

Neutron Identification with Plastic Scintillators



PRISMA+

DETECTOR LAB

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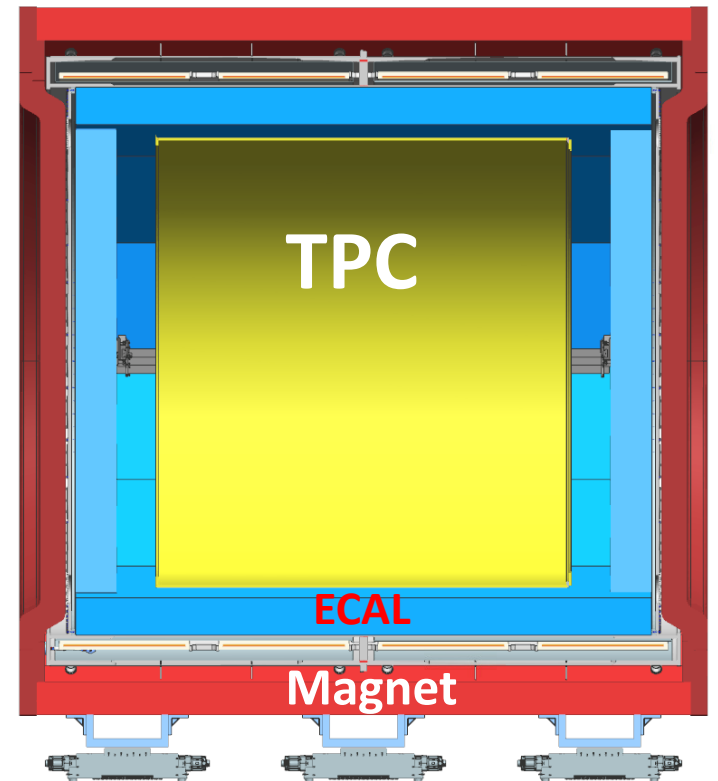


OUTLINE

- Motivation
- What is Neutron ID?
- Methodology
- Old setup + Results
- Improved setup + first Results
- Summary

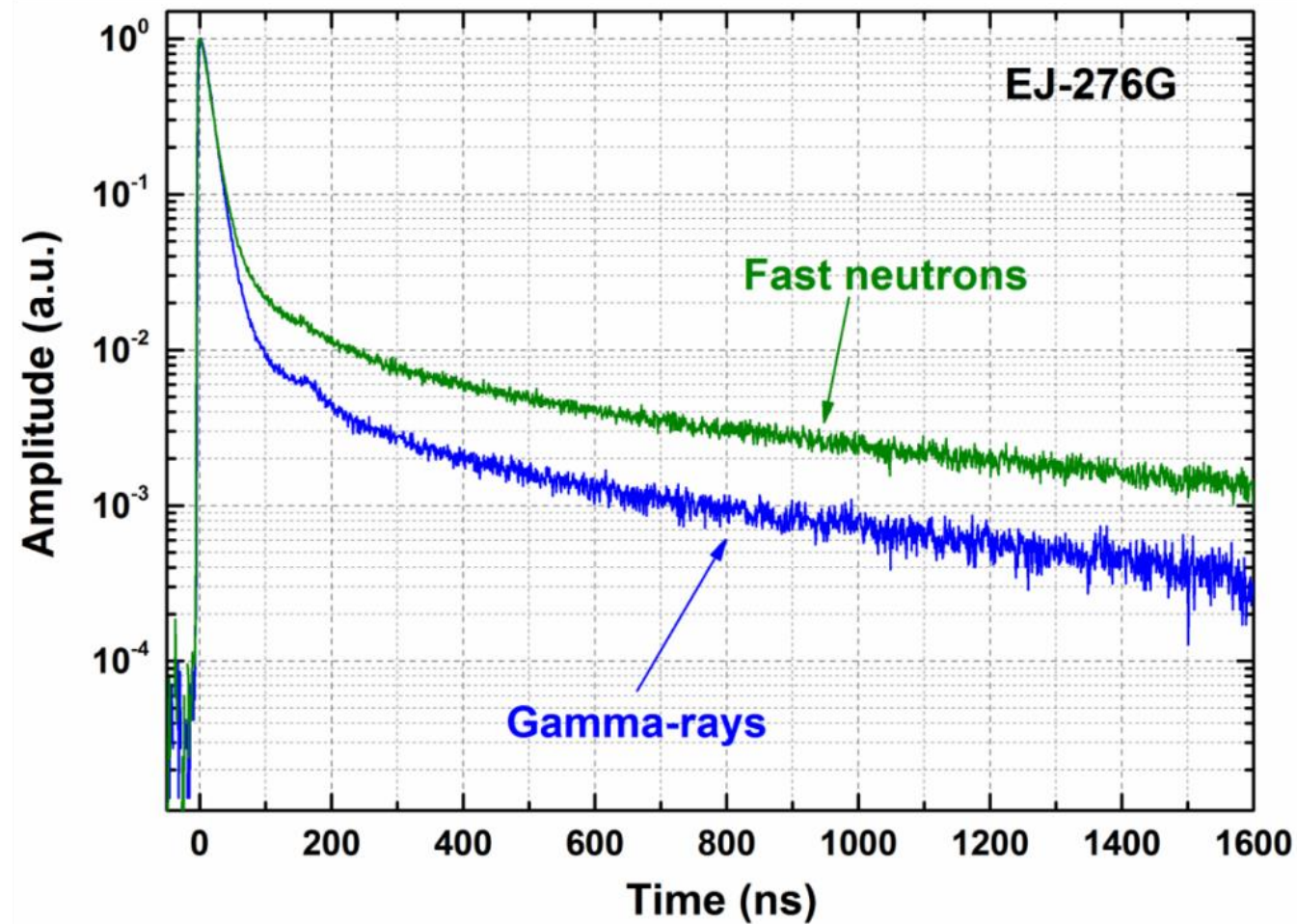
MOTIVATION

- Can we use AHCAL-style design with pulse shape discrimination (PSD) to identify gammas and neutrons?
- Case study: DUNE ND-GAr
 - TPC measures charged particles
 - ECAL can detect neutral particles but not identify (**so far**)
- Special plastic scintillators are optimized for Neutron ID (e.g. EJ-276G)

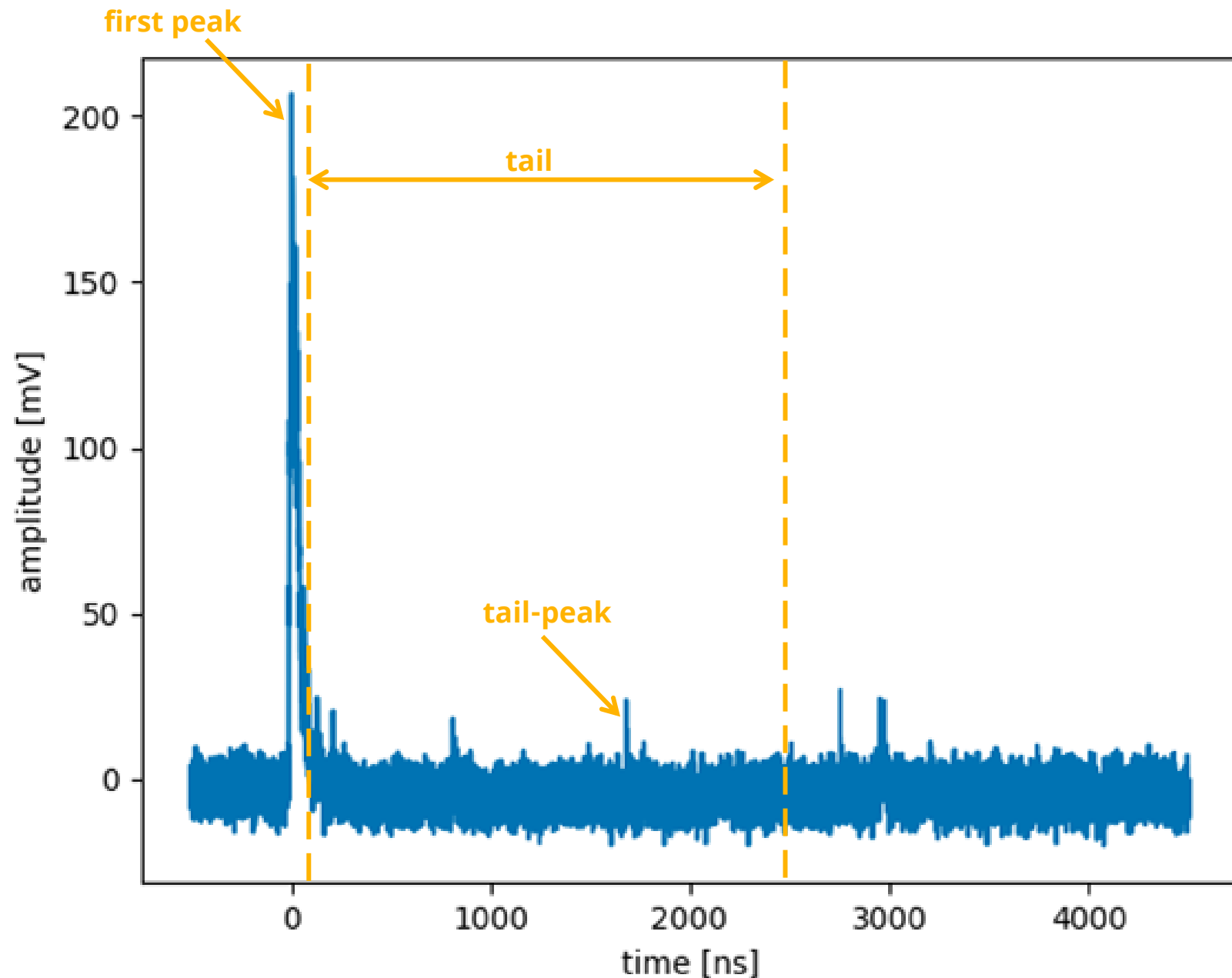


WHAT IS NEUTRON ID?

- Neutron ID: discrimination of incident neutrons by scintillation signature
- Classic process:
 1. Ionizing particles excite scintillator
 2. Light emission at deexcitation
- Neutron:
 1. Neutron scattering on nucleus
 2. Proton recoil
 3. Ionization
- Processes differ in photon timing



particle	Short component Decay const. [ns]	Long component 1 Decay const. [ns]	Long component 2 Decay const. [ns]
Gamma	13 ± 1	110 ± 10	800 ± 80
Neutron	14 ± 1	95 ± 10	800 ± 80



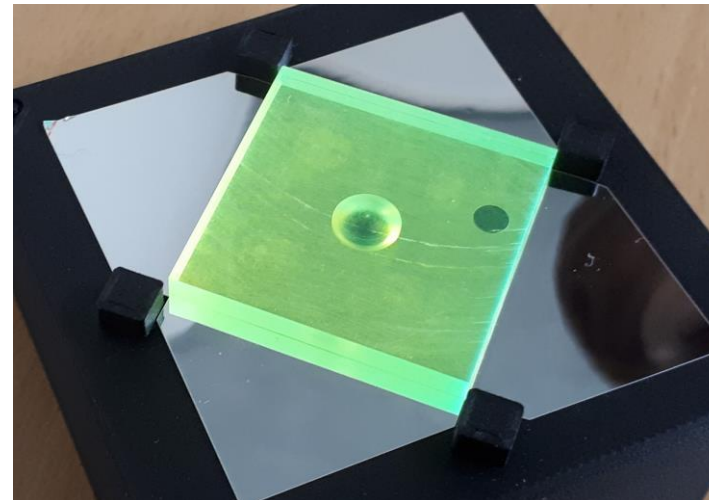
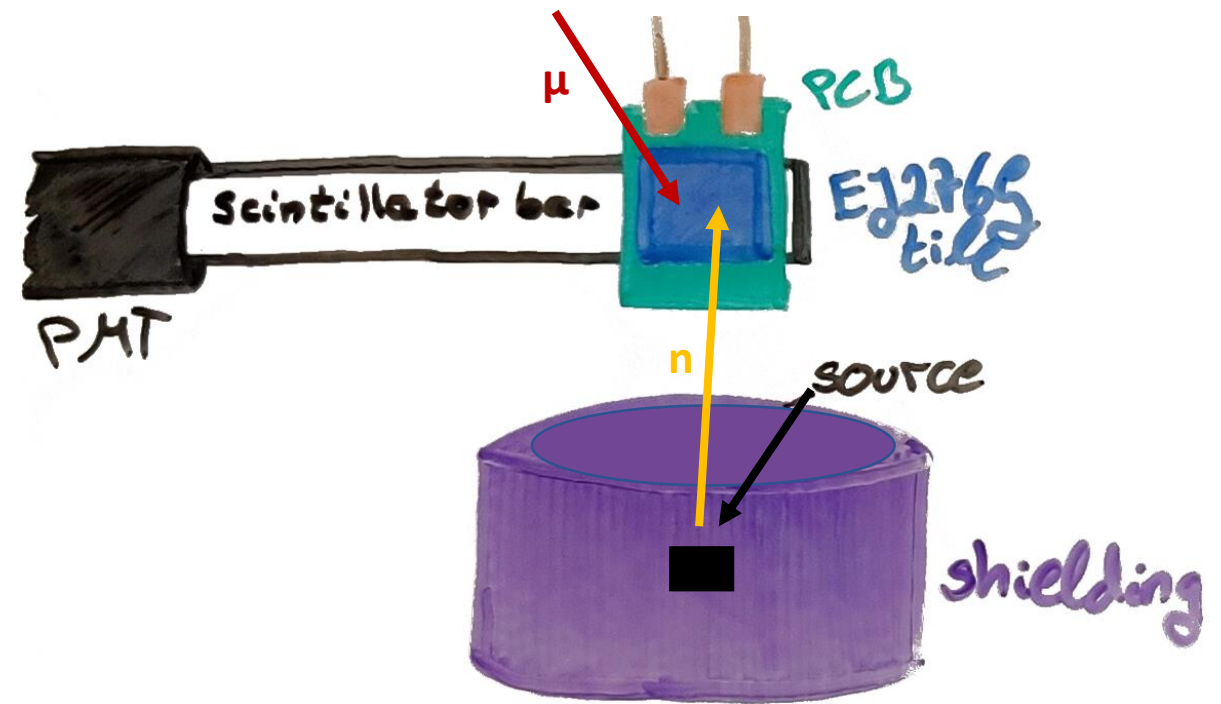
METHODOLOGY

- Compare immediate and delayed light output using full waveform
- In practice: compare
 - Height of initial peak
 - Number of tail peaks
 - Timing of tail peaks

Tail Start:	4 x short component 4 x 15 ns = 60 ns
Tail Stop:	3 x long component 3 x 800 ns = 2400 ns

SETUP

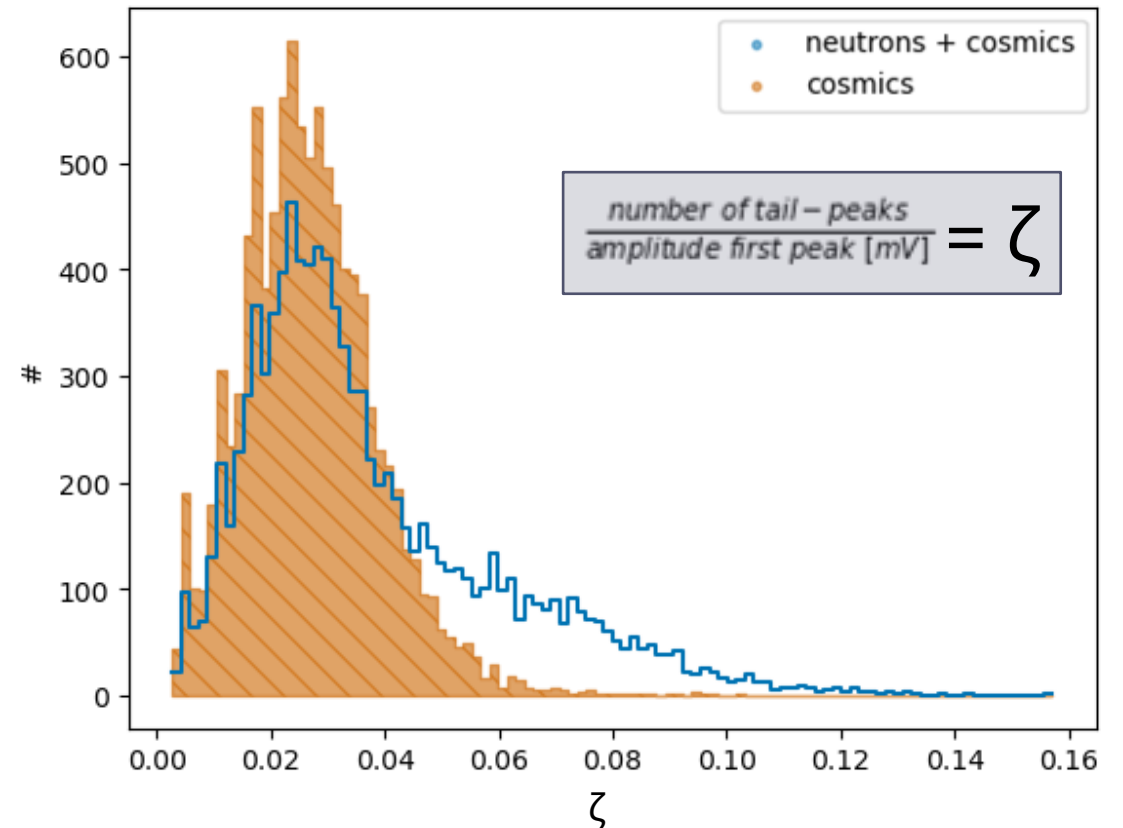
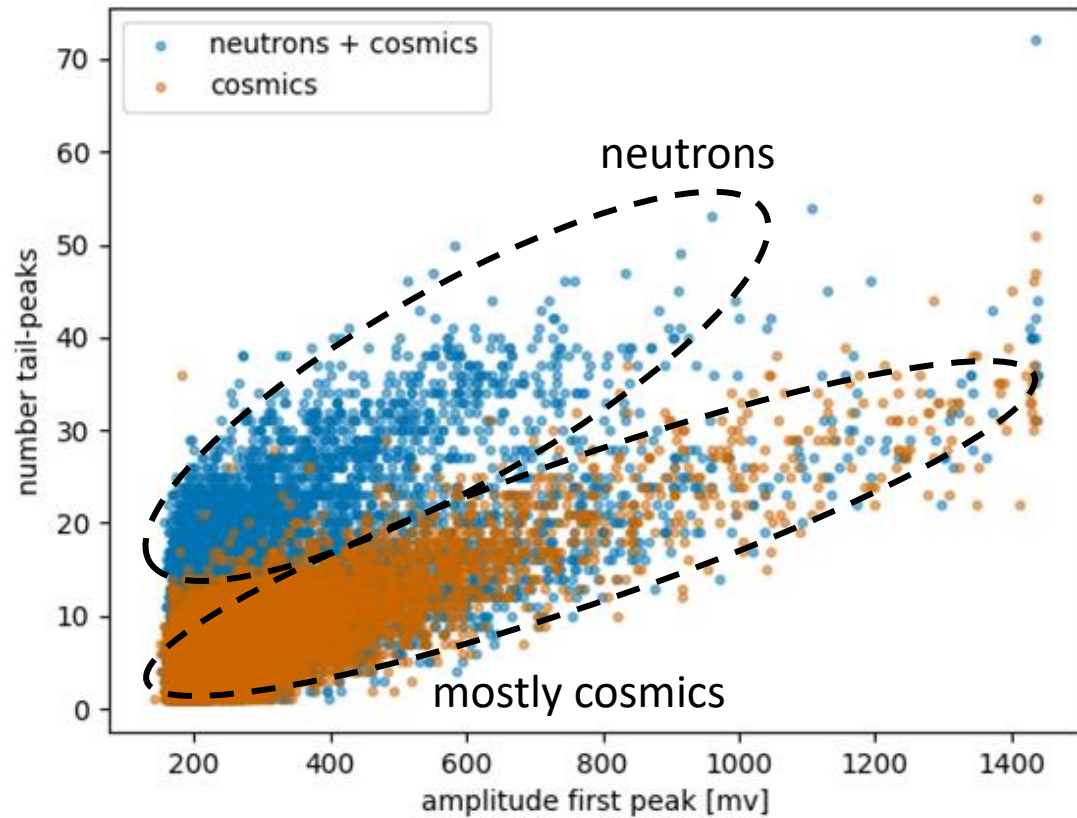
- How do we get neutrons and gammas?
 - Neutrons from shielded AmBe source
 - Cosmic muons used to emulate gammas (similar scintillation signature expected)
- SiPM-on-tile configuration similar to tiles on HBU
 - 30 x 30 x 3 mm EJ-276G tile
 - S13660-1325PE SiPM
- Signal amplified externally and readout by oscilloscope
- Full waveform saved -500 ns to 4500 ns



DATA

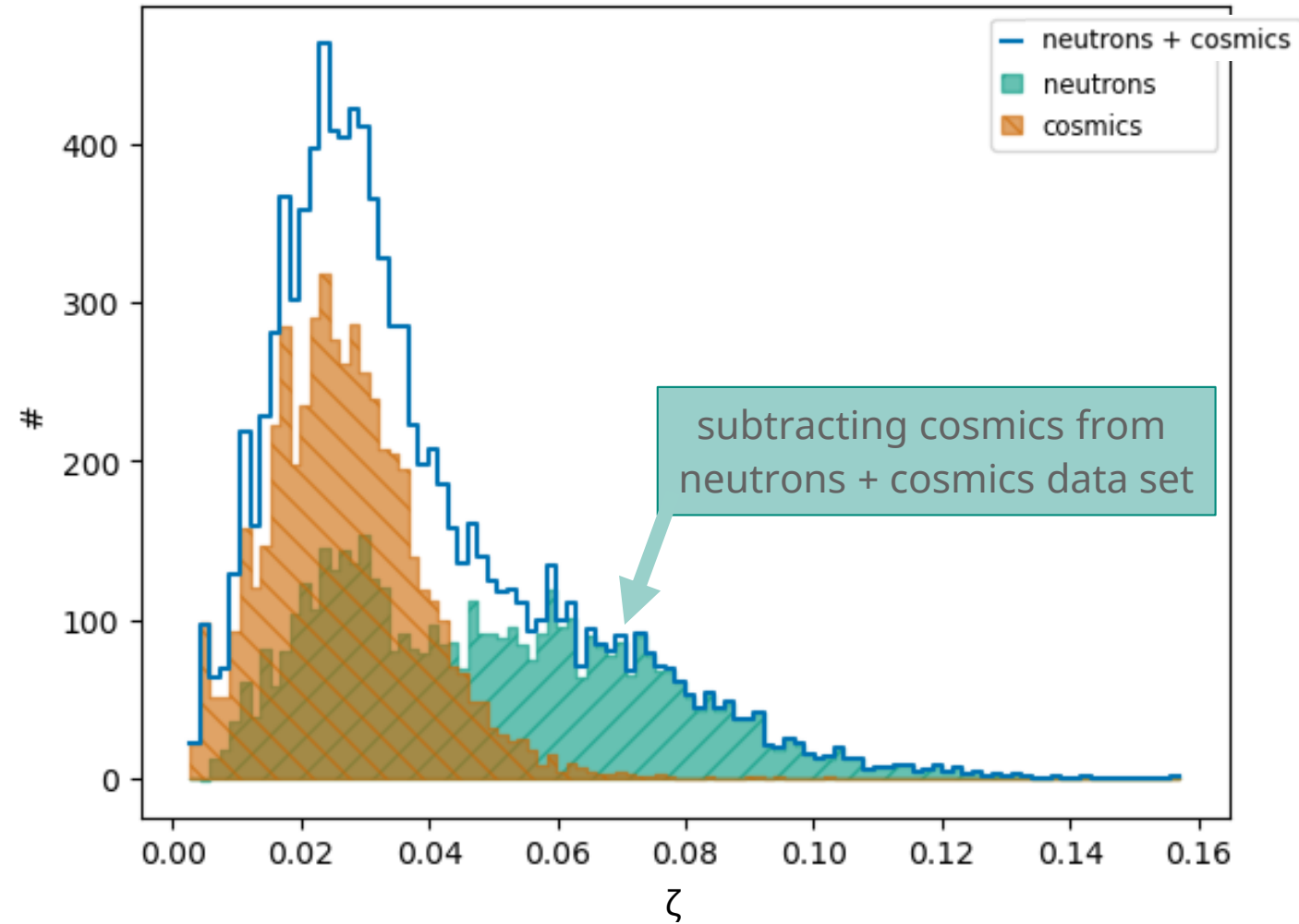
■ Muon + neutron samples

Peak height vs. number of tail peaks per event



DISCRIMINANT

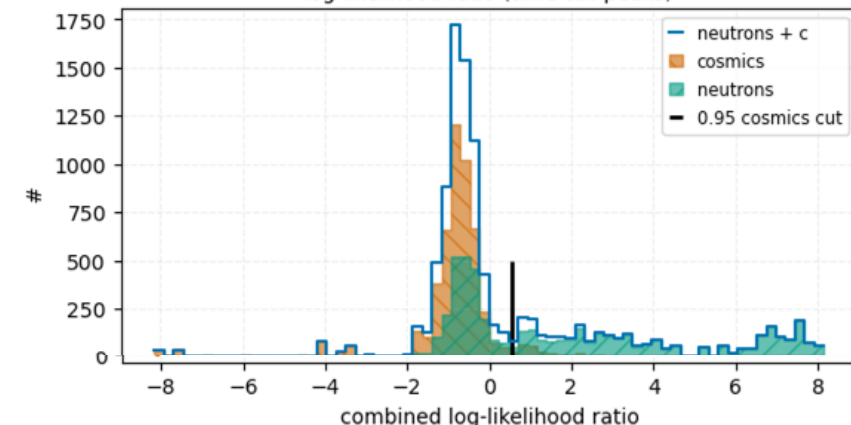
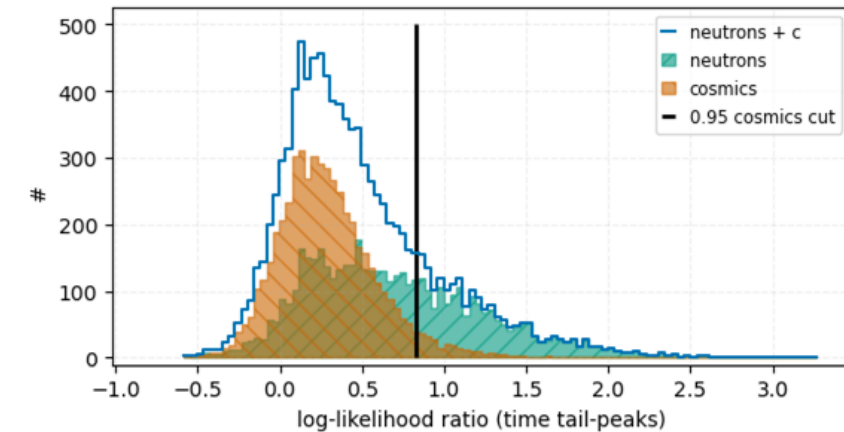
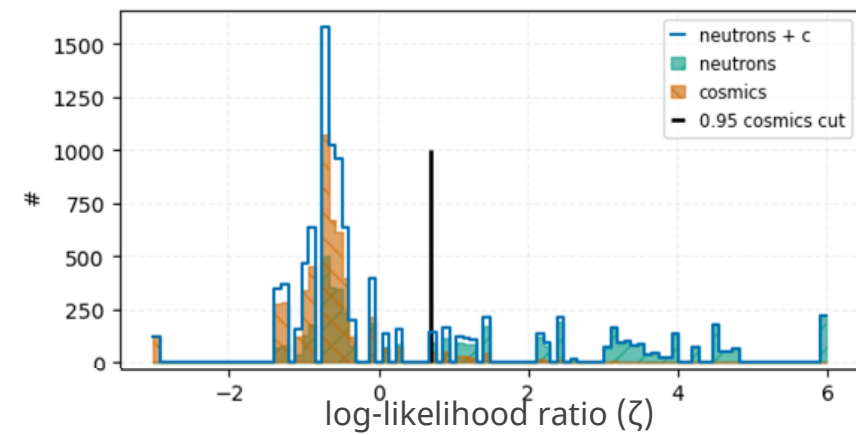
- Data runtime normalized
- Neutron data extracted by subtracting cosmics
- Discrimination efficiency determined based on resulting neutron shape



RESULTS

Three ways to perform Neutron ID:

- *Histogram bin height method*
 - Take log-likelihood for each ζ point
 - Bin height shows probability for each possibility
- *Time distribution Method*
 - Fit exponential to tail peak distribution
 - Log-likelihood based on time distribution fits
- *Combine both methods*
 - Good separation achieved

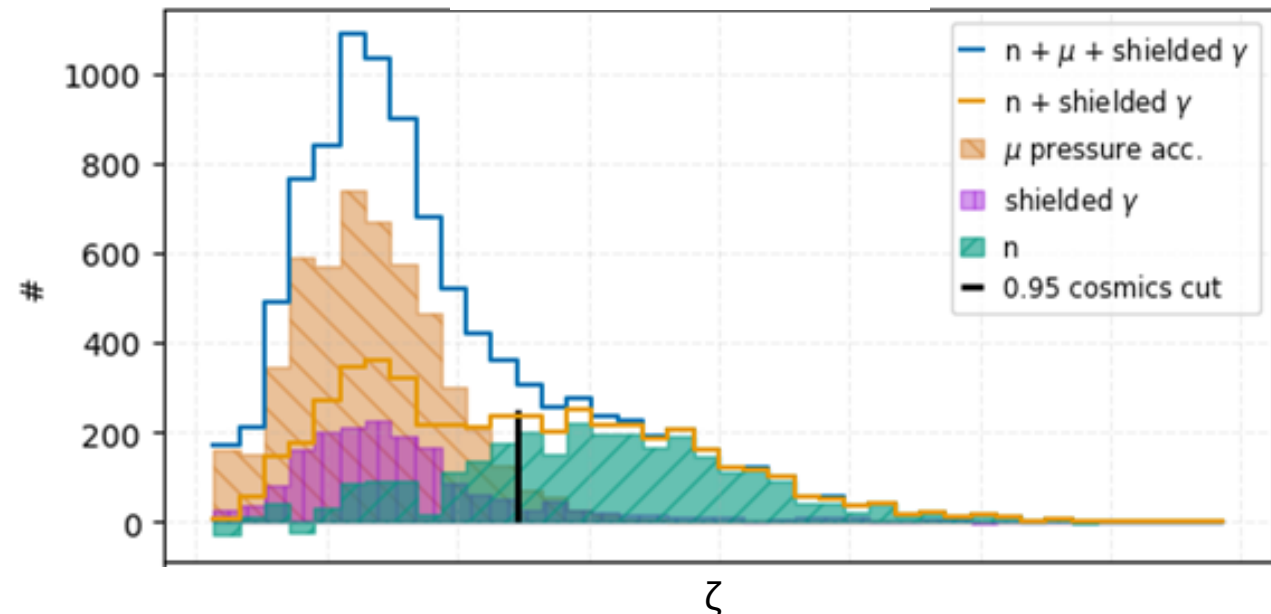
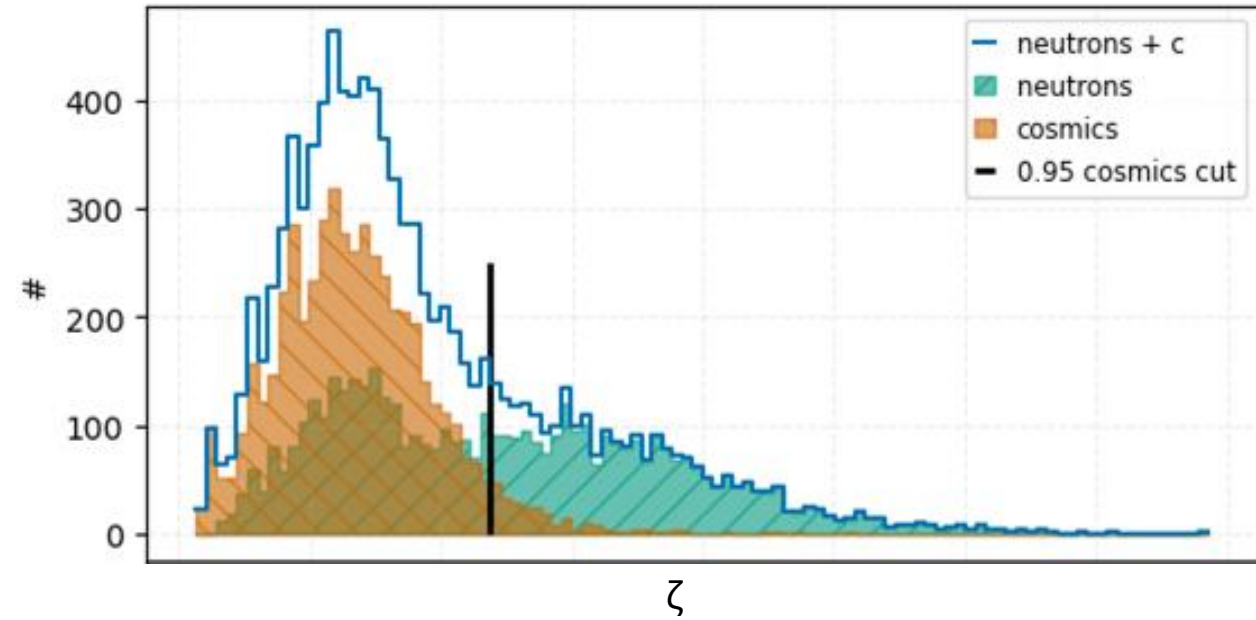


DISCUSSION

- What is the left peak in neutron distribution?
 - Suspicion: gammas penetrating shielding
 - Validation: measurement without shielding
- Neutron distribution corrected
 - Looks good
- Muons indeed look like gammas
 - Which is also good!

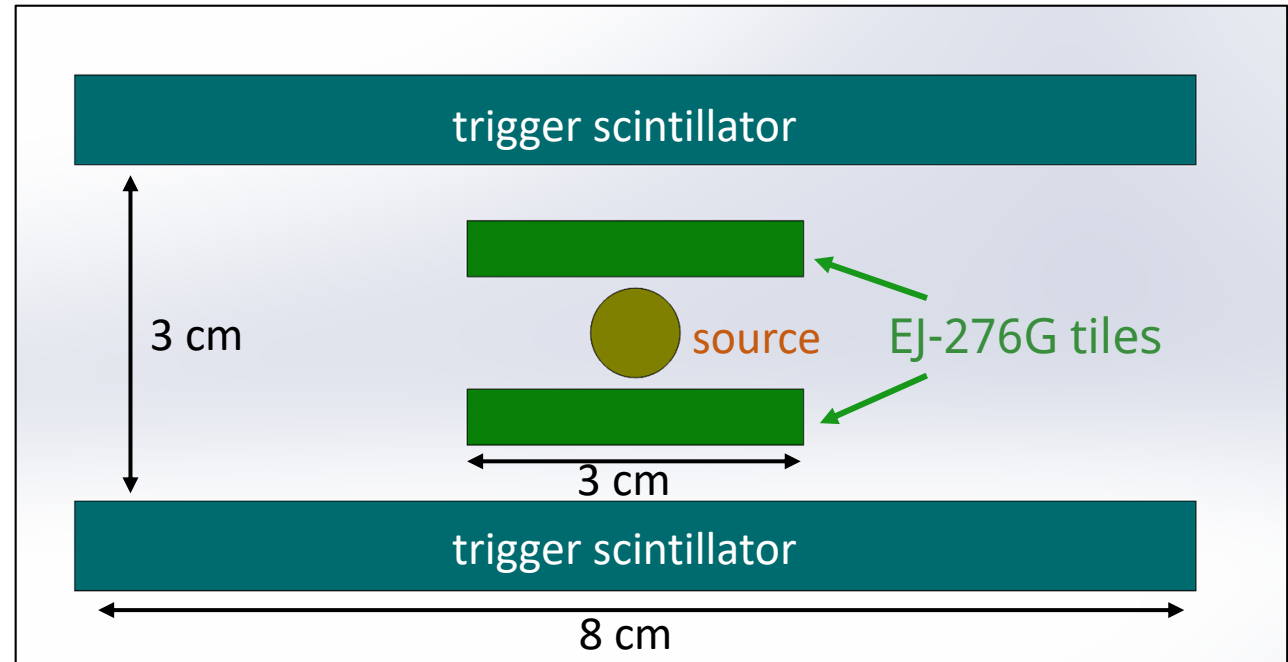
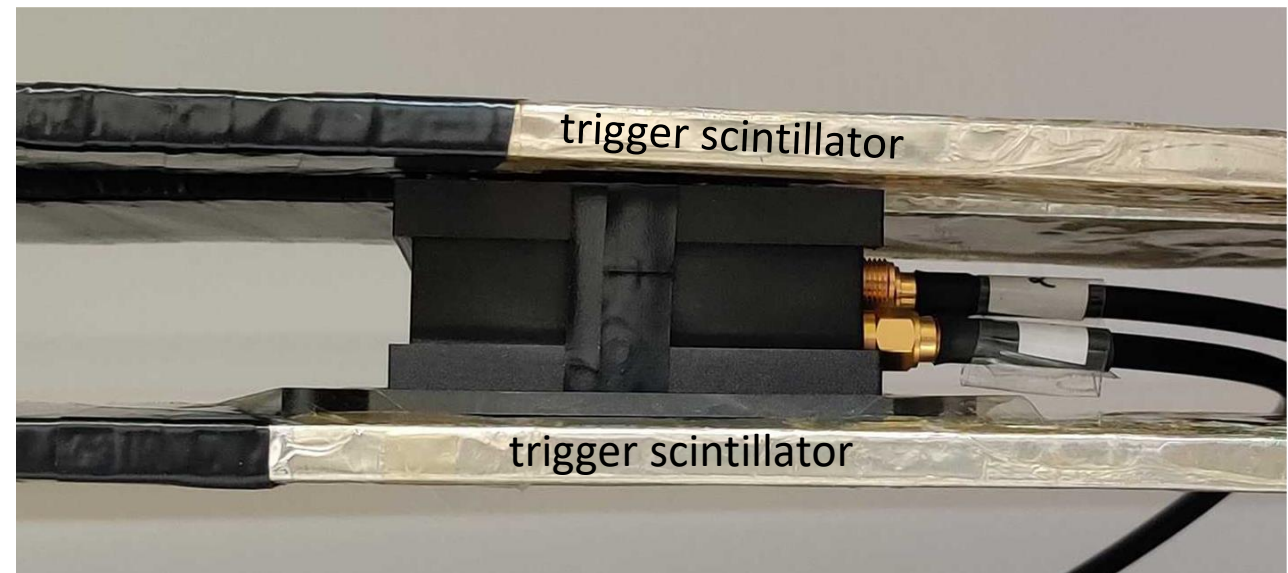
Next steps:

- Separation not good enough
 - SiPM upgrade to increase LY
 - S13660-1325 to -3050 (active area increase x5.3)
- Tag neutrons with coincidence measurement of γ and n

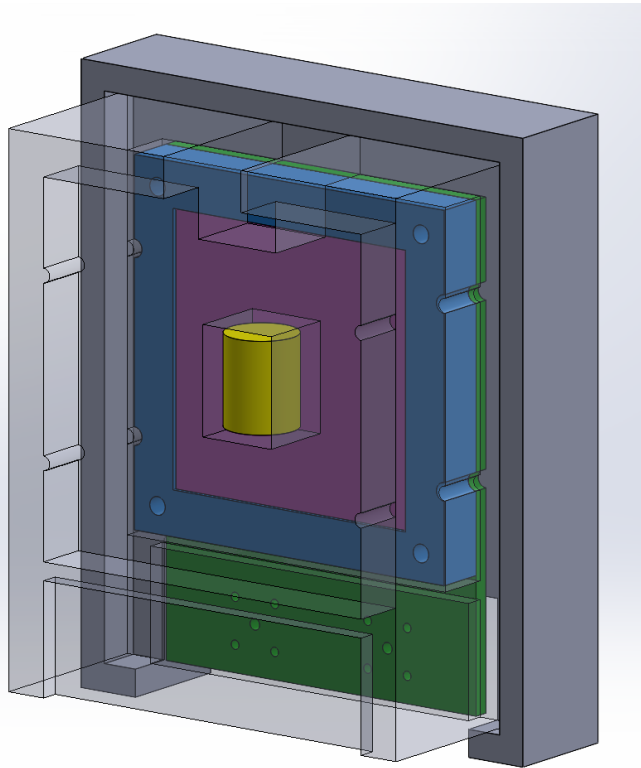


NEW SETUP

- Two PID scintillator tiles on both sides of source to cover large solid angle to maximize coincidences
- Tile thickness increased to 5 mm to boost interaction rate
- Large muon veto on top and bottom
- Additional benefits:
 - Increase proton interaction rate
 - Tag muons with veto



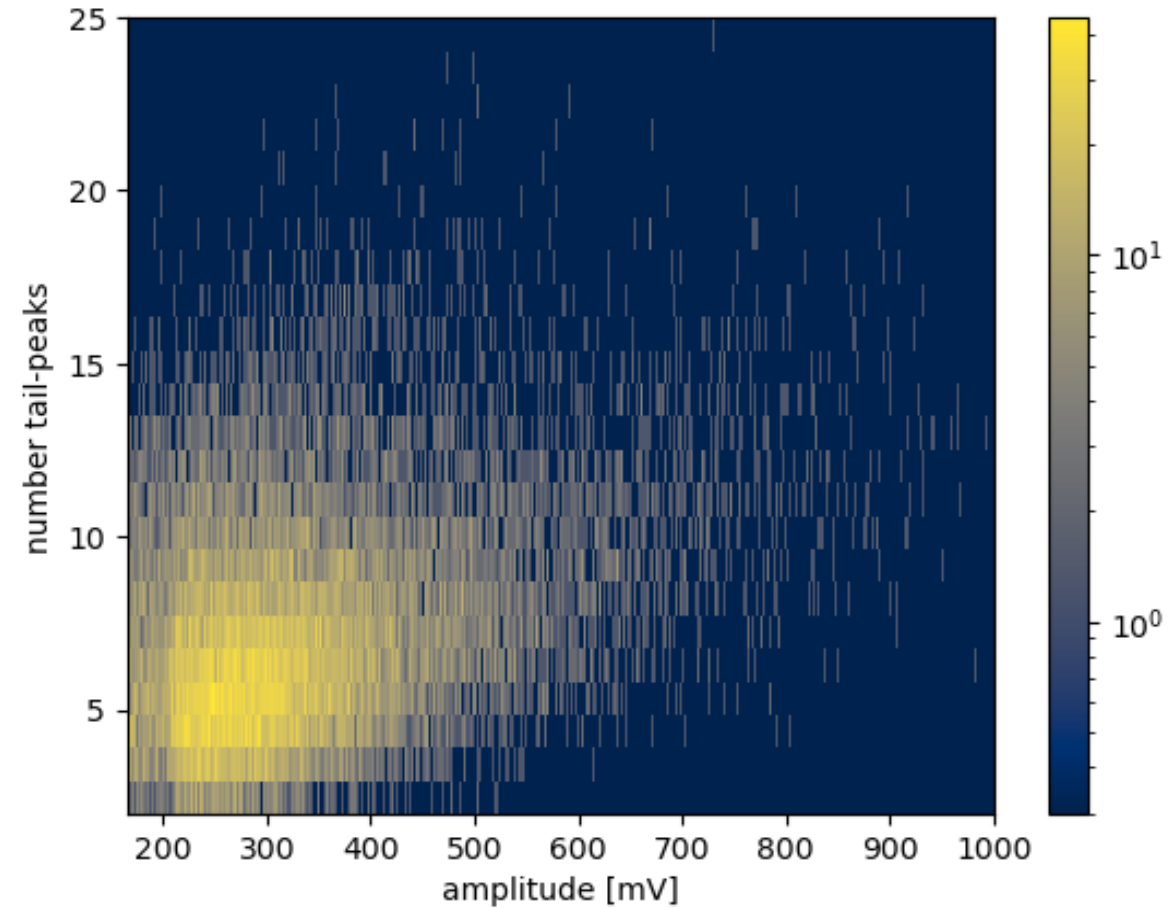
NEW SETUP



FIRST DATA

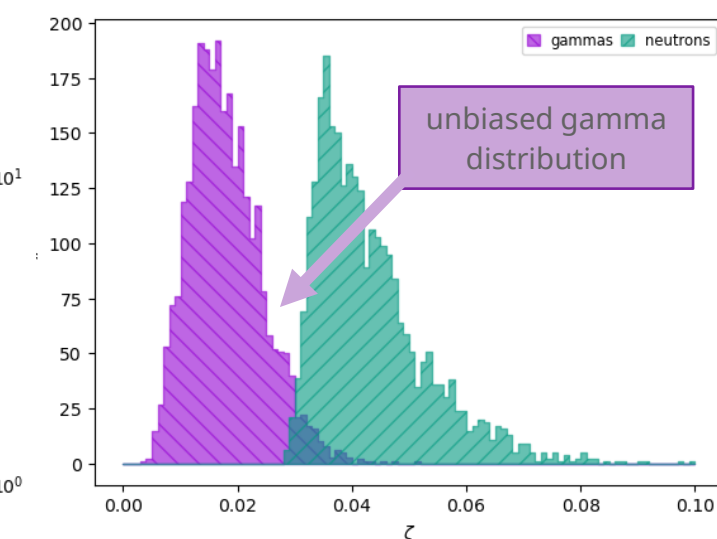
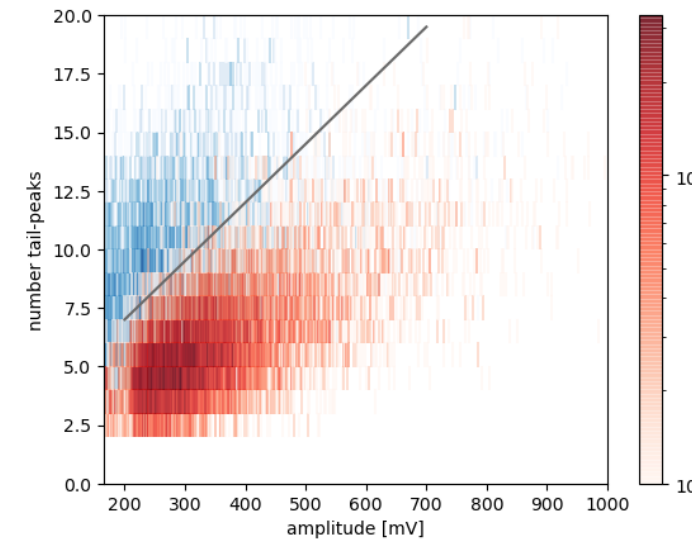
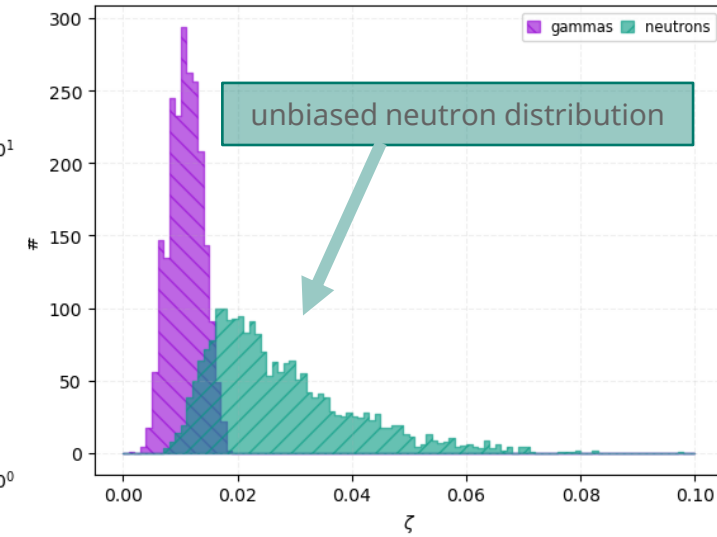
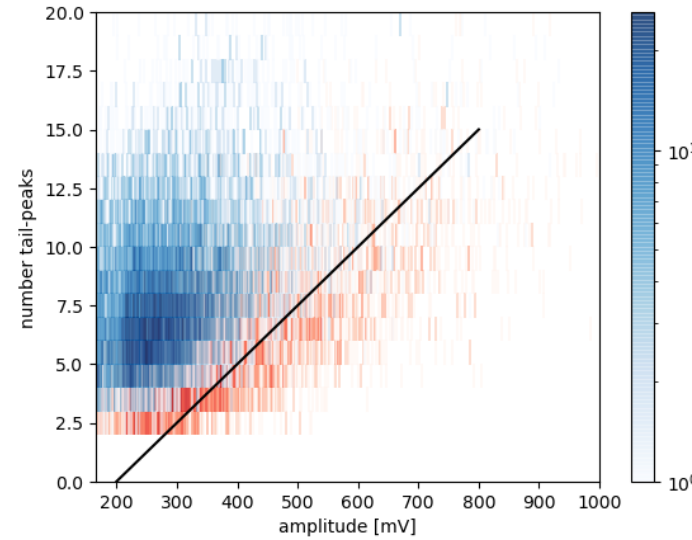
- Data taken for 20h with new setup
- About 2% coincidence events of γ and n

topPMT	x	x	x	-	x	-	-	-
topSiPM	x	-	x	x	-	-	x	-
bottomSiPM	x	x	x	x	x	x	x	x
bottomPMT	x	x	-	x	-	x	-	-
#hits	21	17	108	82	347	285	9424	591240



NEUTRON TAGGING IN ACTION

- Coincidence data used
- Top plots:
 - low cut
 - unbiased neutron sample
- Bottom plots:
 - high cut
 - unbiased gamma sample
- Tagging works
 - clean determination of $n + \gamma$ shapes



SUMMARY

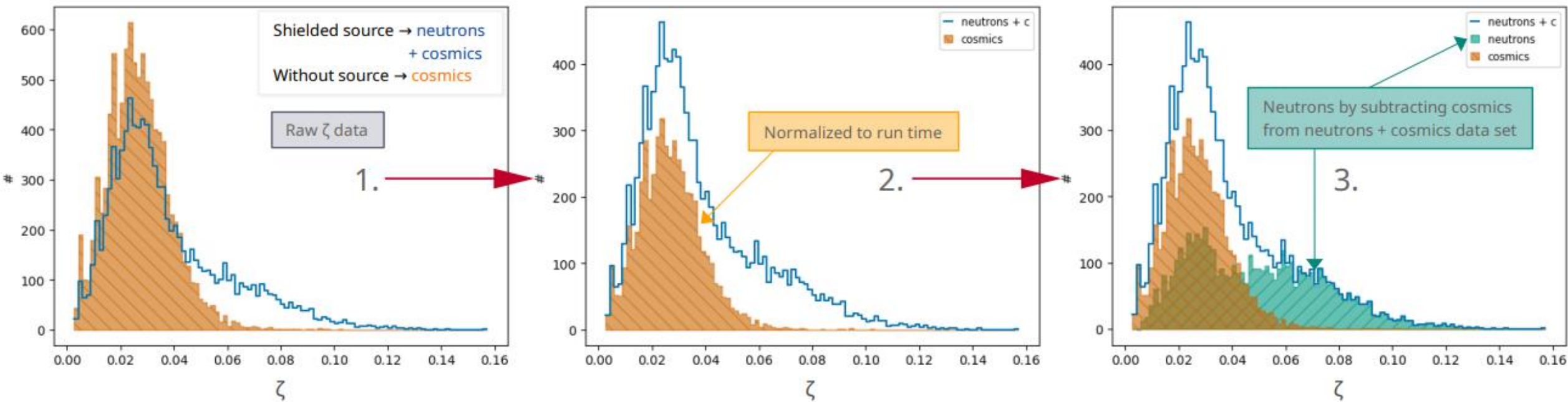
- Neutron ID with EJ-276G scintillator is a success
- Coincident photon + neutron emission from AmBe source can be used to determine neutron scintillation signal
- Next step: Increase LY with larger SiPM to improve separation
- To be discussed: optimal realization in readout electronics (KLauS)



BACKUP

Defining a discriminant

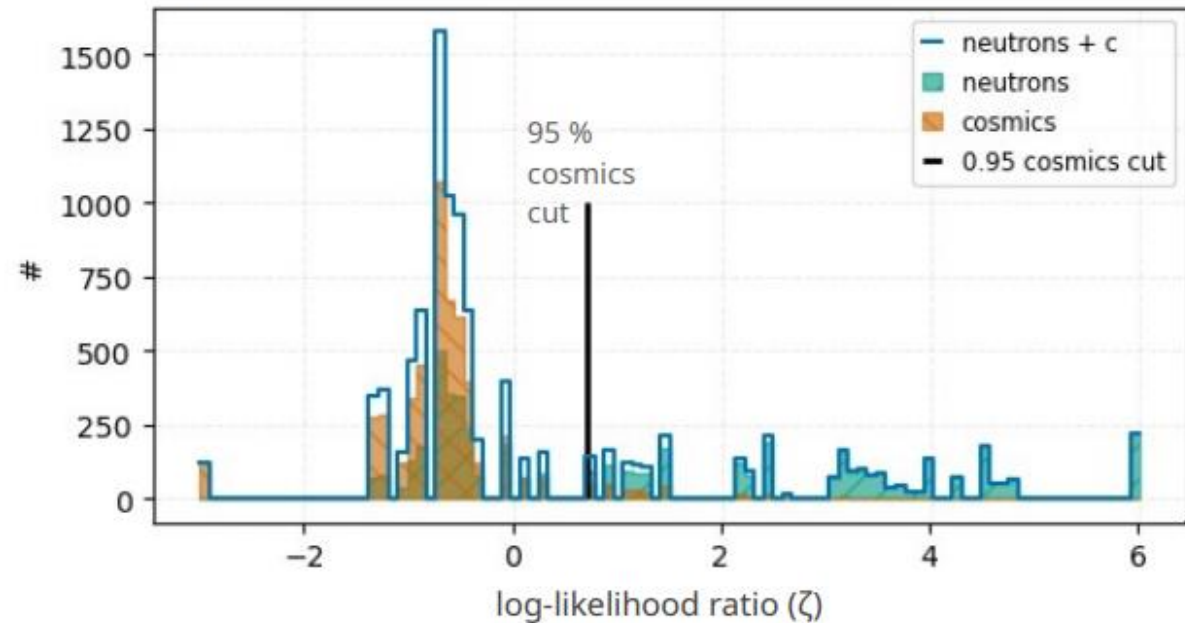
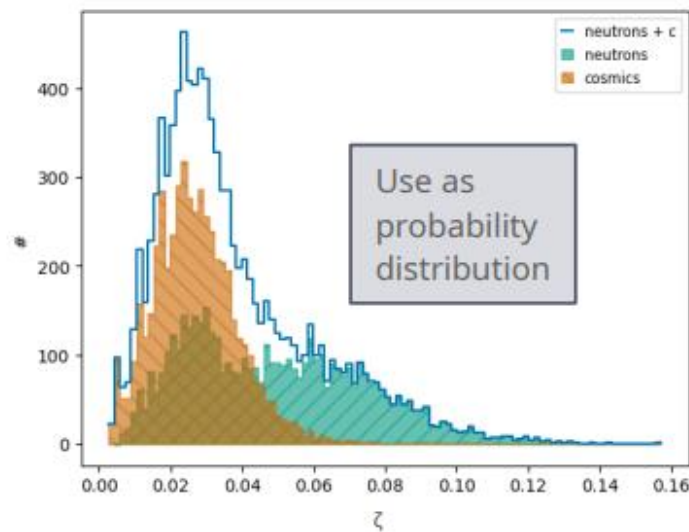
- Separate neutrons and cosmic muons first
 - Cosmic muons and gammas should have similar signal
- Information about amplitude of first peak and number of tail-peaks per event



Distinguish with log-likelihood ratios

Histogram bin height method

- Take for each ζ point the respective bin
- **Bin heights of neutrons and cosmics distribution** as likelihoods to be either one
 - $\log(\text{height neutron}) - \log(\text{height cosmic})$

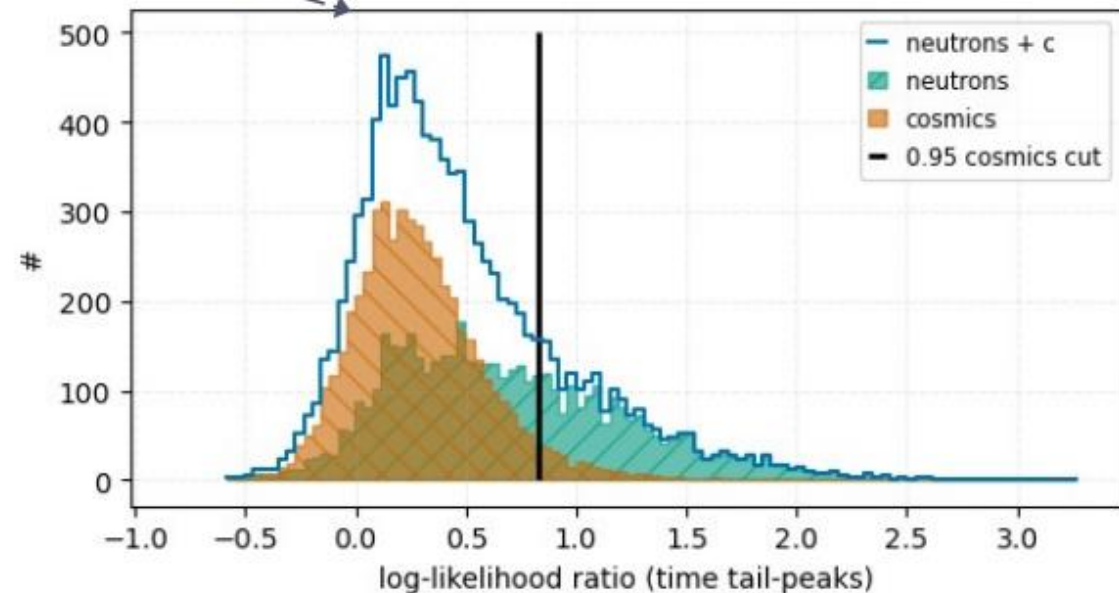
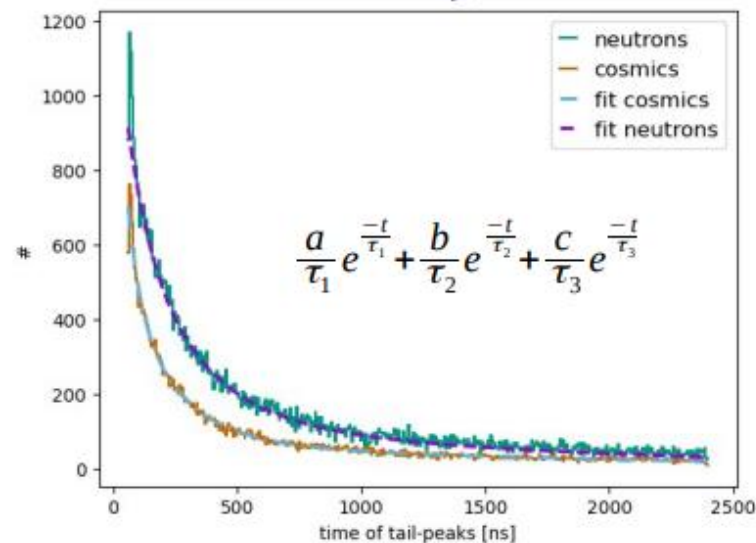


Distinguish with log-likelihood ratios

Time distribution method

- Fit to the time distribution of the tail-peaks of the events
- Calculate log-likelihood ratio of event being **neutron** or **cosmic** by **comparing to the time distribution fits**

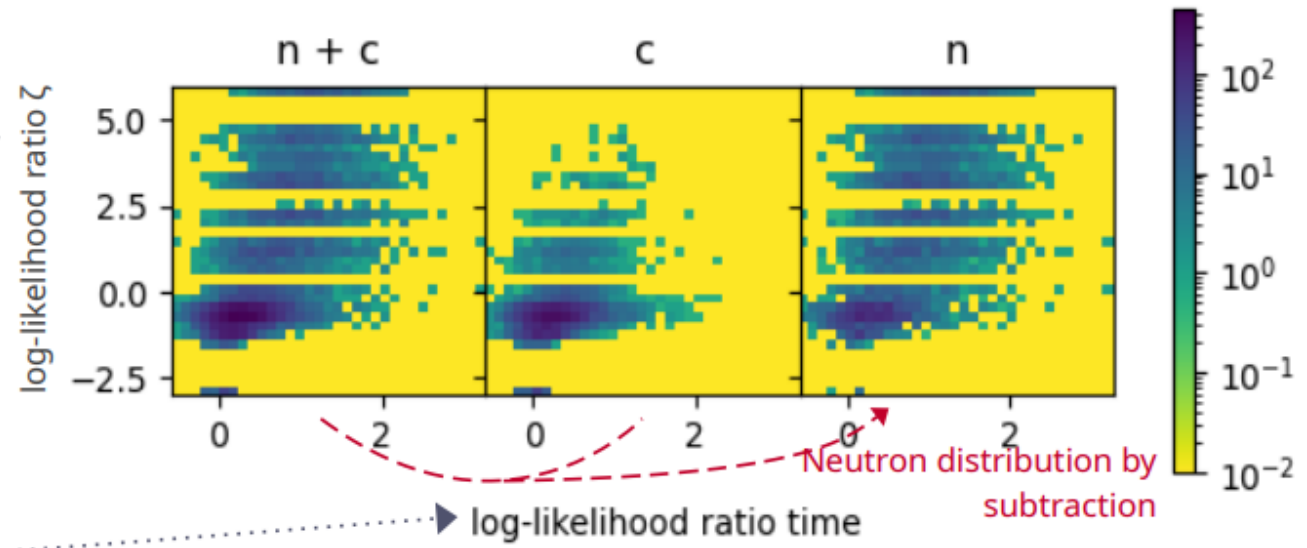
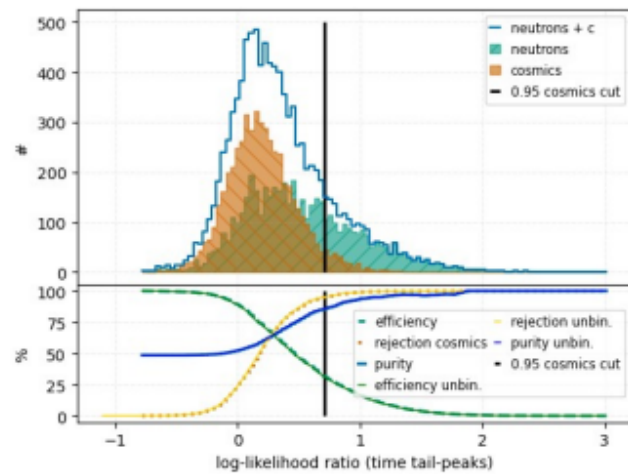
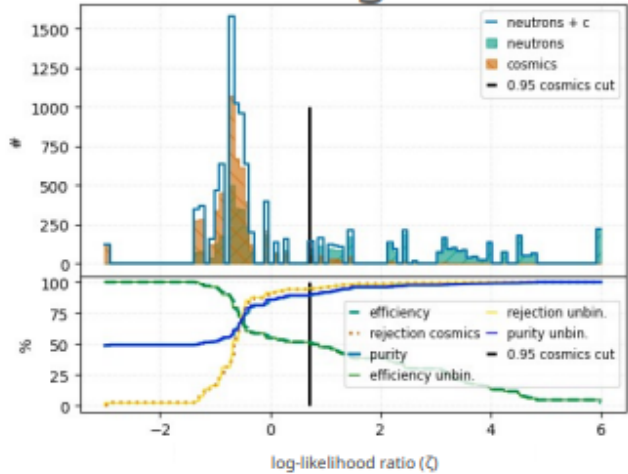
$$\ln \mathcal{L} = \ln \left(\sum \frac{a_n}{\tau_{1,n}} e^{-\frac{t_i}{\tau_{1,n}}} + \frac{b_n}{\tau_{2,n}} e^{-\frac{t_i}{\tau_{2,n}}} + \frac{c_n}{\tau_{3,n}} e^{-\frac{t_i}{\tau_{3,n}}} \right) - \ln \left(\sum \frac{a_\mu}{\tau_{1,\mu}} e^{-\frac{t_i}{\tau_{1,\mu}}} + \frac{b_\mu}{\tau_{2,\mu}} e^{-\frac{t_i}{\tau_{2,\mu}}} + \frac{c_\mu}{\tau_{3,\mu}} e^{-\frac{t_i}{\tau_{3,\mu}}} \right)$$



Combine log-likelihood ratio methods

2d combination

- 2d histograms with time information on x- and bin height method on y-axis



Combine log-likelihood ratio methods

2d combination

- 2d histograms with time information on x- and bin height method on y-axis
- Use 2d combination as new distributions for new log-likelihood ratio

method	Efficiency [%]
LLR ζ	49.60
LLR time tail-peaks	36.19
Combination LLR - LLR	54.03

