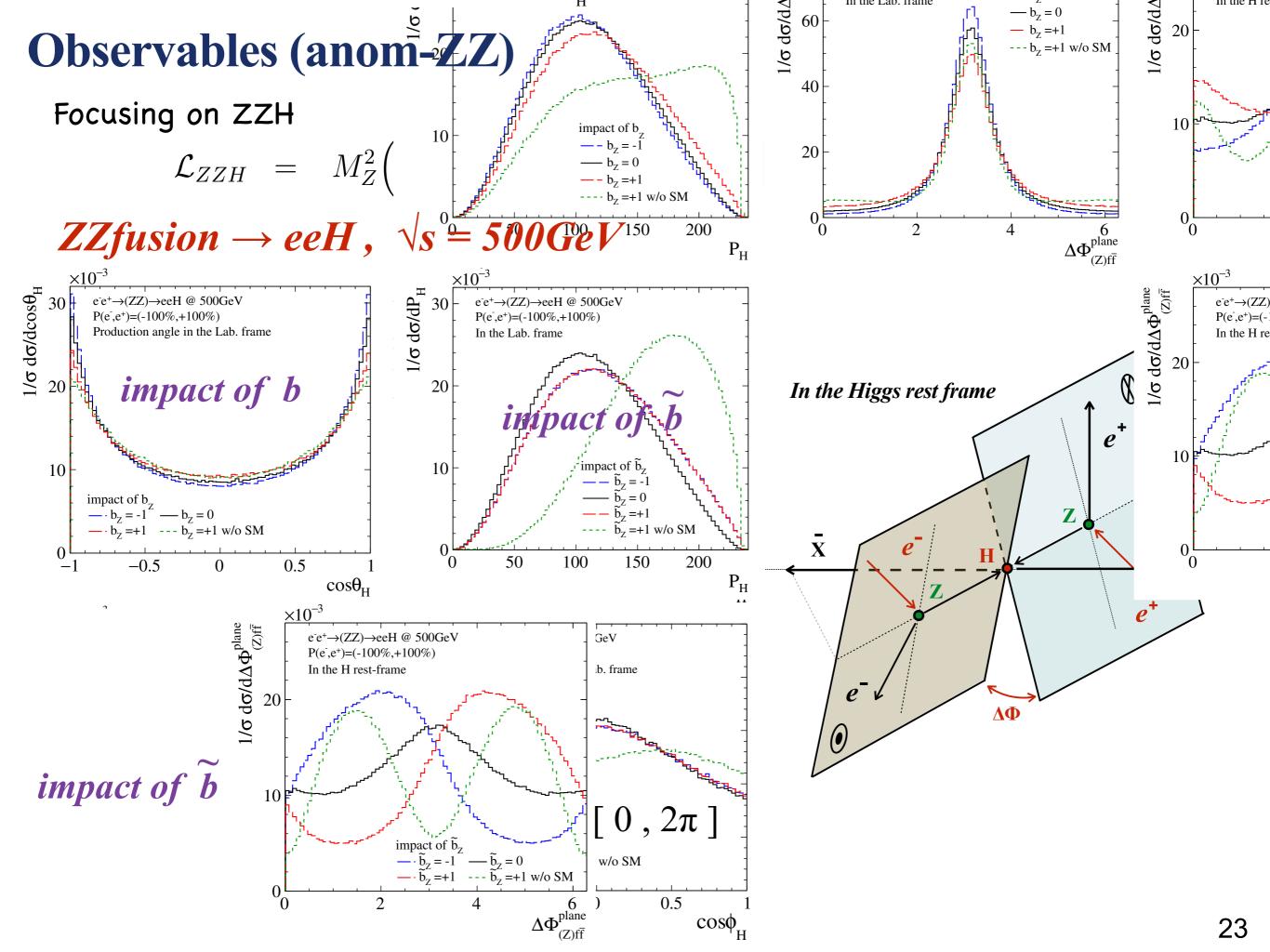
 $cos\theta z$ : a production of the Z.

 $\cos\theta f^*$ : a helicity angle of a Z's daughter.

 $\Delta\Phi$ : an angle b/w two production plane.







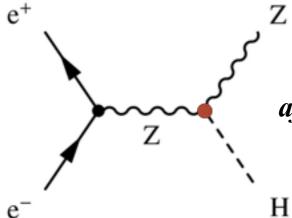
Focusing on WWH

• a different CP-even structure

cusing on WWH SM-like coupling • a different CP-even structure 
$$\mathcal{L}_{WWH} = 2M_W^2 \left(\frac{1}{V} + \frac{a_W}{\Lambda}\right) W_{\mu}^+ W^{-\mu} H + \frac{b_W}{\Lambda} \hat{W}_{\mu\nu}^+ \hat{W}^{-\mu\nu} H + \frac{\tilde{b}_W}{\Lambda} \hat{W}_{\mu\nu}^+ \hat{W}^{-\mu\nu} H + \frac{\tilde{b}_W}{\Lambda} \hat{W}_{\mu\nu}^+ \hat{W}^{-\mu\nu} H$$

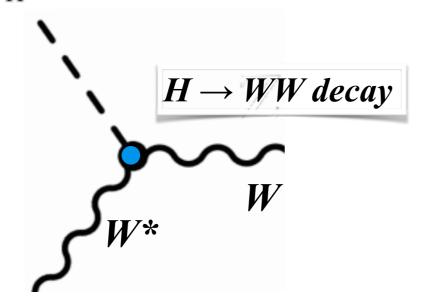
a CP-violating structure

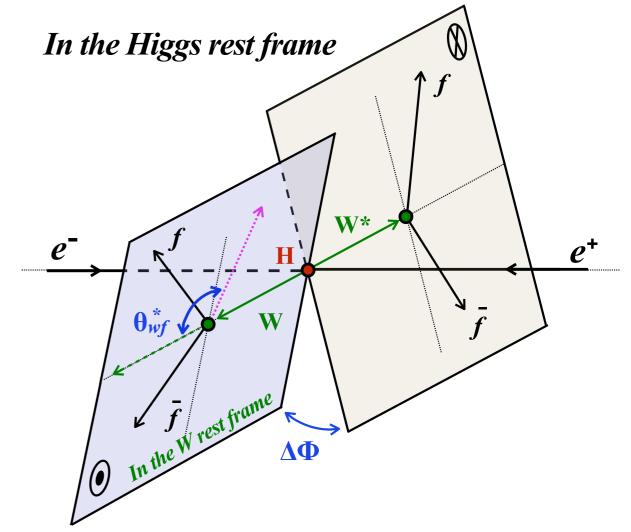
#### The Higgs-straulung



affects ZH production

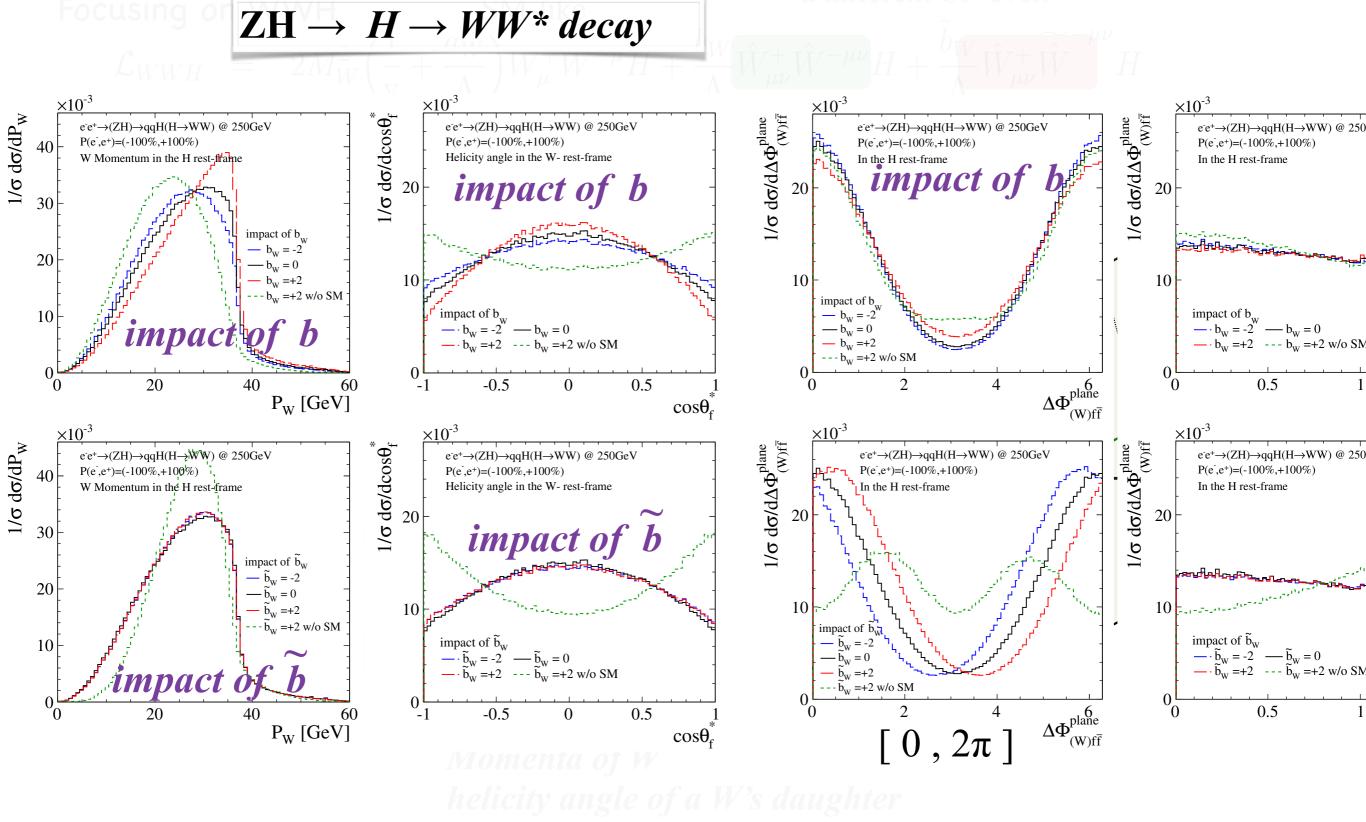
H rest-frame is not affected



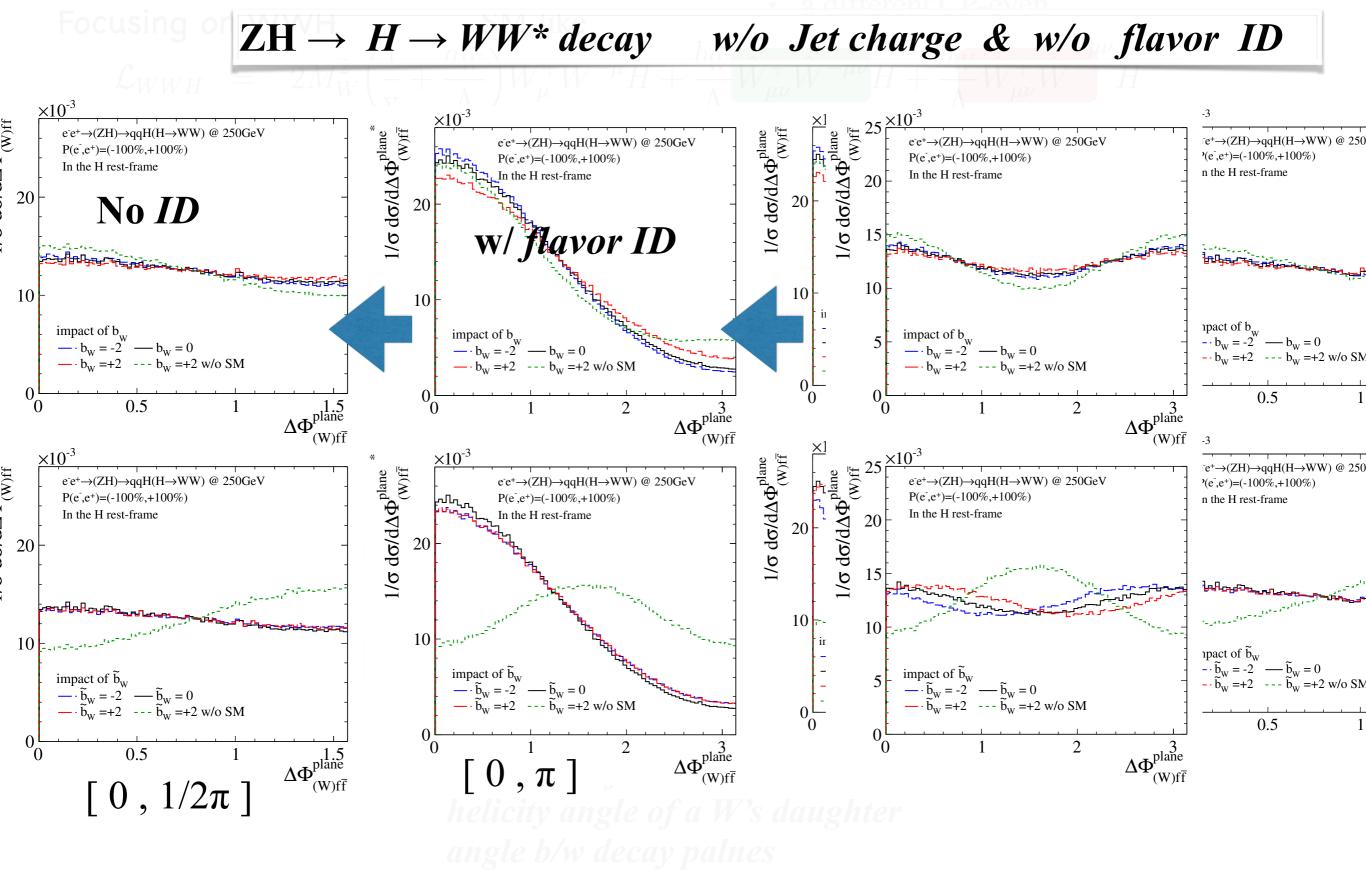


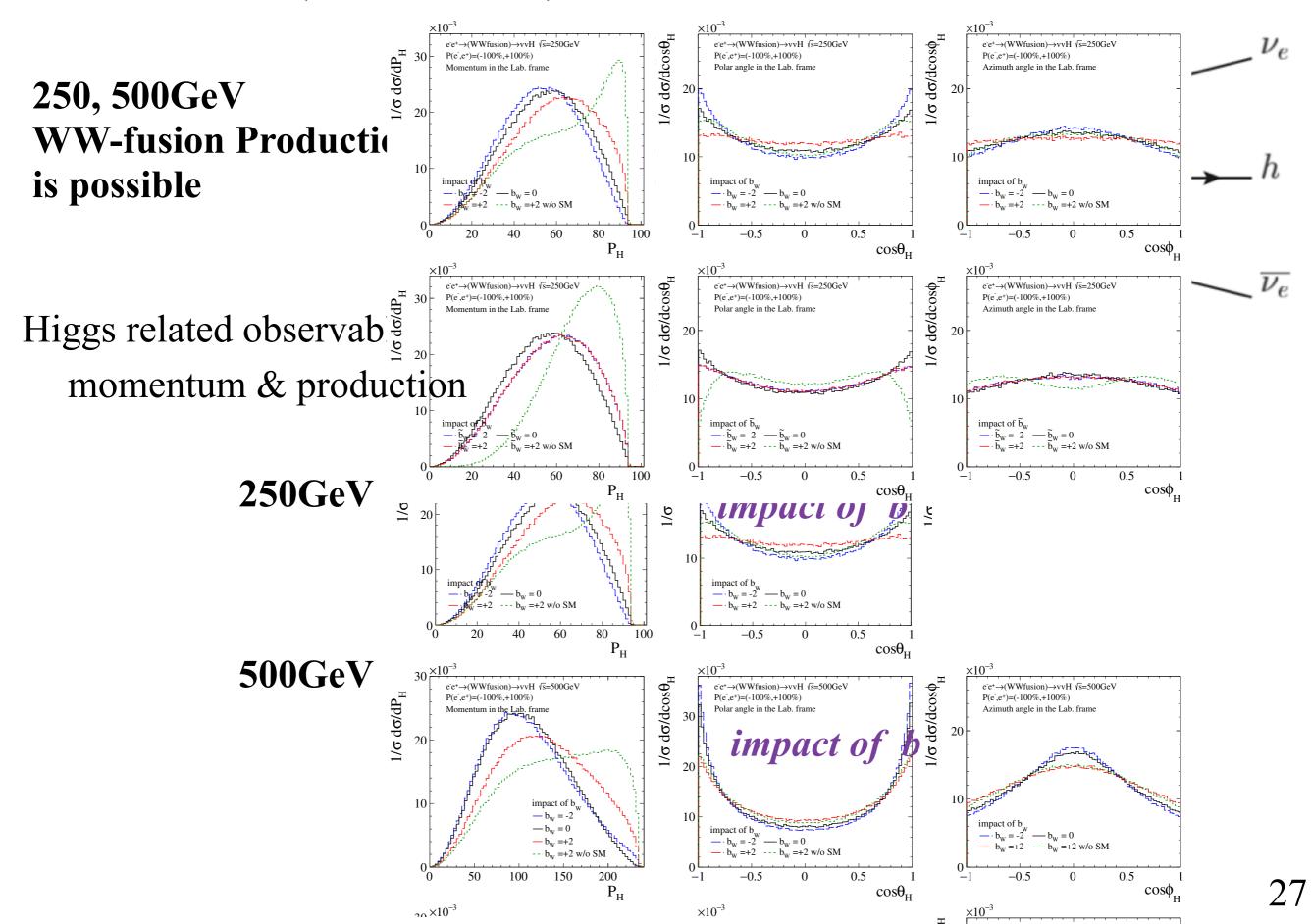
**Observable** 

Momenta of W helicity angle of a W's daughter angle b/w decay palnes



25





## Observables (Prod

$$\mathcal{L}_{ZZH} = M_Z^2 \left(\frac{1}{\mathbf{v}} + \frac{a_Z}{\Lambda}\right) Z_{\mu} Z^{\mu} H + \frac{b_Z}{2\Lambda} \hat{Z}_{\mu\nu} \hat{Z}^{\mu\nu} H + \frac{\tilde{b}_Z}{2\Lambda} \hat{Z}_{\mu\nu} \hat{Z}^{\mu\nu} H^{\mathrm{cos}\theta}_{\mathrm{higg}}$$

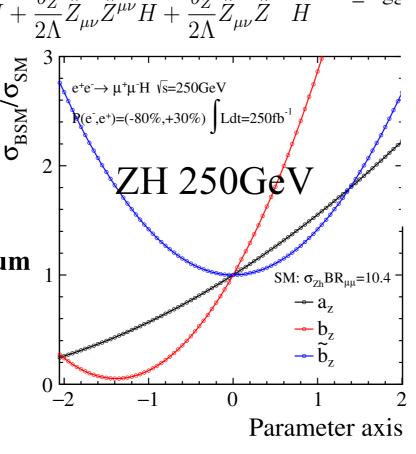
No energy dependence on a Recover the SM with  $-\Lambda/v$ 

b bt vary depending on momentum

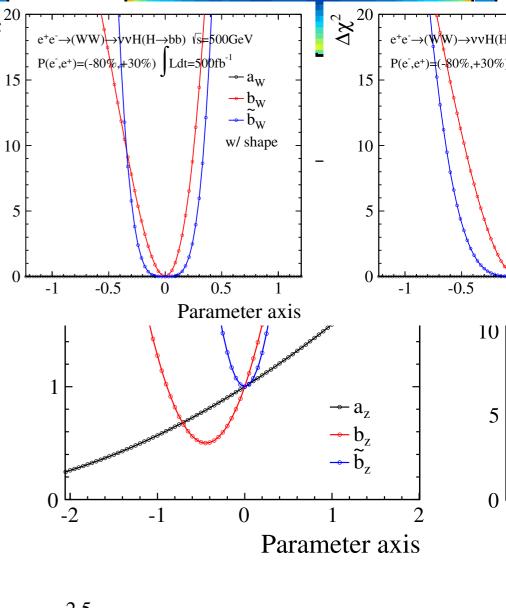
bt change symmetric

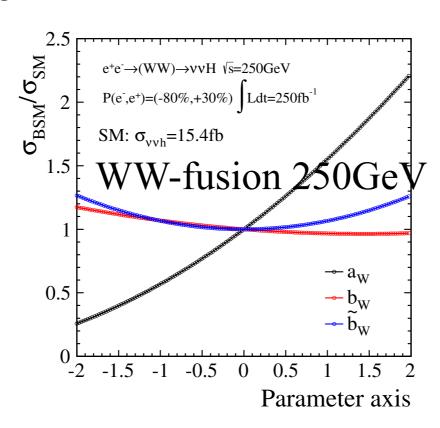
$$(F_{\mu\nu}\tilde{F}^{\mu\nu})^2$$
 gives

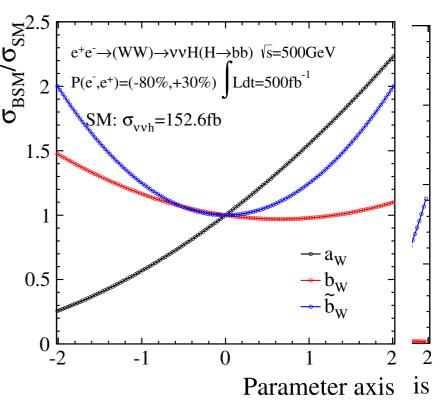
one term with positive sign.



Normalized to a.u.



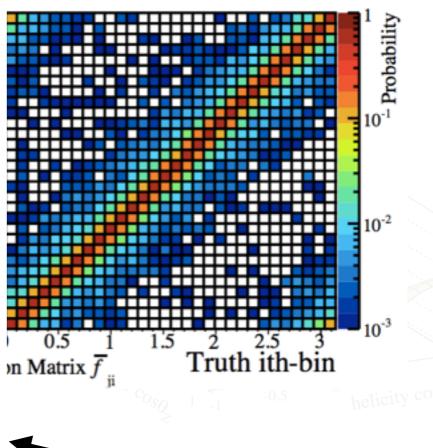




## function

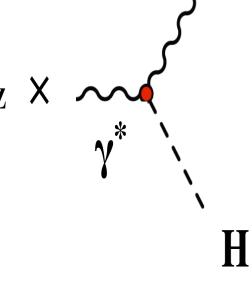
## natrix f

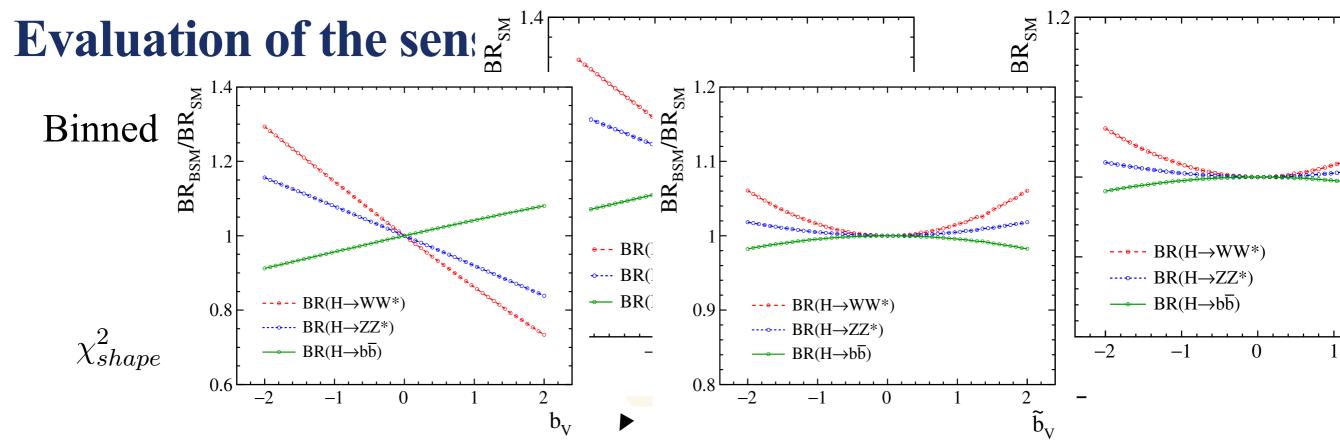
s observed in reality)





Gen





#### Poisson error on each bin

(SM Bkgs are taken into account)

## Normalization (Cross-section)

$$\chi_{norm}^{2} = \left[ \frac{N_{SM} - N_{BSM}(a_{V}, b_{V}, \tilde{b}_{V})}{\delta \sigma_{Zh/eeh} \cdot N_{SM}} \right]^{2}$$

# Relative errors of cross-section measurement

(SM Bkgs are taken into account)

#### **Detector response function**

→ Transfer the theory to "Detector level" distribution

full simulation, T. Barklow et al., "ILC Operating Scenarios", arXiv:1506.07830 [hep-ex]

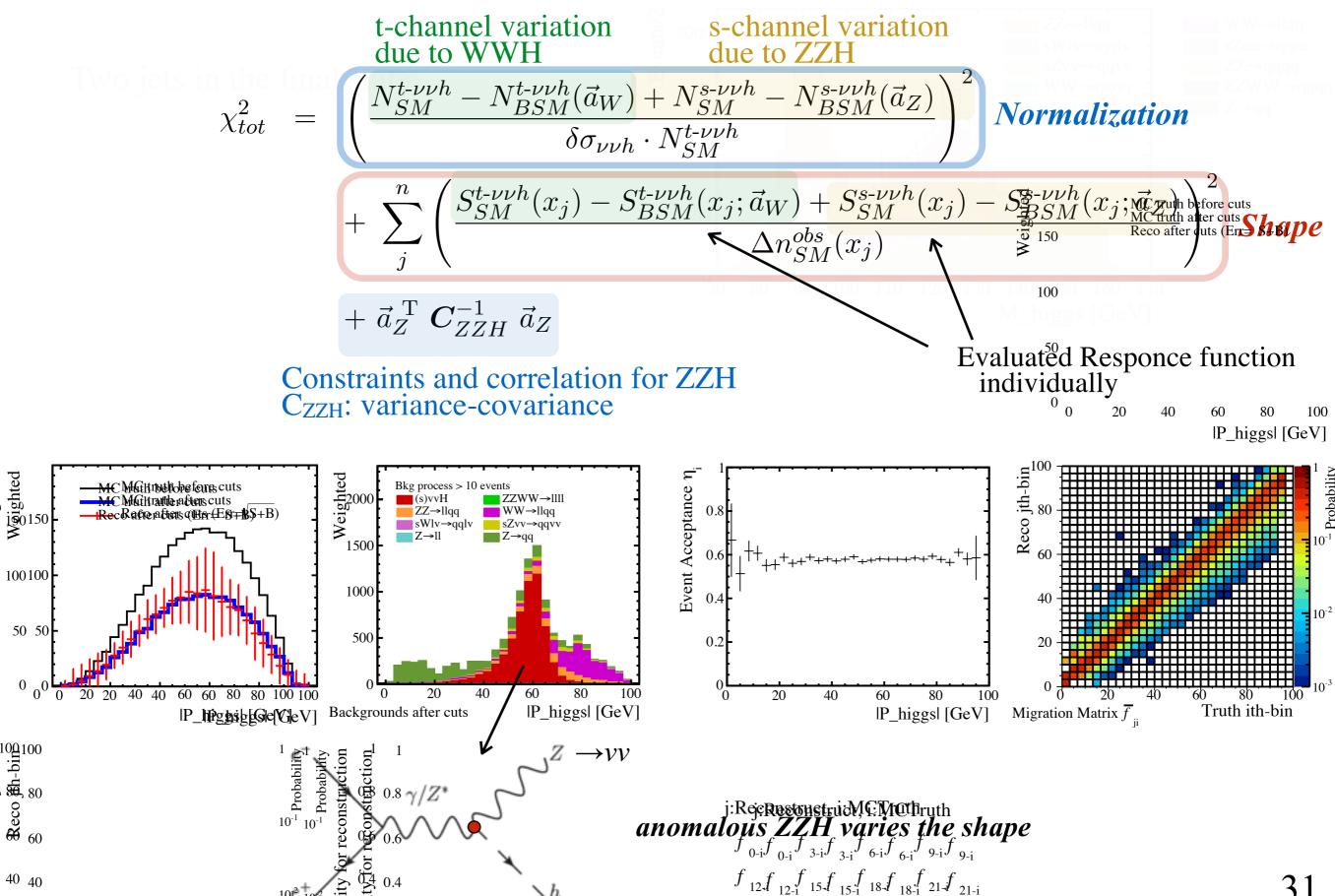
$$\delta\sigma(Zh) = 2.0 \% \text{ and } 3.0 \%$$
 for 250 and 500 GeV

 $\delta\sigma_{eeh}$  are 27.16 % and 5.32 % for 250 and 500 GeV

$$\delta(\sigma_{eeh} \cdot BR_{hbb}) = 27.0 \%$$
 and 4.0 % for 250 and 500 GeV  $\delta BR_{hbb} = 2.9 \%$  and 3.5 % for 250 and 500 GeV

20 20

8.1% and 1.0%.



 $f_{24}f_{24}f_{24}f_{27}f_{27-i}$ 

## ZZH a-b

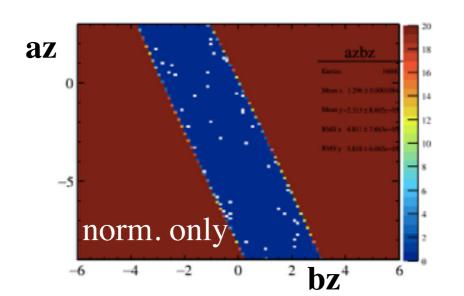
#### Normalization from **ZH** and **ZZ**

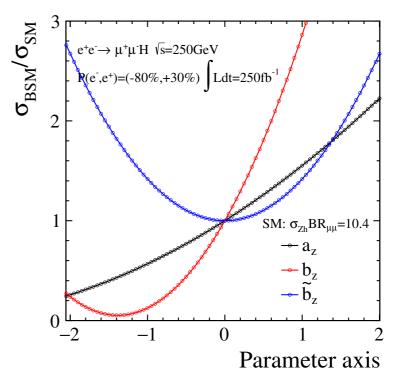
Both az & bz can adjust each other by making σ increase & decrease

Both az & bz make  $\sigma$  increase, and any btz can not adjust since any btz can increase  $\sigma$ . Thus, the bound is quickly restricted

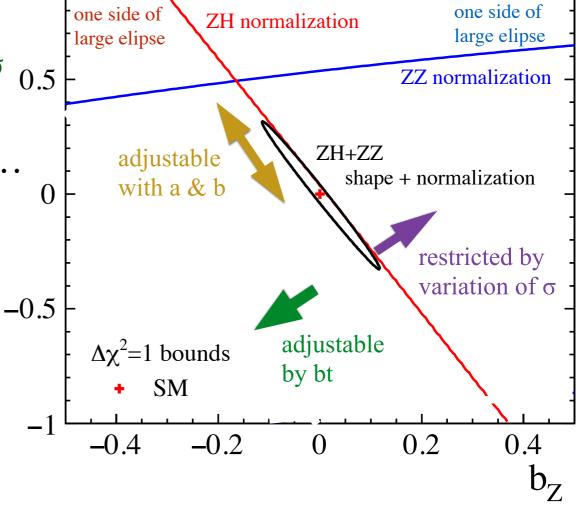
For this direction both az & bz make  $\sigma$  decrease, and btz has huge room to recover the SM value by increasing  $\sigma$ 

Once the shape is included in the analysis ... the bound is strongly constrained.





 $\sqrt{s}=250 \text{GeV}, L_{int}=250 \text{fb}^{-1}, P(e^-,e^+)=(-80\%,+30\%)$ 

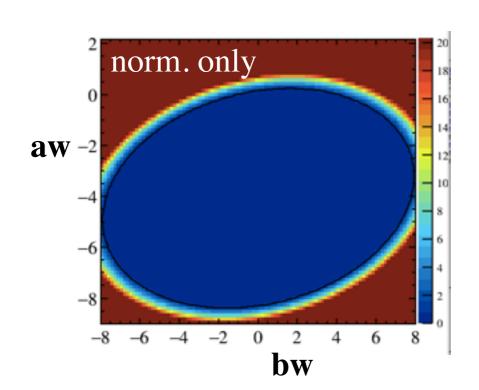


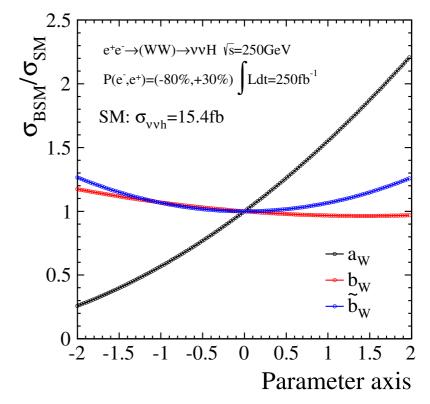
## WWH a-b

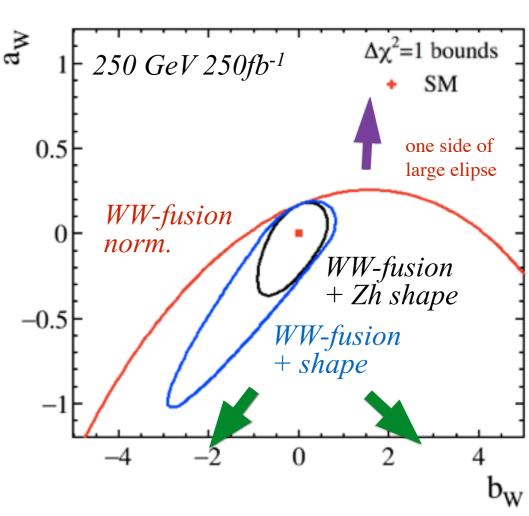
#### Normalization from WW-fusion

az make  $\sigma$  increase, but, this time, bz can not change it largely any btz can not adjust since any btz can increase  $\sigma$ . Thus, the bound is quickly restricted

For this direction both az make  $\sigma$  decrease. Both bz & btz has huge room to recover the SM value by increasing  $\sigma$ 







15

10

## WWH a-b 250 & 500 GeV

