

# FCAL Workshop in Tel Aviv

W. Lohmann, DESY

Why a  $e^+e^-$  Collider

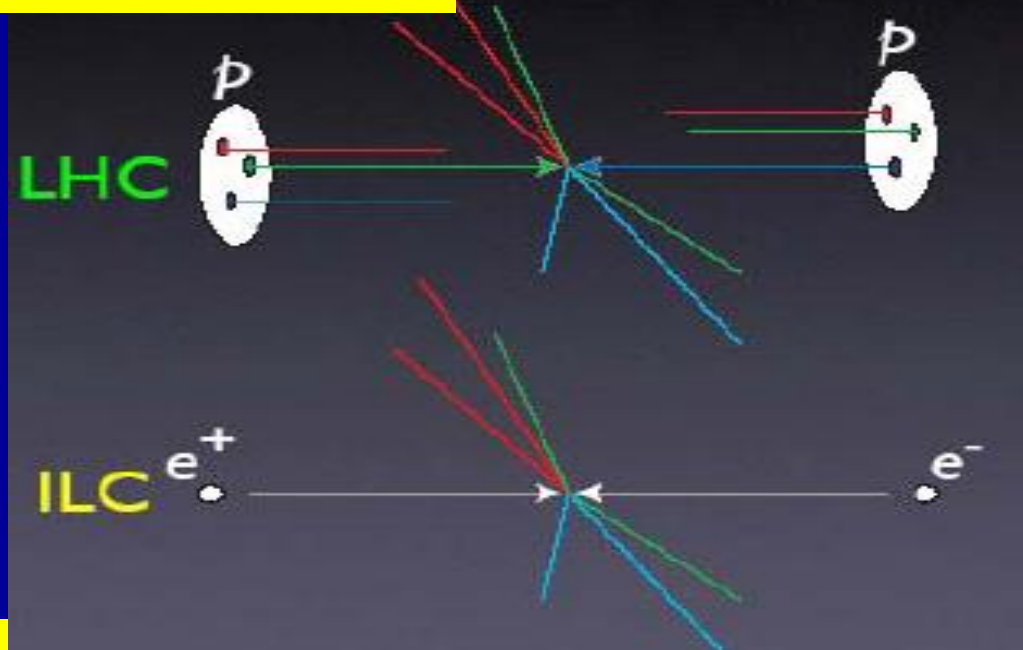
Physics essentials

The Snowmass Adventure

Where we are with FCAL

# Why $e^+e^-$

- Electrons are pointlike
- Energy tunable
- Polarised beams
- Clear events



## Accelerator Design

First stage: 90 - 500 GeV

Second stage: up to 1 TeV

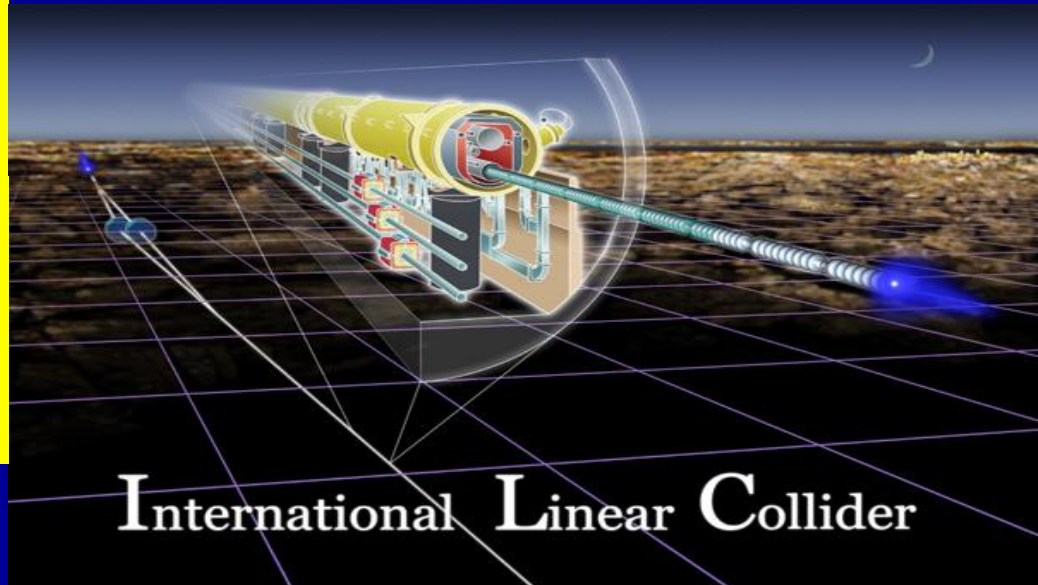
Luminosity:  $500 \text{ fb}^{-1} / \text{year}$

Cold (SC) Technology

Frequency: 5 Hz (trains)

About 3000 bunches per train

300 ns between bunches



International Linear Collider

# Physics essentials

Origin of Mass

Space-Time Structure

Dark Matter

Predict new particles or phenomena in the energy range 100 GeV - 1 TeV:  
The Terascale - the domain of the ILC!

# Origin of Mass

SM of particle physics:

Leptons and Quarks (Fermions,  $s=1/2$ ) form matter

Gauge Bosons ( $S=1$ , Photon, Z,  $W^+$ , Gluons) mediate Interactions

## Higgs Mechanism

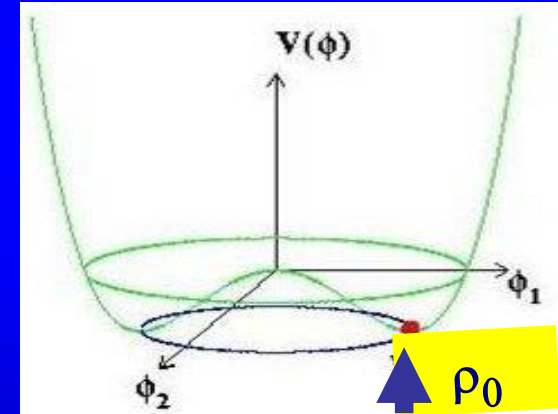
doublets under SU(2)

$$\Phi = \begin{pmatrix} \Phi_1 \\ \Phi_2 \end{pmatrix}$$

$$\mathcal{L}_H = \partial_\nu \Phi^\dagger \partial^\nu \Phi - \mathcal{V}(\Phi)$$

$$\mathcal{V}(\Phi) = -\mu^2 \Phi^\dagger \Phi + \lambda (\Phi^\dagger \Phi)^2$$

$$\Phi = \begin{pmatrix} 0 \\ \frac{1}{\sqrt{2}} \rho_0 \end{pmatrix}$$



Gauge Boson Masses

$$m_W^2 = \frac{1}{4} g_1^2 \rho_0^2 = \frac{e^2 \rho_0^2}{4 \sin^2 \theta_W}$$

$$m_Z^2 = \frac{1}{4} (g_1^2 + g_2^2) \rho_0^2 = \frac{e^2 \rho_0^2}{4 \sin^2 \theta_W \cos^2 \theta_W}$$

Fermion Masses

$$m_e = \frac{1}{\sqrt{2}} \rho_0 c_e$$

$$\rho_0^2 = \text{sqrt}(2) G_F$$

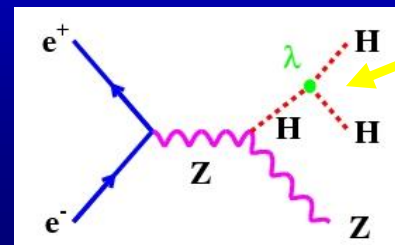
Unkonwn:

$$m_H^2 = 2\lambda\rho_0^2$$

$$\lambda\rho_0 h^3$$

$$\frac{1}{4}\lambda h^4$$

Higgs Field Potential,  $\lambda$



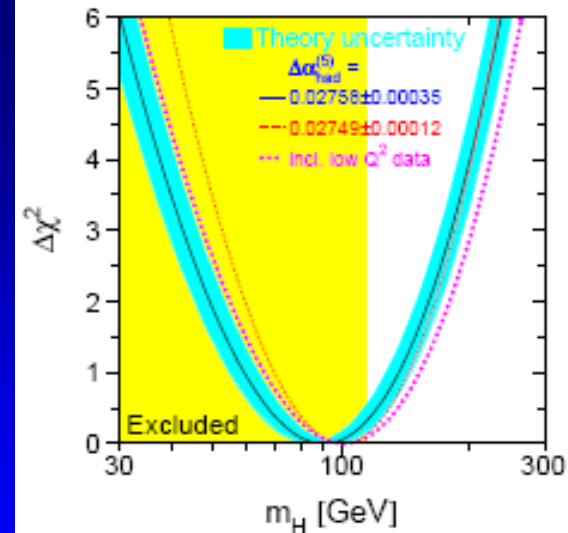
# Higgs Boson

## What we know about the Higgs Boson:

From LEP, SLD, Tevatron  
(Precision measurements)

$m_H = 91^{+45}_{-32} \text{ GeV}, < 186 \text{ GeV @ 95\% CL}$

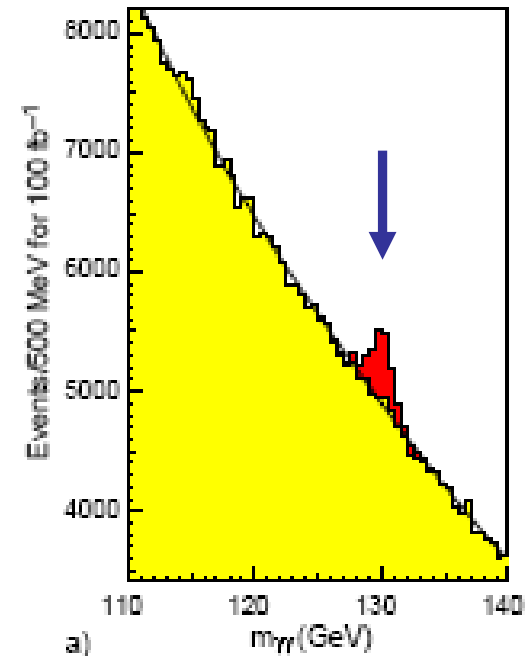
From LEP direct searches:  $m_H > 114 \text{ GeV}$



## What we may know in (a few) years:

LHC/Tevatron will discover  
a 'light' SM Higgs Boson

e.g. CMS  $H \rightarrow \gamma\gamma$   
 $\mathcal{L} = 100 \text{ fb}^{-1}$



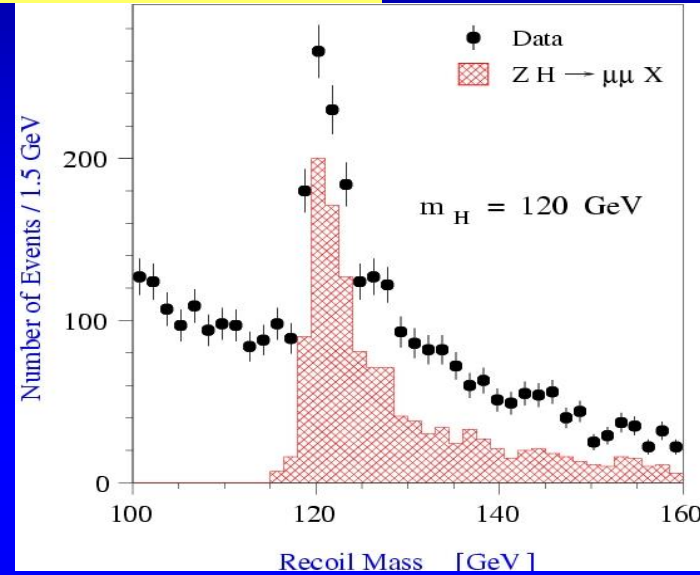
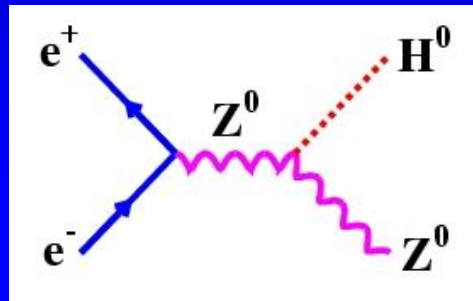
# Higgs Boson

## What we expect from ILC: Understand EWSB!

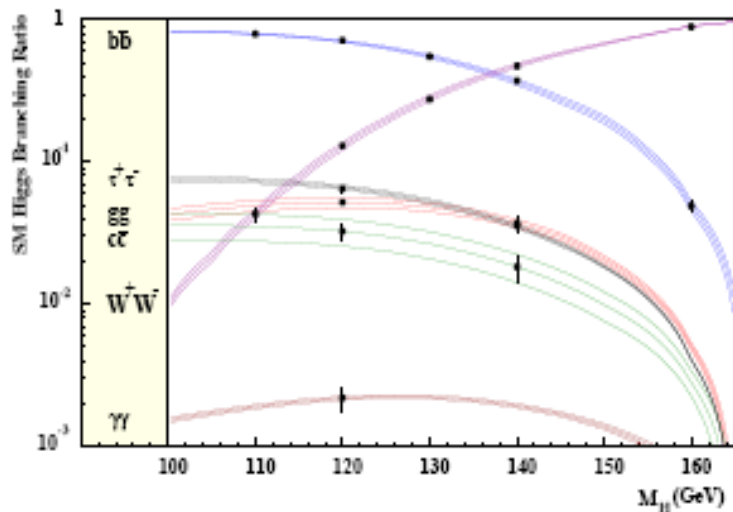
### Identification of the Higgs (Mass, Spin, Parity), Couplings

$e^+e^- \rightarrow ZH \rightarrow l^+l^-X$   
 ('golden physics channel'), with  $\delta(m_{l^+l^-}) \ll \Gamma_Z$

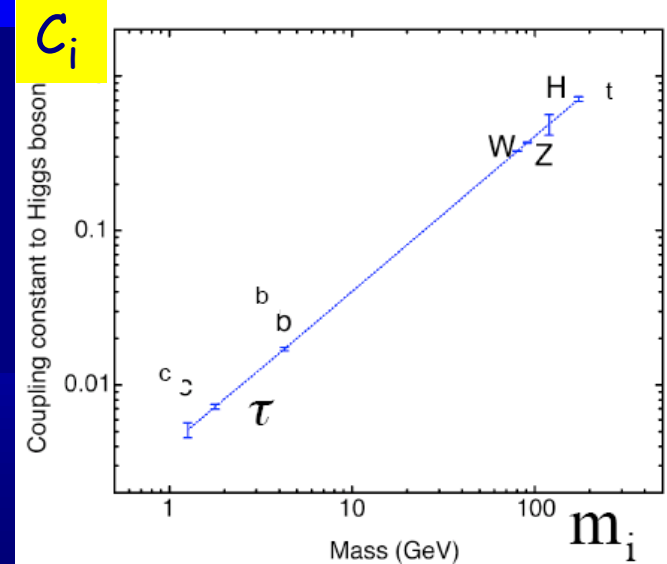
Mass accuracy  $\sim 40$  MeV  
 Momentum and jet energy resolution



### Branching fractions (couplings)



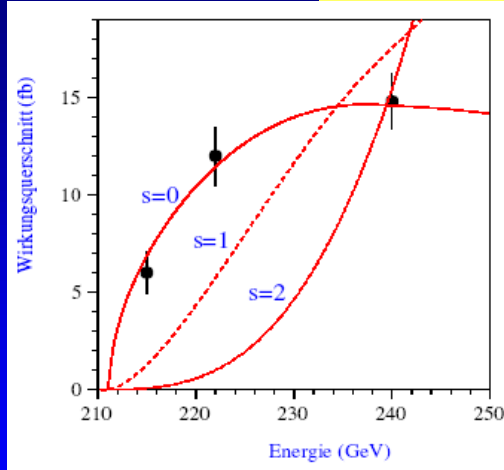
$m_H < 140$  GeV  
 $Z, W, b, \tau, c, t$   
 $m_H > 140$  GeV  
 $Z, W, t, b$   
 Flavour tagging



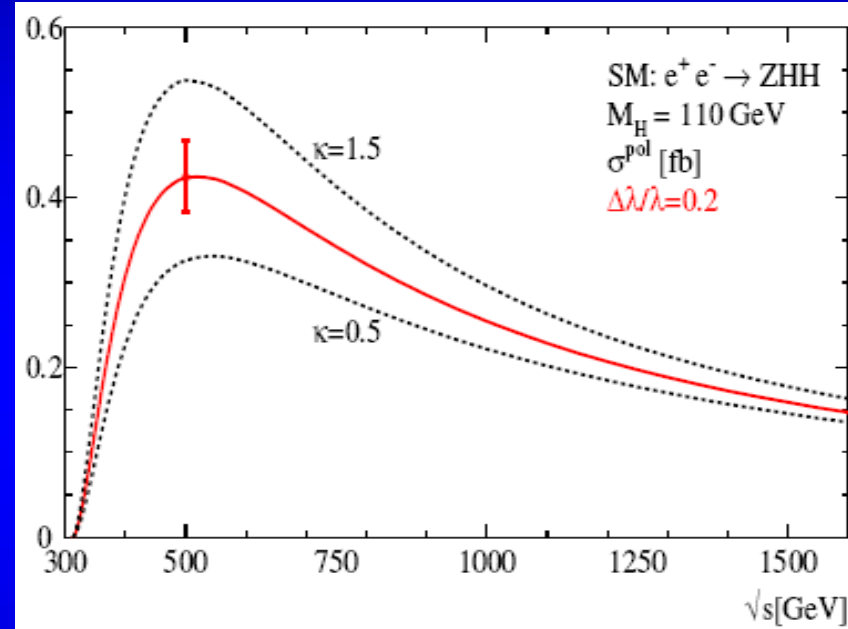
# Higgs Boson

Spin, Parity  
CP

b-tagging,  
 $\tau$ -tagging

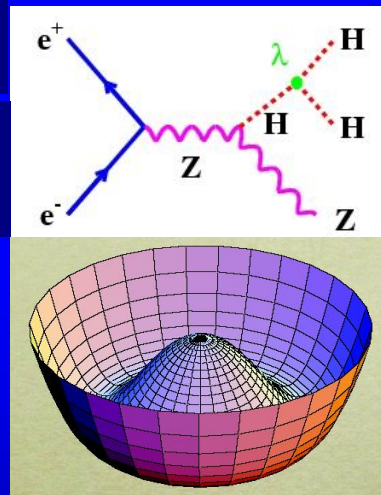


The Higgs boson would be the first elementary particle with spin 0!



Higgs Field Potential,  $\lambda$

Jet energy resolution,  
b-tagging, vertex charge



Beyond SM: more complex Higgs sector, e.g. MSSM

Two CP even states:  $h, H$  ( $m_h < 130$  GeV)

One CP odd state:  $A$

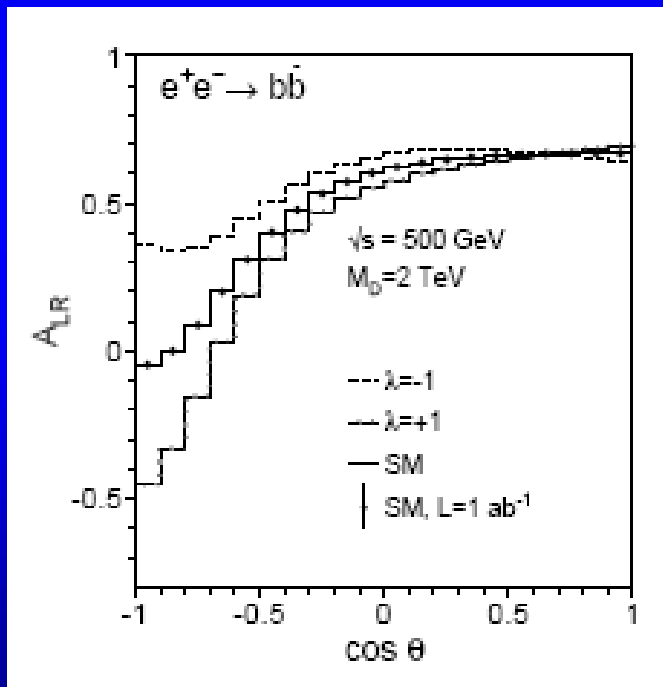
Two charged states:  $H^{\pm}$

# Space-Time Structure

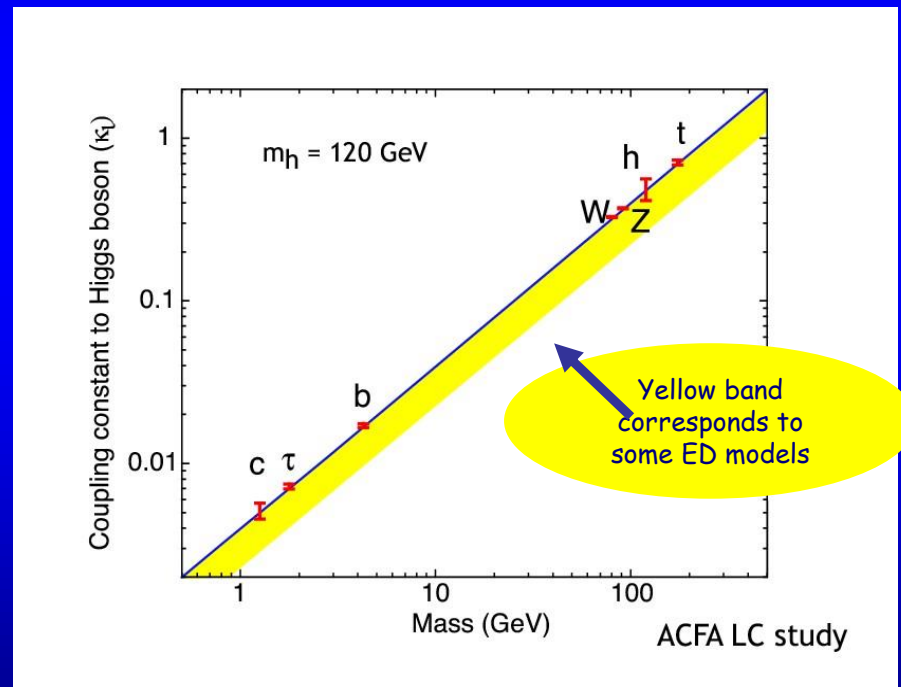
Extra Space Dimensions (Gravity extends to more than three Dimensions, the 'bulk'): Kaluza-Klein towers of states

$$e^+e^- \rightarrow f\bar{f}$$

Scalar Mode: Radion, mixing with the Higgs Boson



b-tagging, vertex charge



B, c-tagging,  
 $\tau$ -tagging

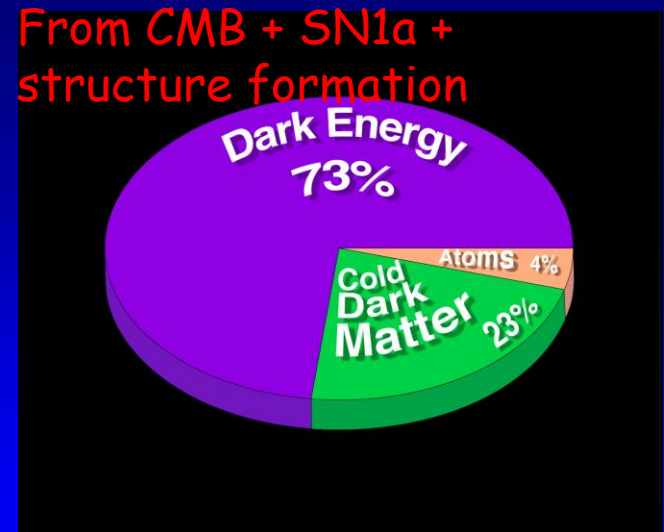


# Dark Matter

There is no Cold Dark Matter particle in the SM!

From Observational Cosmology:

Baryon	4 %
Dark Matter	23 %



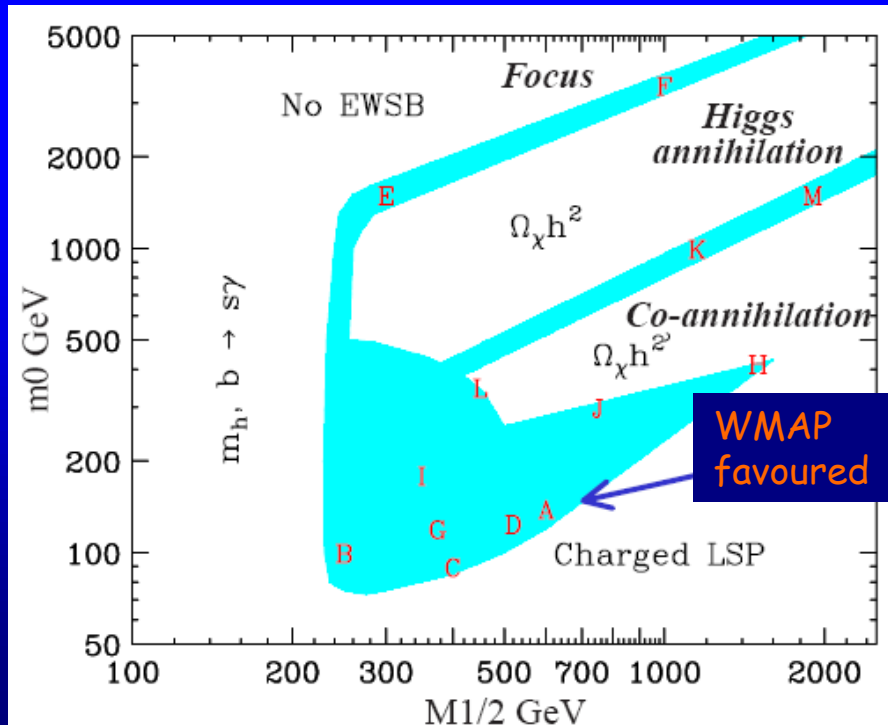
Supersymmetry provides CDM candidates, e.g.:

$$e^+e^- \rightarrow \tilde{\tau}^+ \tilde{\tau}^- \rightarrow \tau^+\tau^-\chi^0\chi^0,$$

$\chi^0$  is LSP

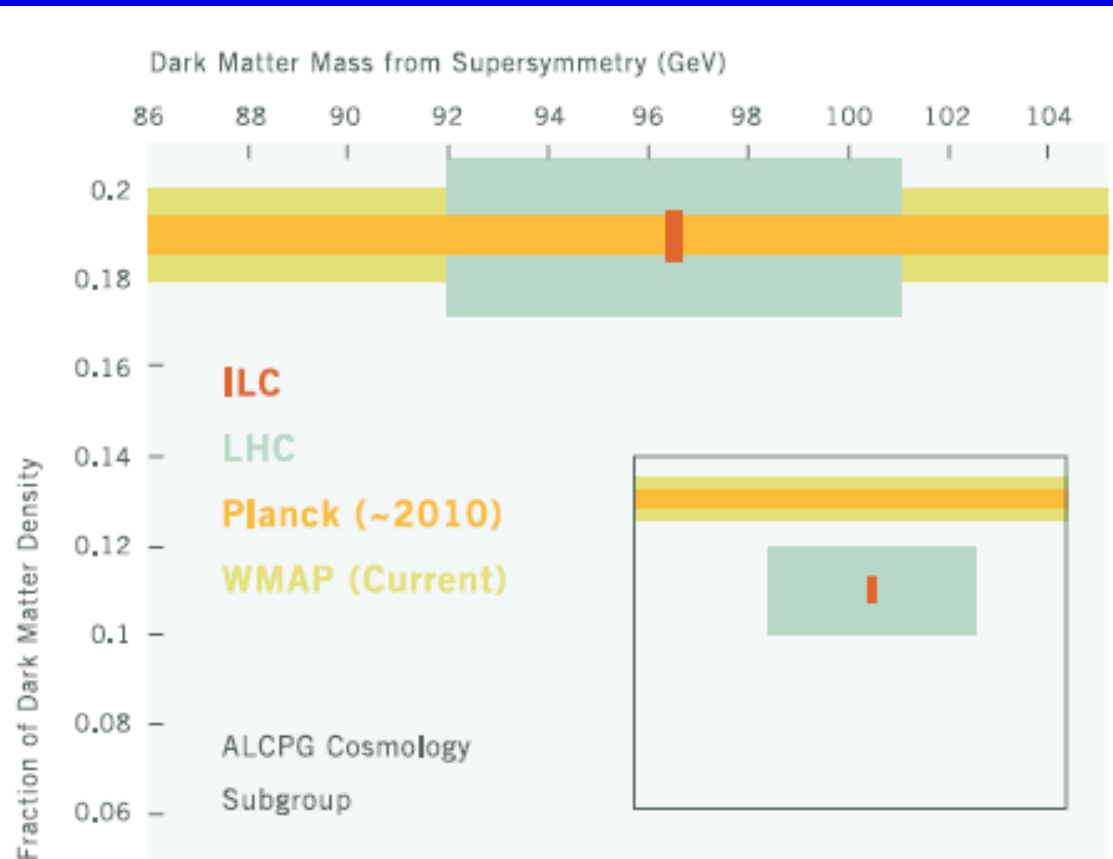
Small  $\Delta m$ , difficult to detect  
Large background from 4f events

Detector hermeticity



# Dark Matter

The target is to discover CDM particles, measure their mass and couplings and compare to observational cosmology

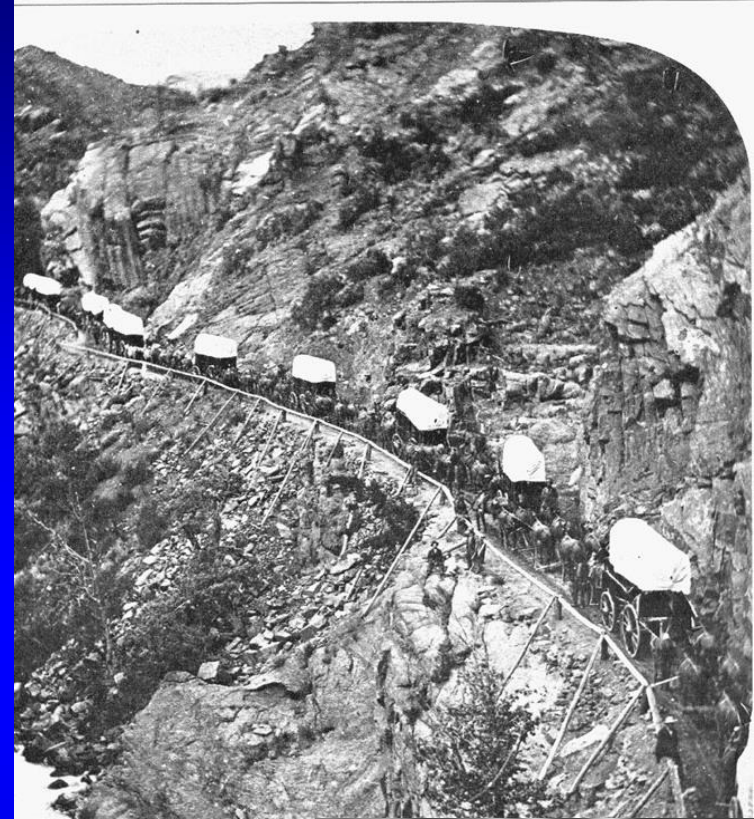


A possible scenario

# The Snowmass adventure

More than 750 physicists from around the world came to work together

A 'virtual' Lab, GDE is formed to manage the world-wide effort (Accelerator, Detector, Physics ..)  
Several working groups are formed,  
People from all parts of the world overtook clear responsibilities



←  
The Lab (GDE) has a director, Berry Barish (and regional directors for Europe, NA and Asia)  
Europe: B. Foster



# The GDE Plan and Schedule

2005

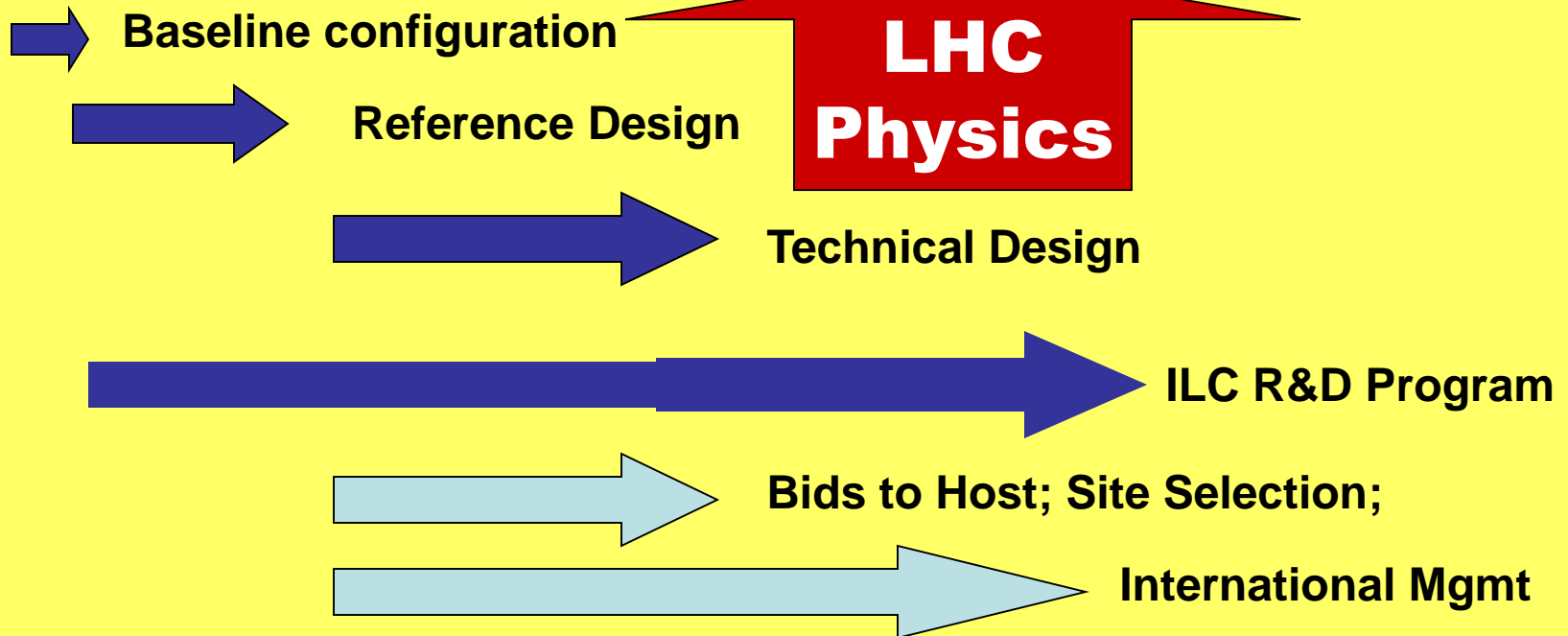
2006

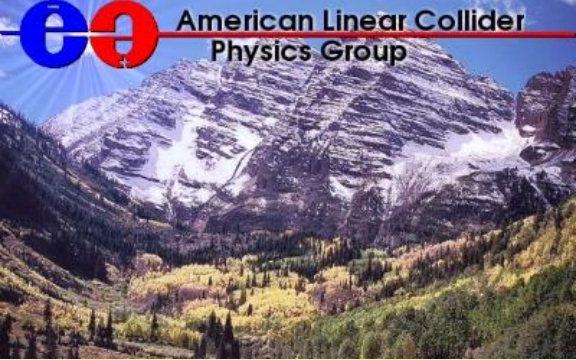
2007

2008

2009

2010





# GDE Timeline

## machine

end of 2005

Baseline Configuration Document

end of 2006

Develop Reference Design Report

## detectors

R&D Report

Detector Outlines

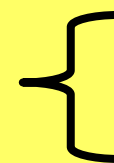
(Mar, 2006)

→ Detector Concept Report

3 volumes: i.) RDR (machine)

ii.) Detector Concept Report

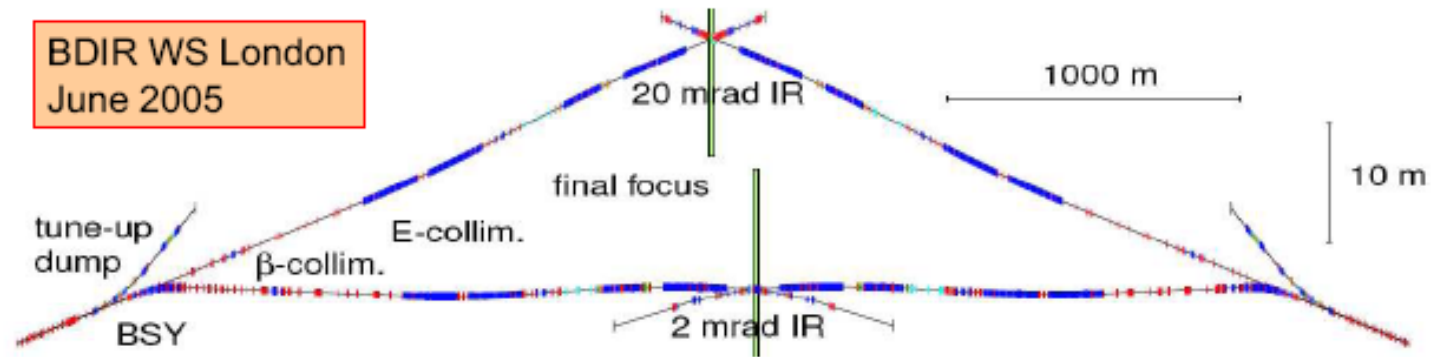
iii.) Exec Summary



Physics  
Detectors

# Interaction Region

to Design at RHUL



Full optics for all beamlines,  
2 mRad and 20 mRad designs explored in detail,  
up/downstream instrumentation present for both IRs.

## Two Detectors, because:

- Confirmation and redundancy
- Complementary Collider options
- Competition
- Efficiency, reliability
- Historical lessons

# ILC-LHC

- The Success of LHC will be a big boost for our field
- We are going ahead aggressively ahead to elaborate the case for the ILC, following our schedule
- Once we have collisions at the ILC an exciting Synergy with LHC will realized

## Historic lesson:

Discovery	Collider	$L_{peak}$ ( $\text{cm}^{-2}\text{s}^{-1}$ )	1st collisions	Observation (Expt.)	Time lag
$W^\pm$	CERN $Spp\bar{p}S$	$1.7 \times 10^{29}$	Aug 1981	Jan 1983 (UA1)	1.5 yr
$Z^0$	CERN $Spp\bar{p}S$	$1.7 \times 10^{29}$	Aug 1981	Jun 1983 (UA1)	2 yr
top	FNAL Tevatron	$2 \times 10^{30}$	Feb 1987	Mar 1995 (CDF)	8 yr
Higgs	CERN LHC	$10^{33} - 10^{34}$			

ILC has a compelling physics case

The accelerator will be SC (great success for the TESLA collaboration)

The Community made an important step to an 'International Organisation'

The R&D program for the ILC detector is exciting (Don't miss it)



# Where are we with FCAL

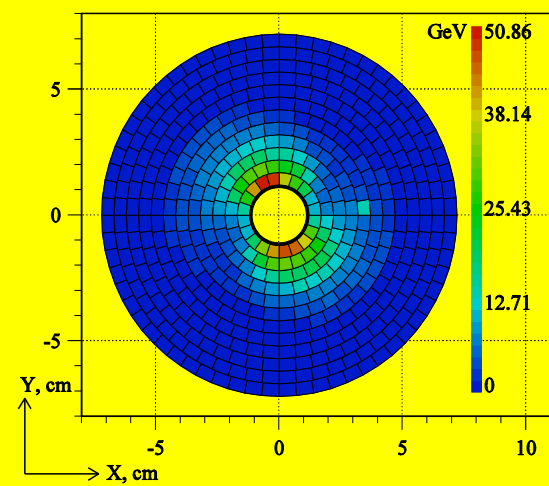
We worked out an advanced design for zero or small crossing angle; This was accepted and acknowledged in Snowmass

Several talks by H. Abramowicz, A. Elagin, P. Bambade (V. Drugakov), myself in Snowmass

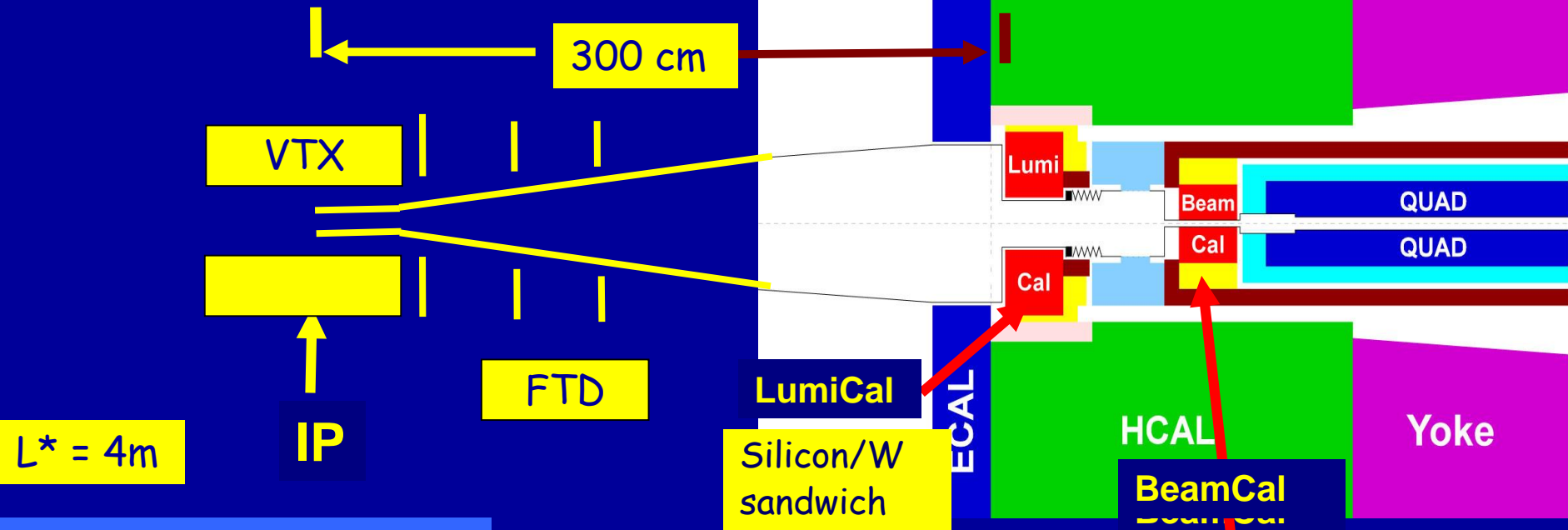
# Very Forward Detectors

- Detection of Electrons and Photons at very low angle – extend hermeticicity
- Measurement of the Luminosity with precision ( $<10^{-3}$ ) using Bhabha scattering

Beamstrahlung  
Depositions:  
20 MGy/year  
Rad. hard sensors  
e.g. Diamond/W  
BeamCal



- Fast Beam Diagnostics



LumiCal:  $26 < \theta < 82$  mrad  
BeamCal:  $4 < \theta < 28$  mrad  
PhotoCal:  $100 < \theta < 400$   $\mu$ rad

R&D for ILC (DESY PRC R&D 02/01):  
Instrumentation of the Very Forward Region of the ILC Detector

## What has to be done:

We need a similar design for 20 mrad crossing angle

- repeating the studies on critical parameters (as done by achim)
- feasibility of beam diagnostics (magnetic field!)
- Studies of background

We have to understand Bhabha phenomenology

- Status of the theory, radiative effects and detector performance
- Comparison of different generators (BHLUMI, SamBha....)
- Background studies
- Optimised segmentation/structure for LumiCal
- Realistic readout scheme

# What has to be done:

## Sensor and Readout

- Continuation of diamond studies (more samples with promising diagnostics, linearity, homogeneity, high radiation doses.)
- Si sensor studies (to learn to work with them).
- Si sensor radiation test.
- Assembly of full sensor planes → prototype test.
- Readout electronics design for the prototype  $O(1000)$  channels.
- Concept Design for the 'fast readout' and fast diagnostics (related to Eurotev).

# This meeting at TAU

- Reports on all topics mentioned

- Discussions on 'critical' issues  
'canonic' geometry in GEANT4

- sensor development

- others

- preparation of the vienna workshop

- EUDET, INTAS ...

Lets make a WORKshop and see at the  
End where we'll go