# The use of beam polarisation in the search for dark matter at the ILC





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# Probing Dark Matter with e<sup>+</sup>e<sup>-</sup>

## **Mono-photon signature**

The mono-photon signature is considered to be the most general way to look for DM particle production in future  $e^+e^-$  colliders.



DM can be pair produced in the  $e^+e^-$  collisions via exchange of a new mediator particle, which couples to both electrons (SM) and DM states

This process can be detected, if additional hard photon radiation from the initial state is observed in the detector...





## Talk to be given at SPIN'2021

The use of beam polarization in the search for dark matter at the ILC

The International Linear Collider has the opportunity to discover dark matter particles in the monophoton signature  $e^+e^- \rightarrow \gamma + (missing)$ . The sensitivity to new physics in this channel depends crucially on the control and correct estimation of backgrounds. This talk will explain how beam polarization can be used at the ILC to improve the reach of this search.

Impact of polarisation already included in the existing studies:

- M. Habermehl, Dark Matter at the International Linear Collider, PhD thesis, University of Hamburg, 2018, https://bib-pubdb1.desy.de/record/417605?
- M.Habermehl, M.Berggren, J.List, Phys. Rev. D 101 (2020) 075053, arXiv:2001.03011
- J.Kalinowski, W.Kotlarski, K.Mekala, P.Sopicki, A.F.Z., arXiv:2004.14486, arXiv:2107.11194 for details see also May 12th presentation: https://agenda.linearcollider.org/event/9223/

#### But maybe some more information can be extracted?





Beam polarisation can be used to suppress/control background levels combination of different polarisations helps to reduce the systematics



It can also enhance mono-photon signal (depending on the scenario).



#### arXiv:2001.03011

Beam polarisation can be used to suppress/control background levels combination of different polarisations helps to reduce the systematics



Mass scale limits improve significantly with proper choice of polarisation.



#### arXiv:2001.03011

Beam polarisation can be used to suppress/control background levels combination of different polarisations helps to reduce the systematics



Polarisation choice depends on the BSM scenario...



#### arXiv:2001.03011

Beam polarisation can be used to suppress/control background levels combination of different polarisations helps to reduce the systematics



 $\Rightarrow$  combined analysis of data taken with different polarisations

## Light mediator exchange



**Cross section limits** for events with tagged photon

arXiv:2107.11194

as a function of mediator mass,  $M_Y$ for fixed  $\Gamma/M = 0.03$ 

Dirac fermion DM,  $m_{\chi} = 50 \text{ GeV}$ 

Vector





#### Impact of polarisation reduced for small mediator masses

A.F.Żarnecki (University of Warsaw)

Polarisation in the search for dark matter

# Light mediator exchange



**Cross section limits** for events with tagged photon arXiv:2107.11194

as a function of mediator mass,  $M_{\rm Y}$  for fixed  $\Gamma/M=0.03$ 

V–A coupling

Dirac fermion DM,  $m_{\chi} = 50 \, \text{GeV}$ 

V+A coupling



Impact of polarisation reduced for small mediator masses

# Light mediator exchange



**Cross section limits** for events with tagged photon arXiv:2107.11194

as a function of mediator mass,  $\ensuremath{\mathsf{M}}_{\ensuremath{\mathsf{Y}}}$  for vector mediator

 $\Gamma/M = 0.01$ 

Dirac fermion DM,  $m_{\chi} = 50 \text{ GeV}$ 

 $\Gamma/M = 0.5$ 



Impact of polarisation reduced for narrow mediator

A.F.Żarnecki (University of Warsaw)

Polarisation in the search for dark matter



## Systematic uncertainties

following ILD study: Phys. Rev. D 101, 075053 (2020), arXiv:2001.03011

Considered sources of uncertainties:

- Integrated luminosity uncertainty of 0.26% uncorrelated between polarisations
- Luminosity spectra shape uncertainty correlated between polarisations
- Uncertainty in neutrino background normalisation of 0.2% (th+exp) correlated between polarisations
- Uncertainty in Bhabha background normalisation of 1% (th+exp) correlated between polarisations
- Uncertainty on beam polarisation of 0.02–0.08% correlated for runs with same beam polarisation at ILC
- $\Rightarrow$  11 nuisance parameters in the model fit



### **How important is the polarisation uncertainty!?** Radiative cross section limits (for events with tagged photon) as a

function of the polarisation uncertainty scaling factor

 $\Gamma/M = 0.03 \qquad \qquad \Gamma/M = 0.10$ 



#### Systematic uncertainties relevant only for heavy mediator exchange



How important is the polarisation uncertainty!? Limits on the EFT mass scale  $\Lambda$  vector mediator, Dirac fermion DM as a function of the polarisation uncertainty scaling factor



Significant impact of polarisation even with much higher uncertainty  $\Rightarrow$  polarisation constrained by the data itself